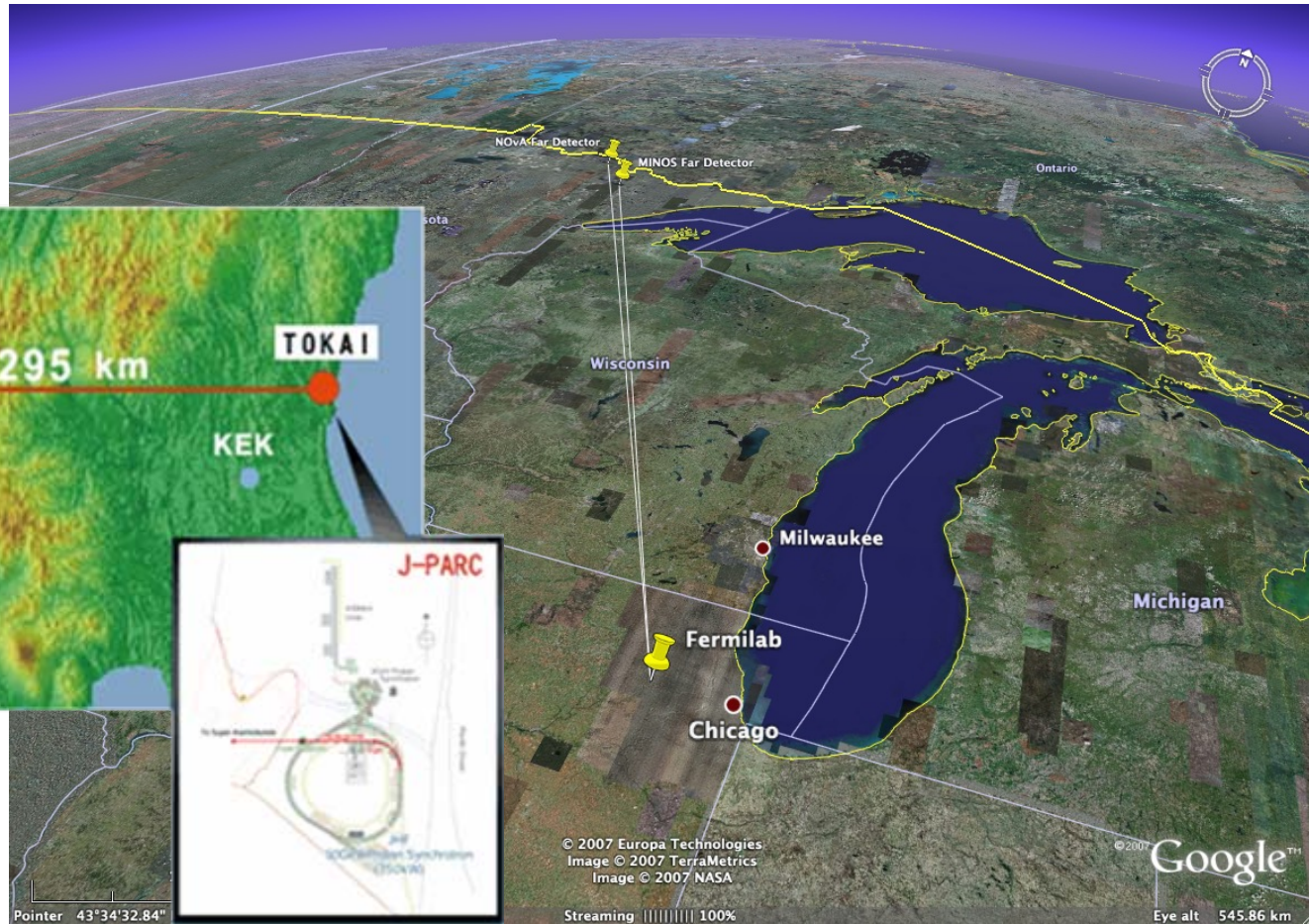
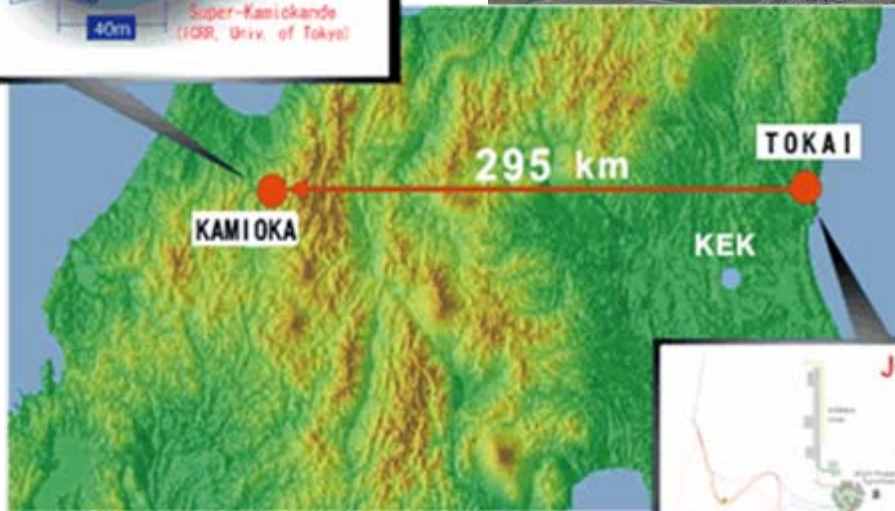
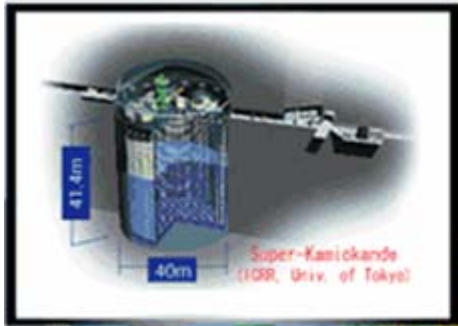


# Accelerator Neutrino Experiments



And DEEP CORE

# Accelerator Long-Baseline

## Neutrino Physics Goals

$\nu_e$  appearance

- Test the  $\nu_\mu$  to  $\nu_x$  oscillation hypothesis

- Measure precisely  $|\Delta m_{32}^2|$  and  $\sin^2(2\theta_{23})$

- Search for  $\nu_\mu$  to  $\nu_e$  oscillations

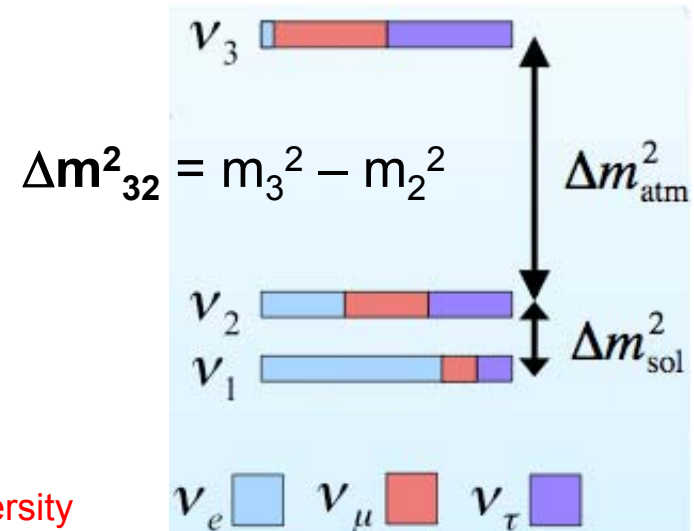
- sensitive to  $\theta_{13}$

- Determine the mass hierarchy

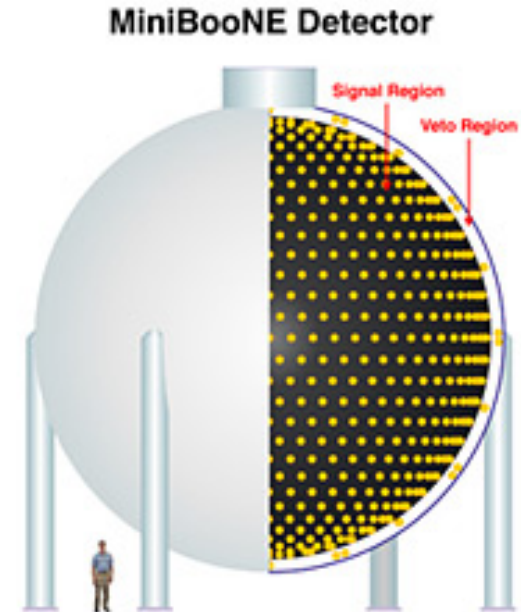
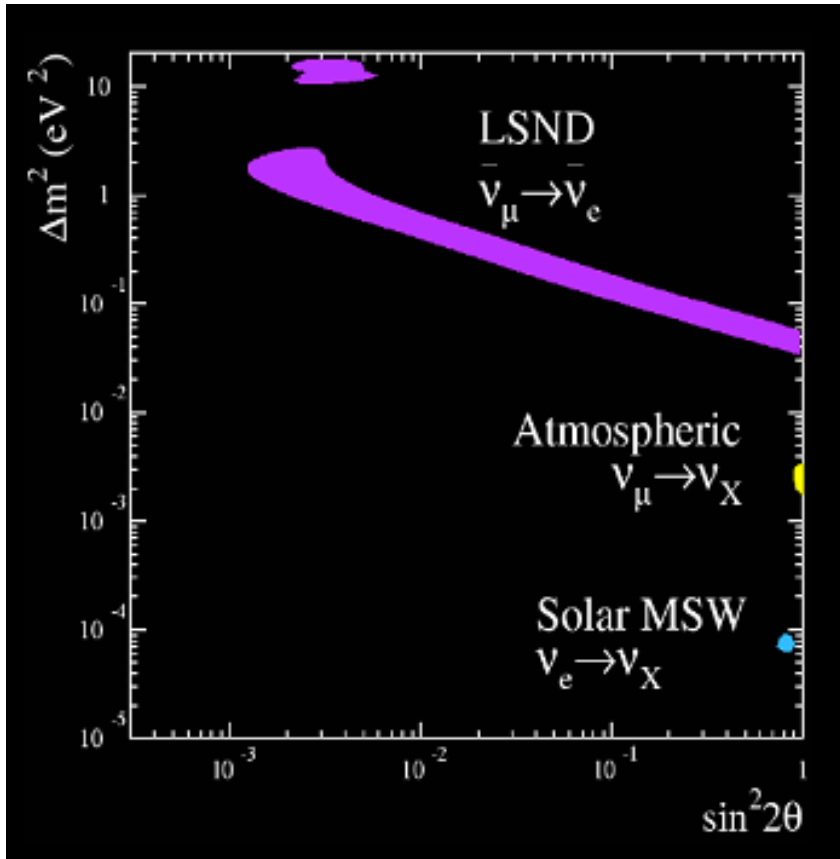
- Measure  $\delta_{CP}$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$\nu_\mu$  disappearance



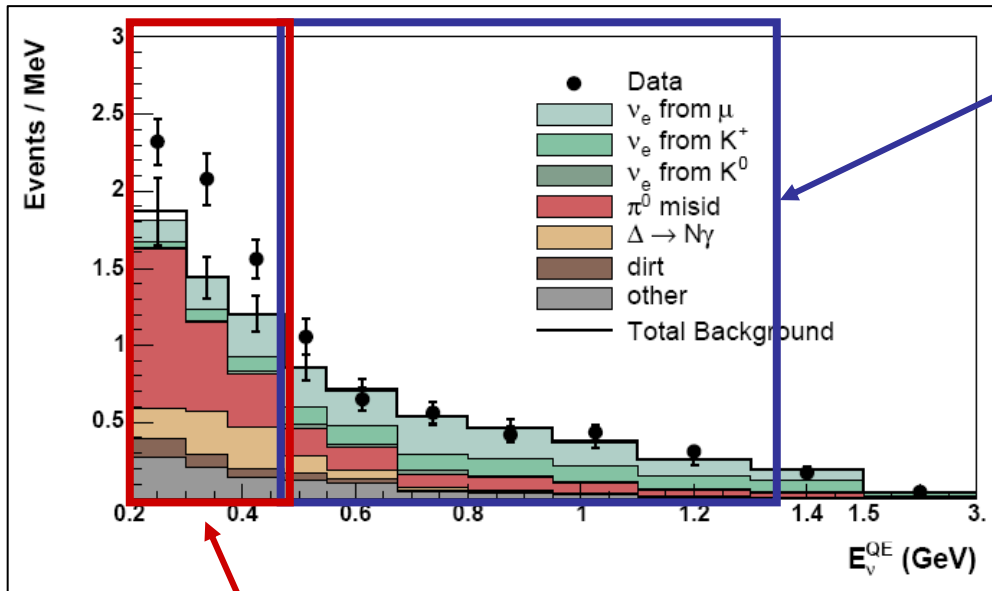
# MiniBooNE - A search for $\nu_e$ appearance at $\Delta m^2 \sim 1 \text{ eV}^2$



LSND Best fit:  $\sin^2(2\theta) = 0.003$ ,  $\Delta m^2 = 1.2 \text{ eV}^2$

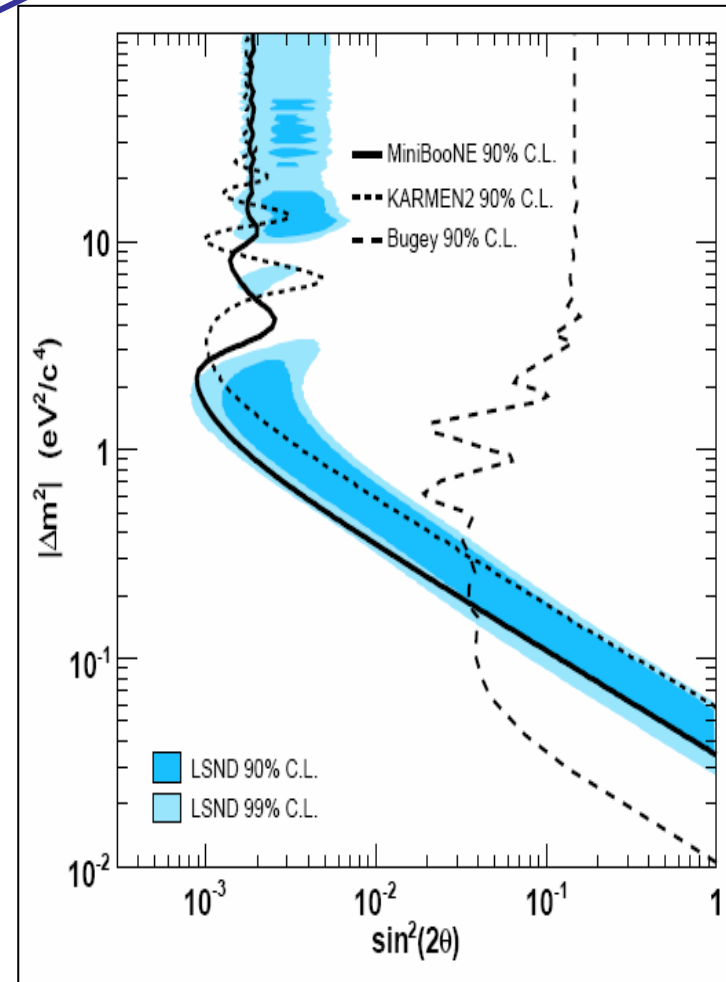
**Observed no excess consistent with the LSND two-neutrino oscillation**

$\nu_\mu \rightarrow \nu_e$  signal region

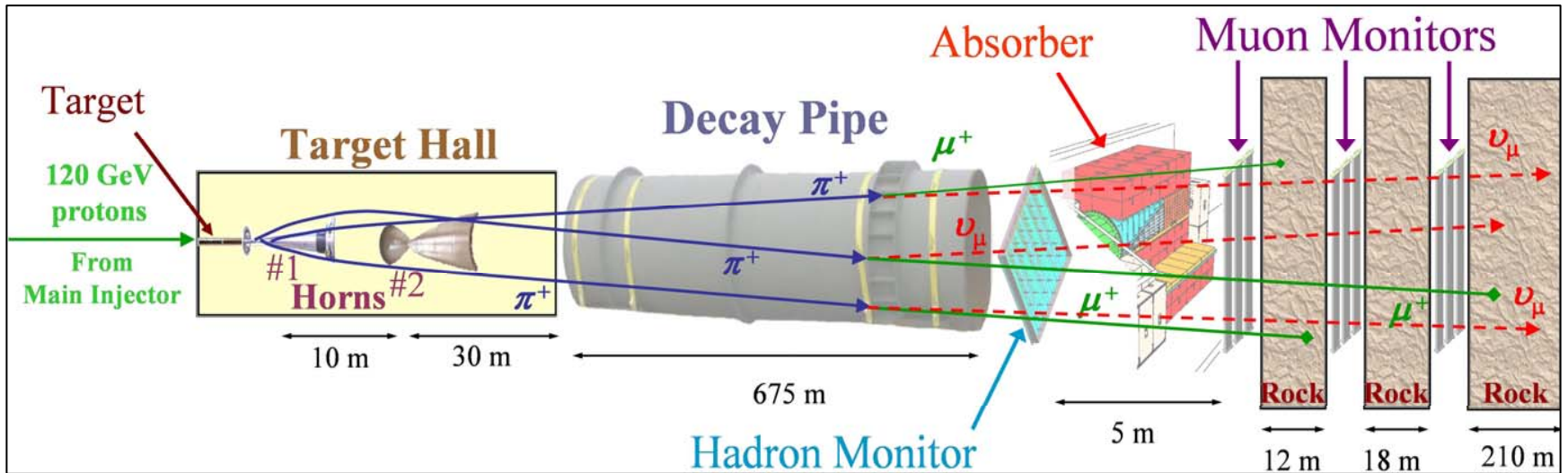


**low energy excess region**  
 **$128.8 \pm 43.4$  excess events ( $3.0\sigma$ )**

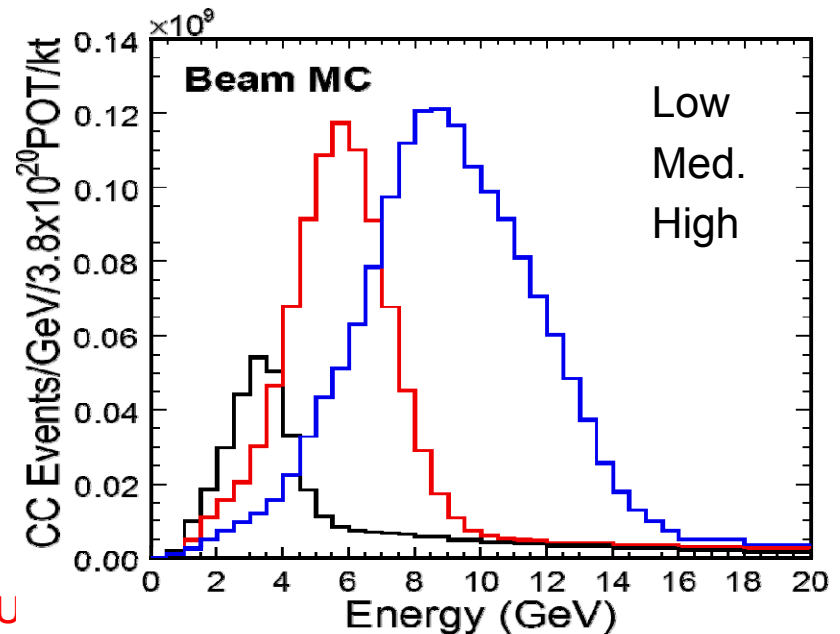
**No sign of excess in recent  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  data**



# Neutrino Beam (NuMI)



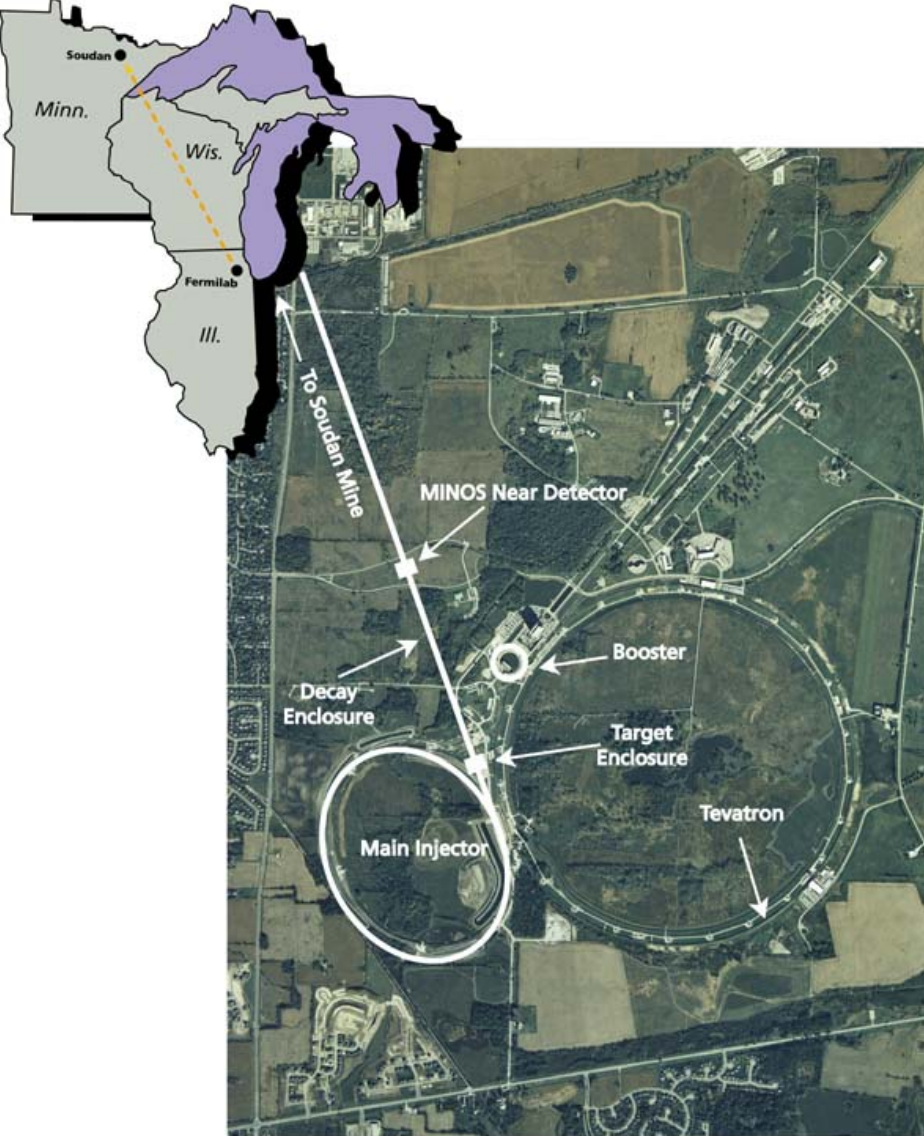
- 120 GeV protons strike target
- 10  $\mu$ s long pulse of  $3 \times 10^{13}$  protons every 2.2 seconds (275 kW)
- Two magnetic horns focus secondary  $\pi/K$ 
  - decay of  $\pi/K$  produce neutrinos
- Variable neutrino beam energy



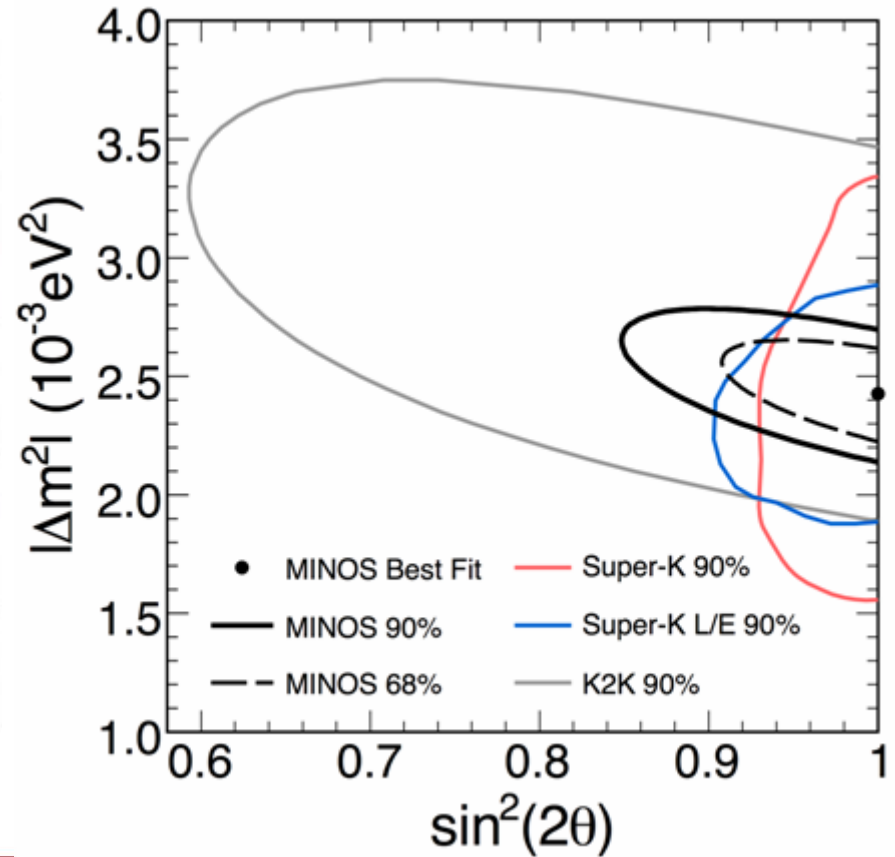
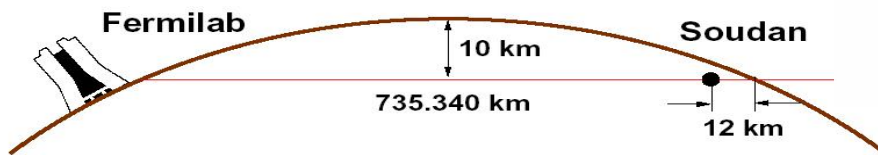
# MINOS

$$|\Delta m^2| = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2$$

(68% C.L.)



FERMILAB #98-1321D



rsity

# Far Detector Energy Spectrum

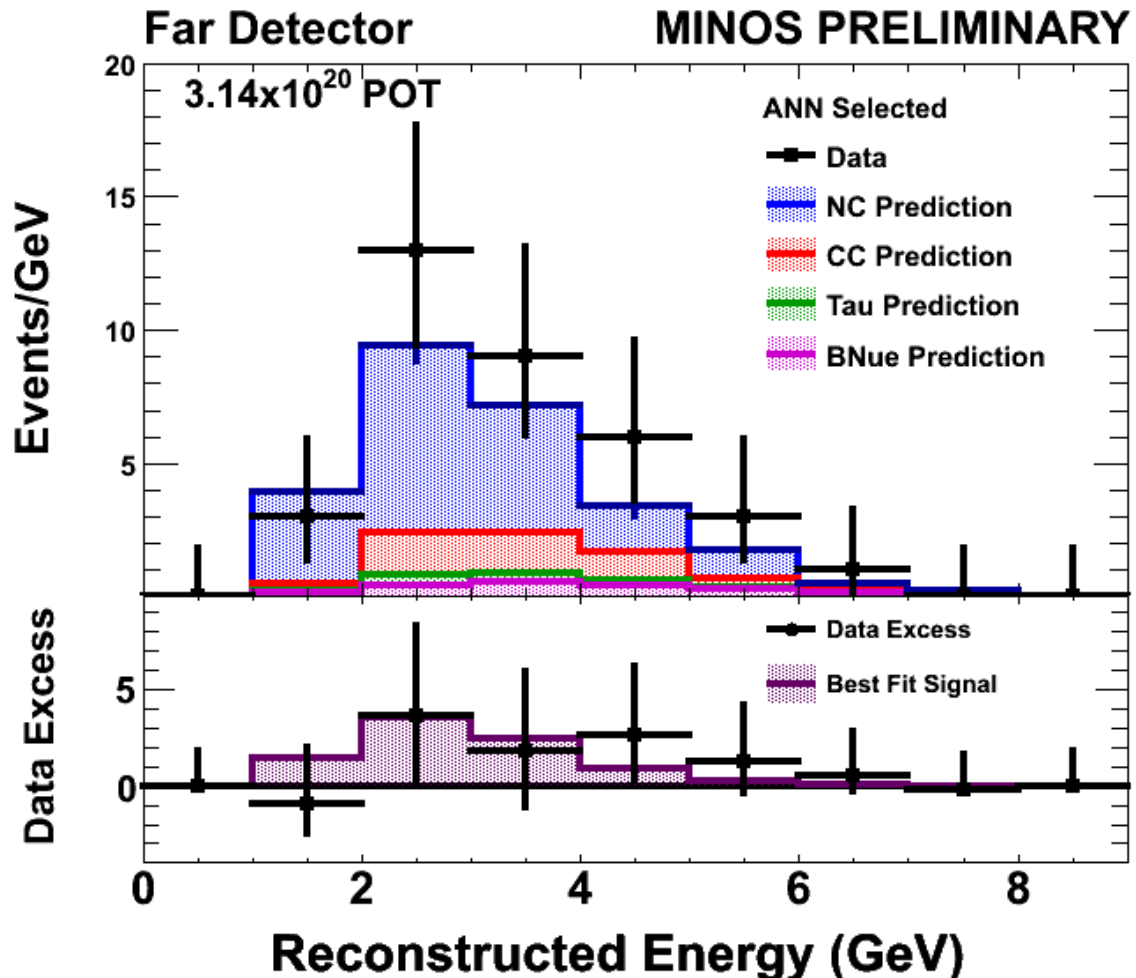
- A blind analysis was performed:
  - all procedures for calculating background and signal were finalised before the Far detector data were looked at
- Expected background:
 

**$27 \pm 5(\text{stat}) \pm 2(\text{sys})$**
- Observed events:
 

**35**
- A  $1.5\sigma$  excess over background prediction

Workshop on "Low Energy"  
Neutrino Physics and Astrophysics

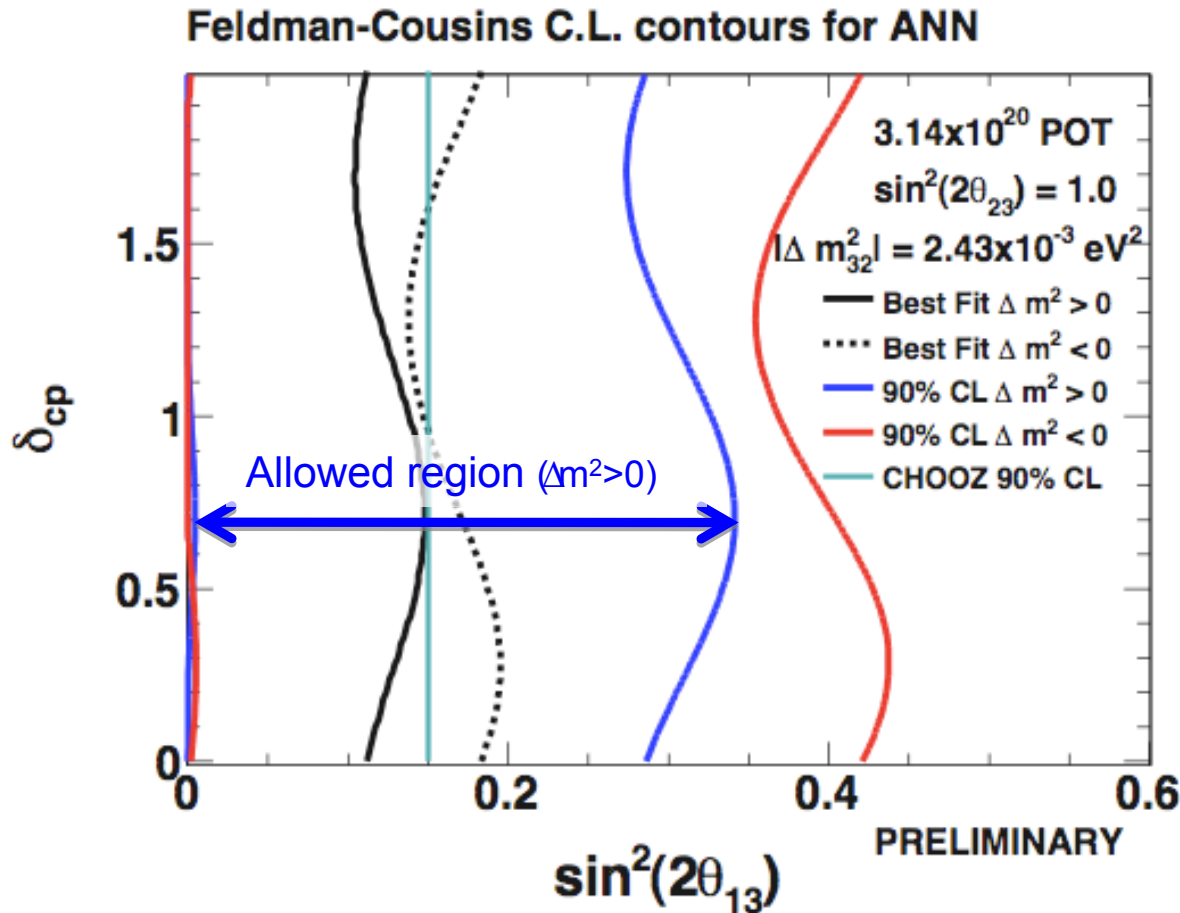
Scott Menary – York University



Fit the data to the oscillation hypothesis, obtain the signal prediction for the best fit point

# Allowed Region

- A Feldman-Cousins method was used
- Fit simply to the number of events from 1-8 GeV
- Best fit and 90% C.L. limits are shown:
  - for both mass hierarchies
  - at MINOS best fit value for  $\Delta m^2_{32}$  &  $\sin^2(2\theta_{23})$



Normal hierarchy ( $\delta_{CP}=0$ ):

$$\sin^2(2\theta_{13}) < 0.29 \text{ (90\% C.L.)}$$

Inverted hierarchy ( $\delta_{CP}=0$ ):

$$\sin^2(2\theta_{13}) < 0.42 \text{ (90\% C.L.)}$$



# CERN Neutrinos to Gran Sasso

long base-line appearance experiment:

- Produce muon neutrino beam at CERN
- Measure tau neutrinos in Gran Sasso

Experiments:

OPERA (1200 ton), ICARUS (600 ton)

$4.5 \cdot 10^{19}$  pot/year (200 days, nominal intensity)

→  $2.2 \cdot 10^{17}$  pot/day  $\sim 10^{17} \nu_{\mu}$  /day  $\sim 10^{11} \nu_{\mu}$  /day in detector

→ 3600  $\nu_{\mu}$  interactions/year in OPERA (charged current interactions)

→ 2-3  $\nu_{\tau}$  interactions detected/year in OPERA



$\sim 1 \nu_{\tau}$  observed interaction with  $2 \cdot 10^{19}$  pot

CNGS Run 2008:  $1.78 \cdot 10^{19}$  pot

# The Known and Unknown in Neutrino Physics

## The CP Violation Parameter

Three Neutrino PMNS Mixing Matrix:

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

$$= \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

From Atmospheric and Long Baseline Disappearance Measurements

From Reactor Disappearance Measurements

From Long Baseline Appearance Measurements

From Solar Neutrino Measurements

## In Vacuum the Oscillation Probability is:

- $P(\nu_\mu \rightarrow \nu_e) = P_1 + P_2 + P_3 + P_4$

- $P_1 = \sin^2(\theta_{23}) \sin^2(2\theta_{13}) \sin^2(1.27 \Delta m_{13}^2 L/E)$

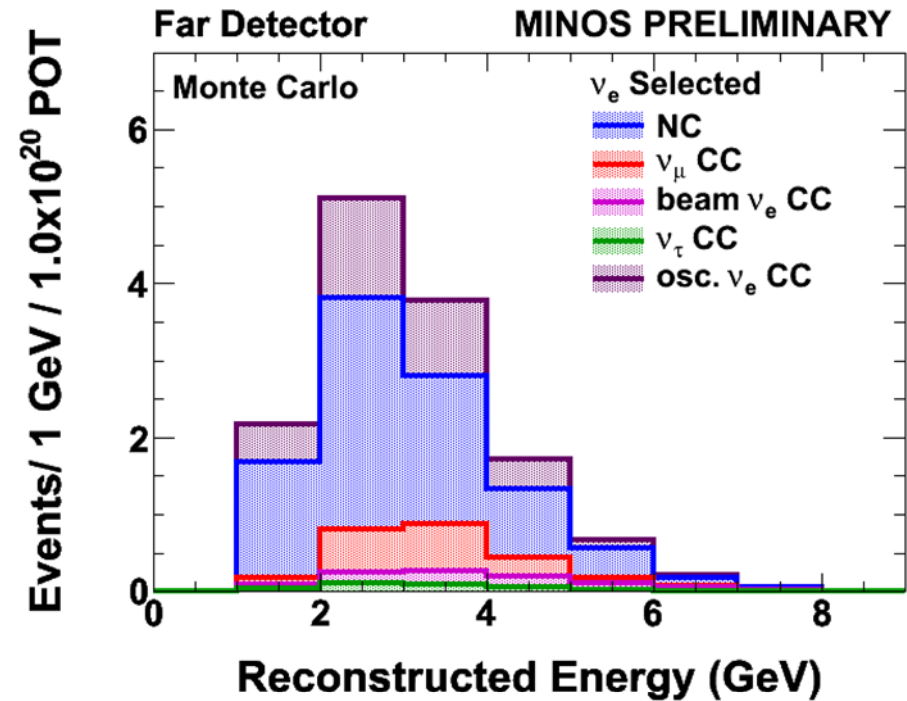
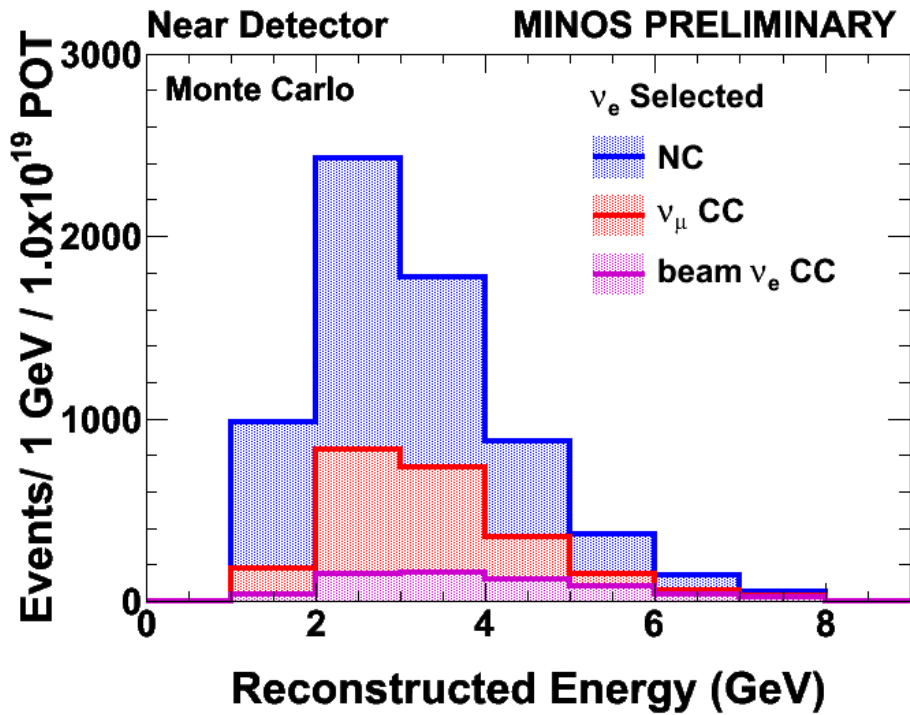
- $P_2 = \cos^2(\theta_{23}) \sin^2(2\theta_{12}) \sin^2(1.27 \Delta m_{12}^2 L/E)$

- $P_3 = J \sin(\delta) \sin(1.27 \Delta m_{13}^2 L/E)$

- $P_4 = J \cos(\delta) \cos(1.27 \Delta m_{13}^2 L/E)$

where  $J = \cos(\theta_{13}) \sin(2\theta_{12}) \sin(2\theta_{13}) \sin(2\theta_{23}) \times$   
 $\sin(1.27 \Delta m_{13}^2 L/E) \sin(1.27 \Delta m_{12}^2 L/E)$

# MINOS MC Event Composition in 2 Detectors



- Primary background from NC events, also
  - high-y  $\nu_\mu$  CC, beam  $\nu_e$ , oscillated  $\nu_\tau$  at Far detector
- Right plot: purple shows an appearance signal at the Chooz limit ( $\sin^2 2\theta_{13} = 0.15$ )

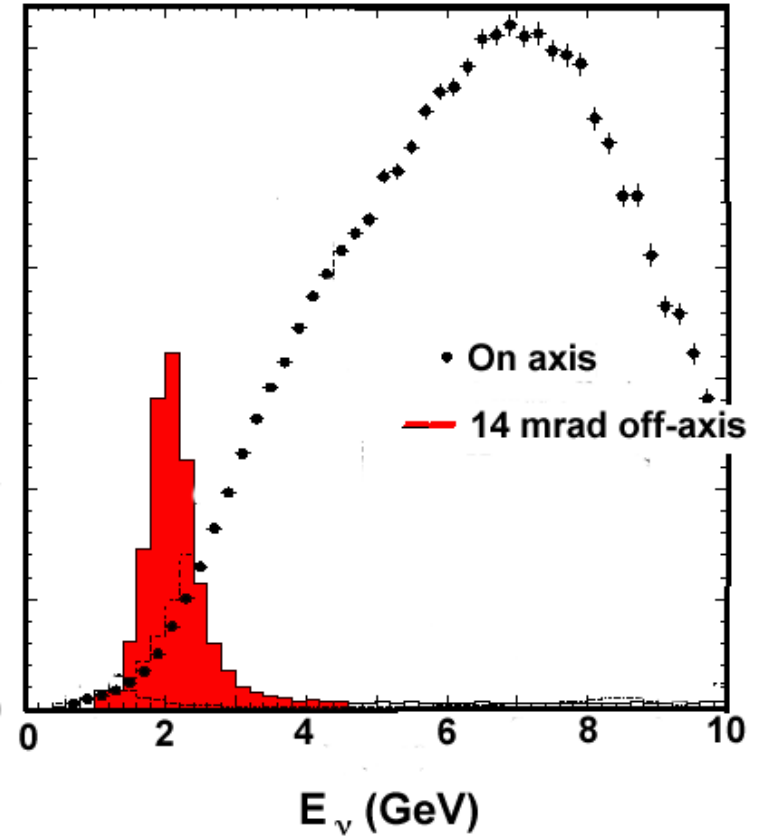
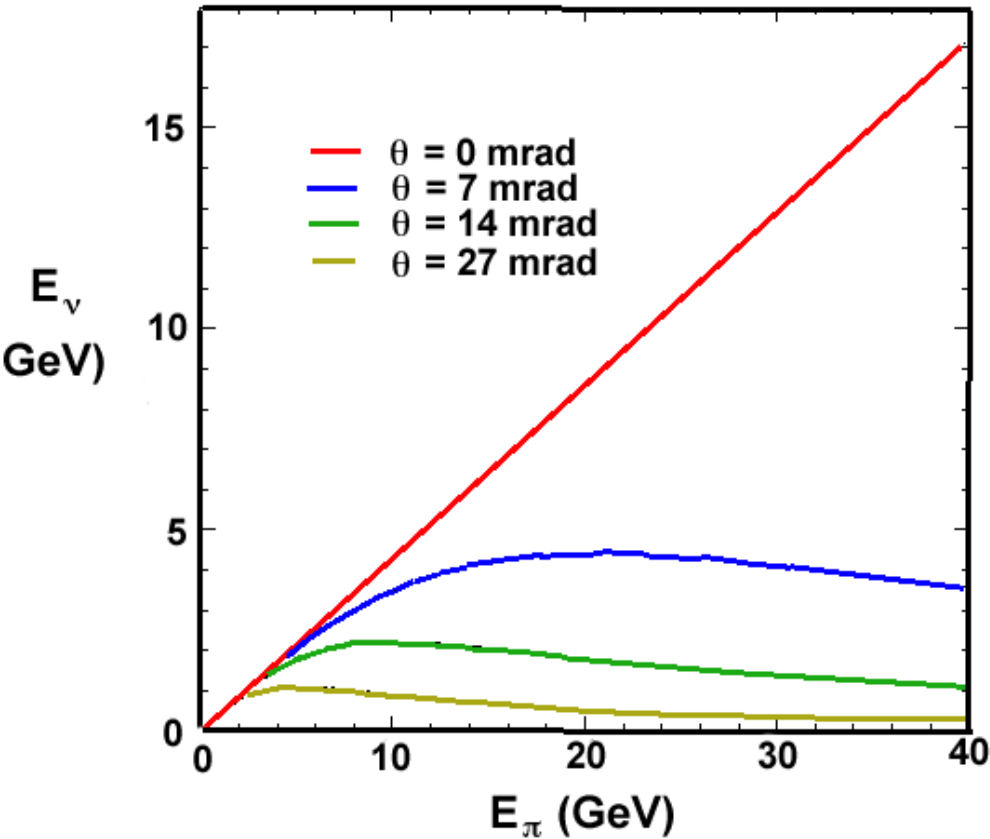
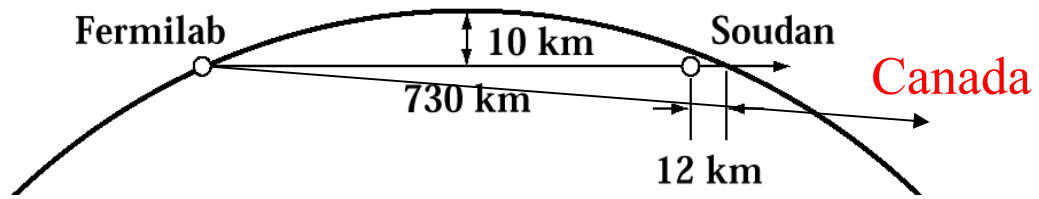
- In matter,  $P_1$  will be approximately multiplied by  $(1 \pm 2E/E_R)$  and  $P_3$  and  $P_4$  will be approximately multiplied by  $(1 \pm E/E_R)$ , where the top sign is for neutrinos with normal mass hierarchy and antineutrinos with inverted mass hierarchy.
- Different baselines “pick out” different terms which helps to break some of the degeneracies.

$$P = f(\sin^2 2\theta_{13}, \delta, \text{sgn}(\Delta m_{13}^2), \Delta m_{12}^2, \Delta m_{13}^2, \sin^2 2\theta_{12}, \sin^2 2\theta_{23}, L, E)$$

**3 unknowns**, 3 parameters under control  $L, E + \nu_\mu / \text{anti } \nu_\mu$

To fight neutral-current background could use a narrow-band beam and a detector technology which does a good job of  $e$  VS  $\pi^0$  identification

# The "Off-axis" Beam



$P_{\max}$  for  $L/E \sim 500$  so for  $E \sim 2$  GeV you want detector at  $L \sim 1000$  km - Canada

# Where to Put NO<sub>ν</sub>A?

“The Ash River site is the furthest available site from Fermilab along the NuMI beamline. This maximizes NO<sub>ν</sub>A’s sensitivity to the mass ordering.”  
M. Goodman, NOW 2008 (Sept)

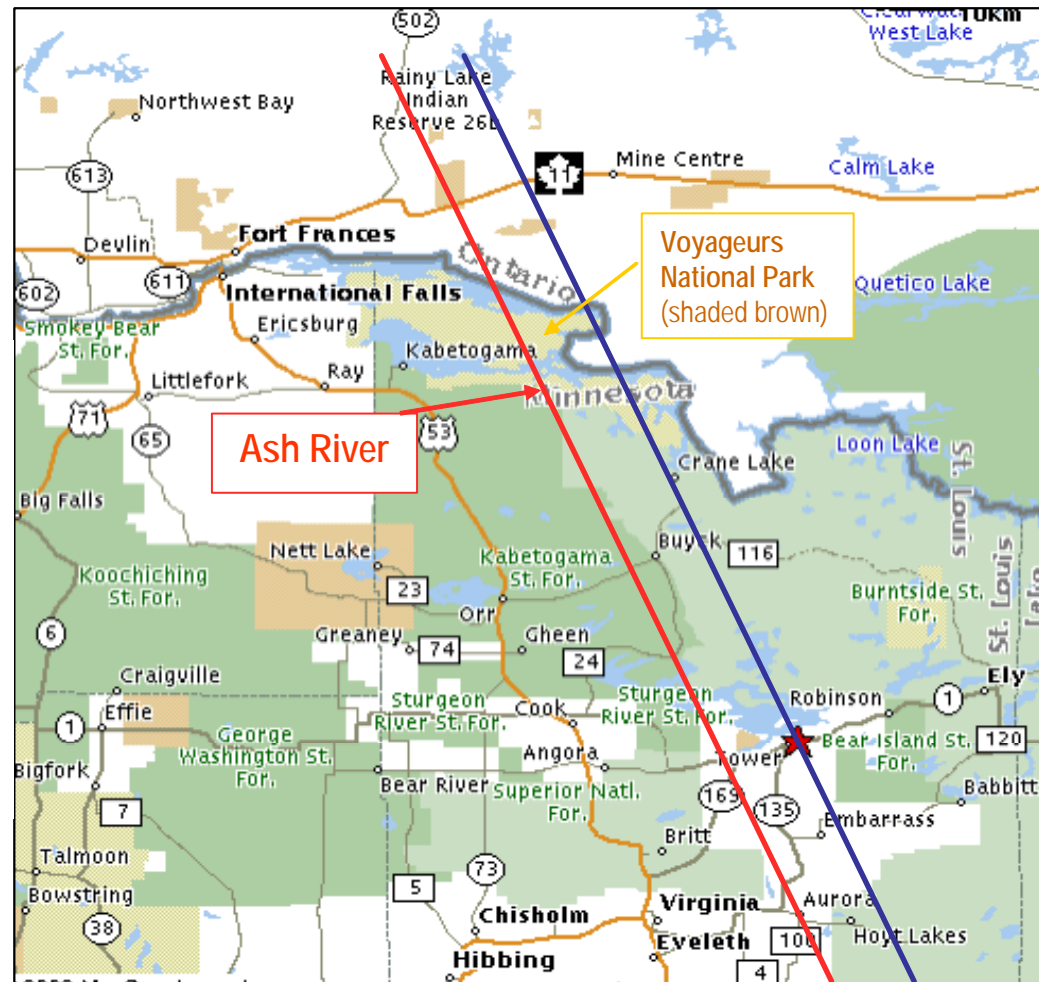
experiment. The main baseline experiment (NO<sub>ν</sub>A) was originally designed with a lower tracking calorimeter to be placed off-axis in the NuMI beam [37]. Since the then preferred baseline of  $L \simeq 712$  km turned out to be too short to be complementary to the T2K experiment in Japan, longer baselines were suggested in Refs. [21, 38–40].<sup>2</sup> As a rule of thumb,

$$\left(\frac{L}{E}\right)_{\text{NuMI}} \gtrsim \left(\frac{L}{E}\right)_{\text{T2K}} \quad (3)$$

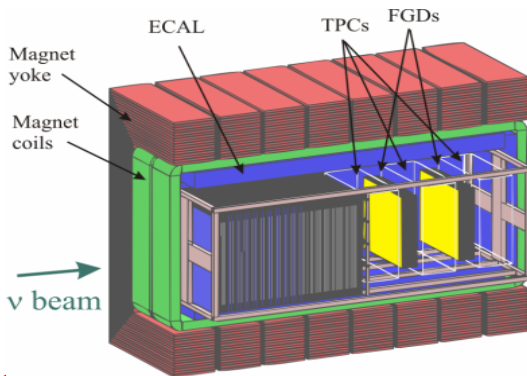
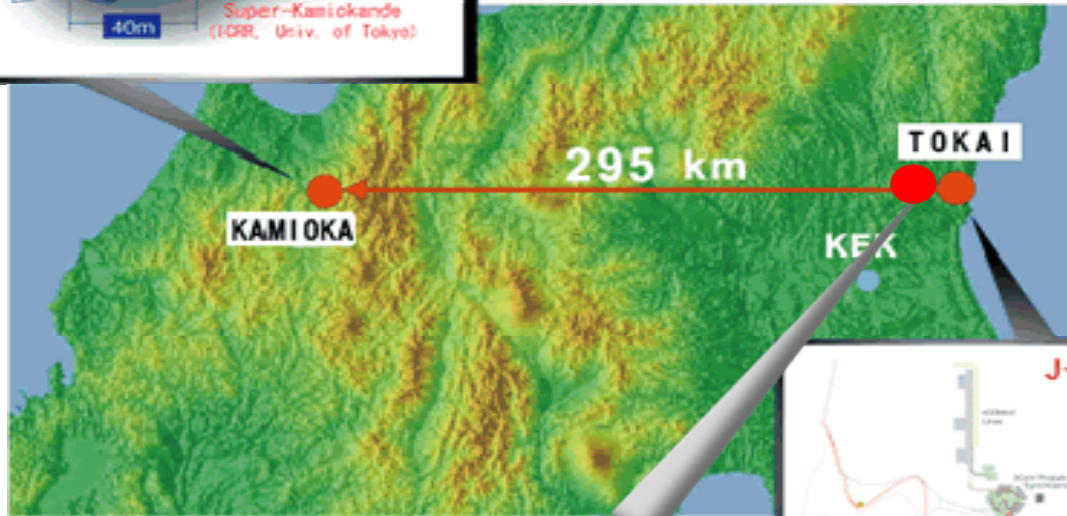
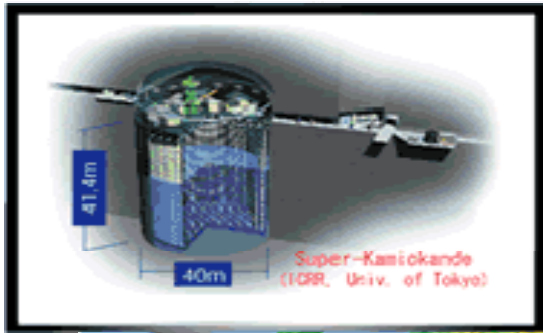
was found in order to yield synergistic physics, which translates to  $L_{\text{NuMI}} \gtrsim 862$  km for  $0.72^\circ$  off-axis angle. Hence, a longer baseline,  $L \simeq 810$  km to the Ash River site in Minnesota, has been proposed for the NO<sub>ν</sub>A experiment [42]. A typical off-axis angle suggested is  $0.85^\circ$ , corresponding to 12 km off-axis at this baseline. In addition, a Totally Active Scintillating Detector (TASD) is the now accepted detector technology, often considered with a mass of 25 kt. Alternative sites with longer baselines in Canada (because of the beam geometry), such as Vermillion Bay, with potentially attractive physics potential are not actively being considered.

[V. Barger](#), [P. Huber](#), [D. Marfatia](#),  
[W. Winter](#)

[arXiv:hep-ph/0703029](https://arxiv.org/abs/hep-ph/0703029)



# T2K – J-PARC to SuperK



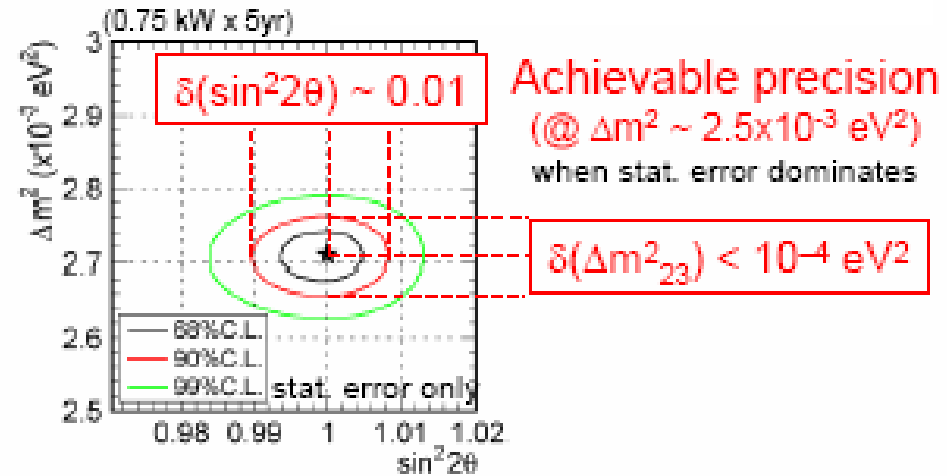
- 30 GeV Protons first transported to **J-PARC Hadron Experimental Hall** in January of this year.
- Were on track for first neutrinos this year.
- Opposite situation to NO<sub>v</sub>A – here they have a detector but not a beam.
- Much smaller matter effect than NO<sub>v</sub>A



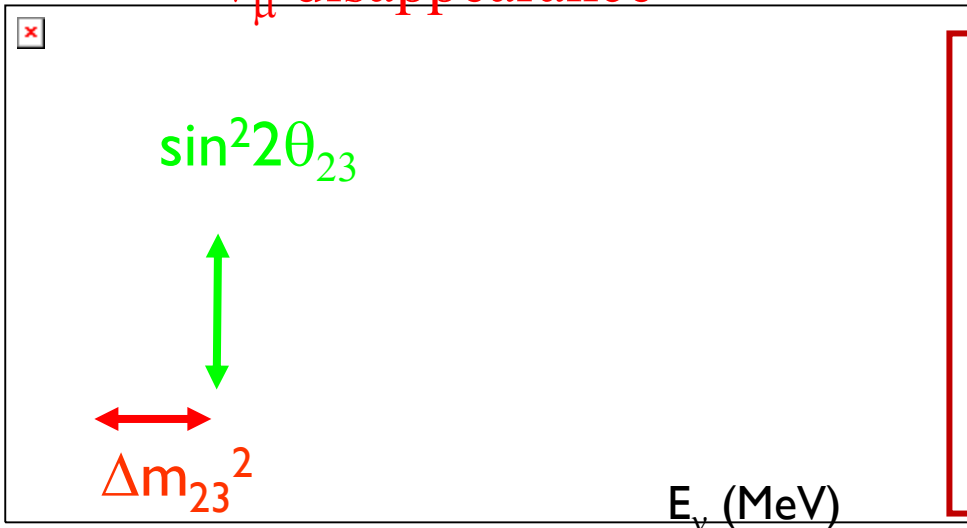
# Main T2K Measurements: $\sin^2 2\theta_{23}$ , $\Delta m^2_{23}$

## Phase I:

- ▶ 5 years X 0.75 MW beam
- ▶  $5 \times 10^{21}$  pot
- ▶ Measurement of mixing angles

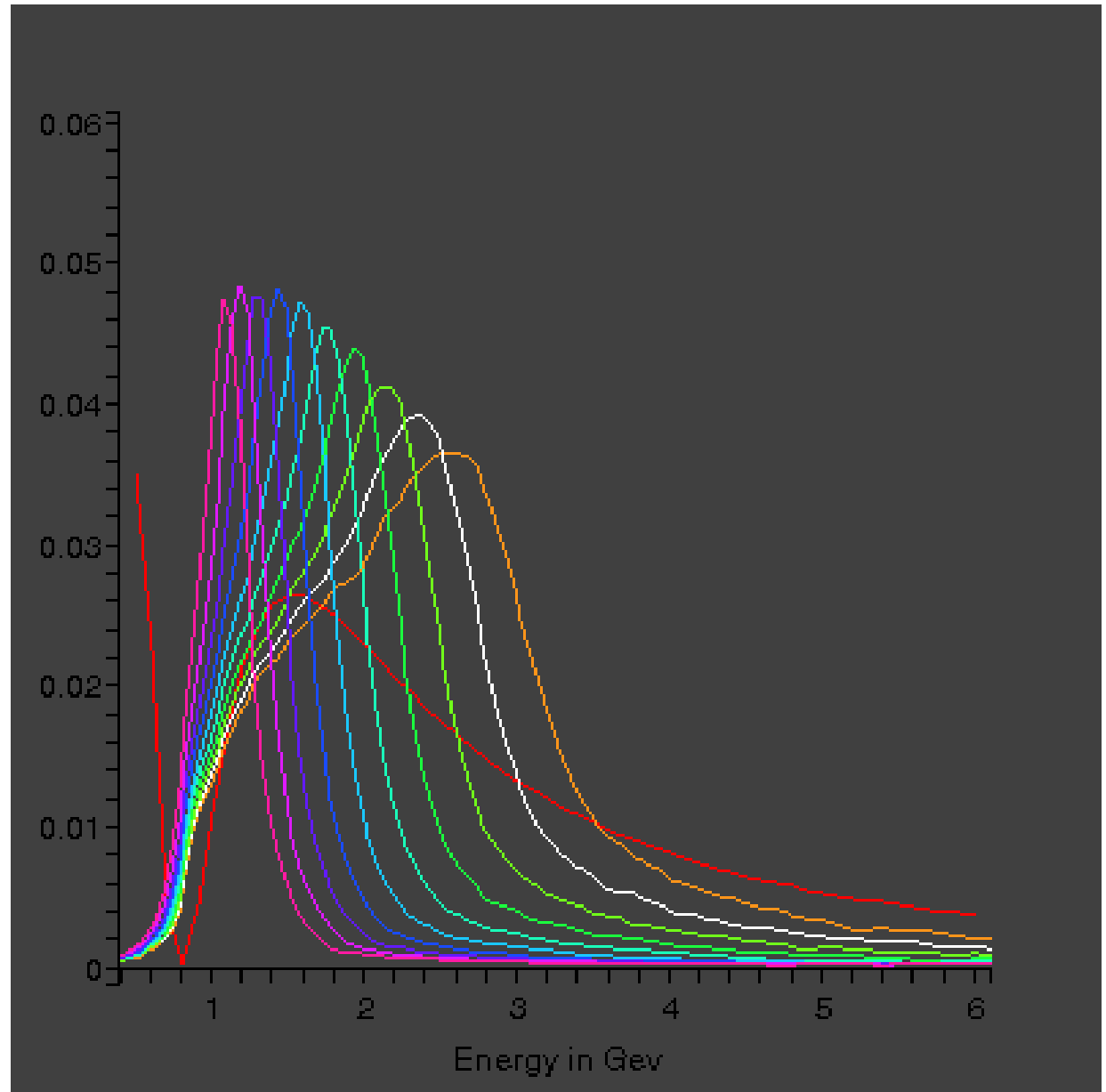


$\nu_{\mu}$  disappearance



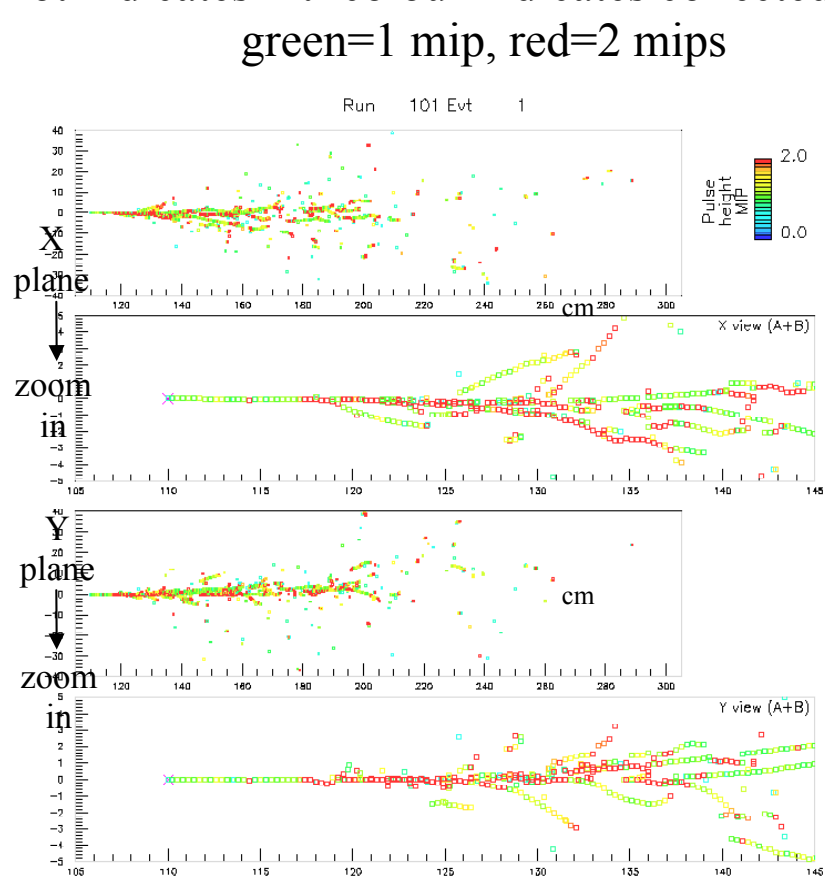
- Use CC Quasi Elastic Events
  - Can reconstruct Neutrino Energy.
  - Background from non-CCQE interactions.

The narrow-band beam at first seemed ideal but the energy at which the oscillation maximum occurs can be quite dependent on the value of  $\delta_{cp}$ . The **Report of the US long baseline neutrino experiment study** [arXiv:0705.4396](https://arxiv.org/abs/0705.4396) recognized that there is a great deal of power in using a wide-band beam (the first two oscillation maxima in the same experiment) but you must have excellent  $e$  versus  $\pi^0$  recognition - you need a Liquid Argon TPC (LArTPC)

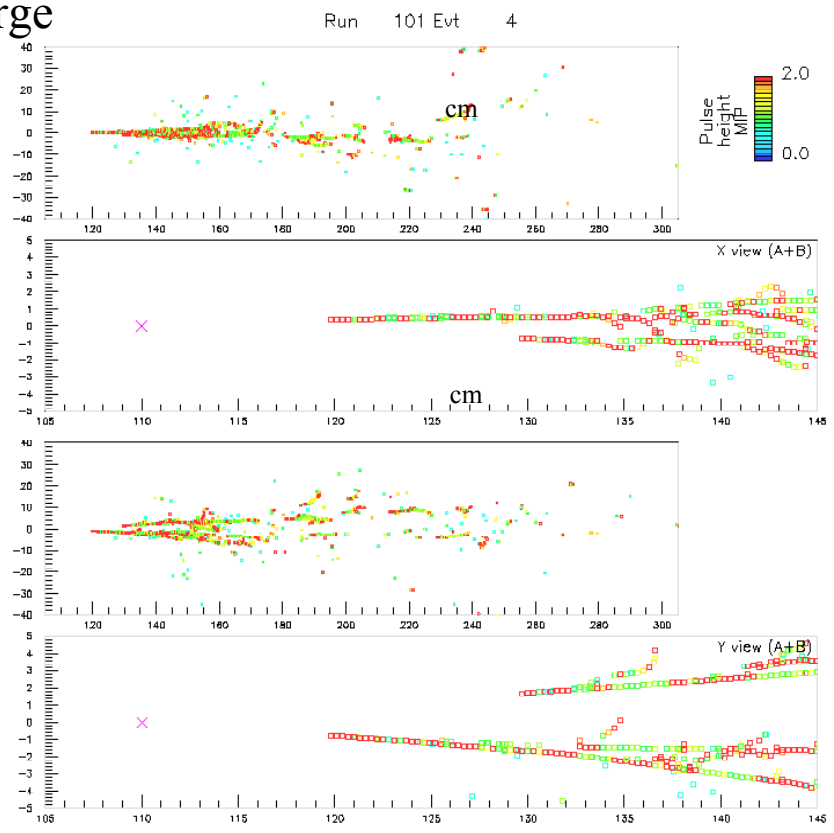


# Electrons versus $\pi^0$ 's at 1.5 GeV in a Liquid Argon TPC

Dot indicates hit- colour indicates collected charge  
 green=1 mip, red=2 mips



X plane  
 ↓  
 zoom in  
 Y plane  
 ↓  
 zoom in

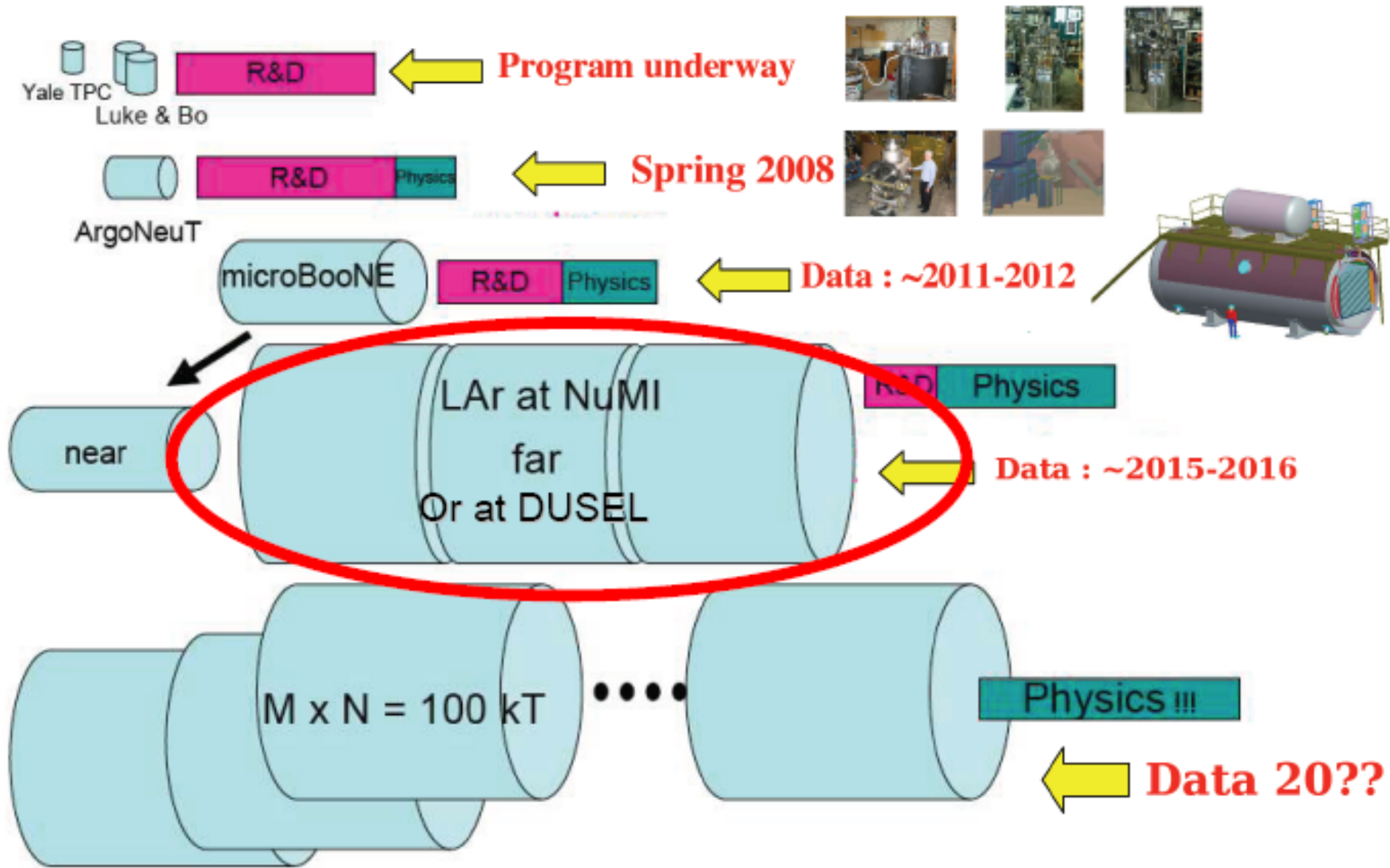


**Electrons-** Single track (mip scale) starting from a single vertex

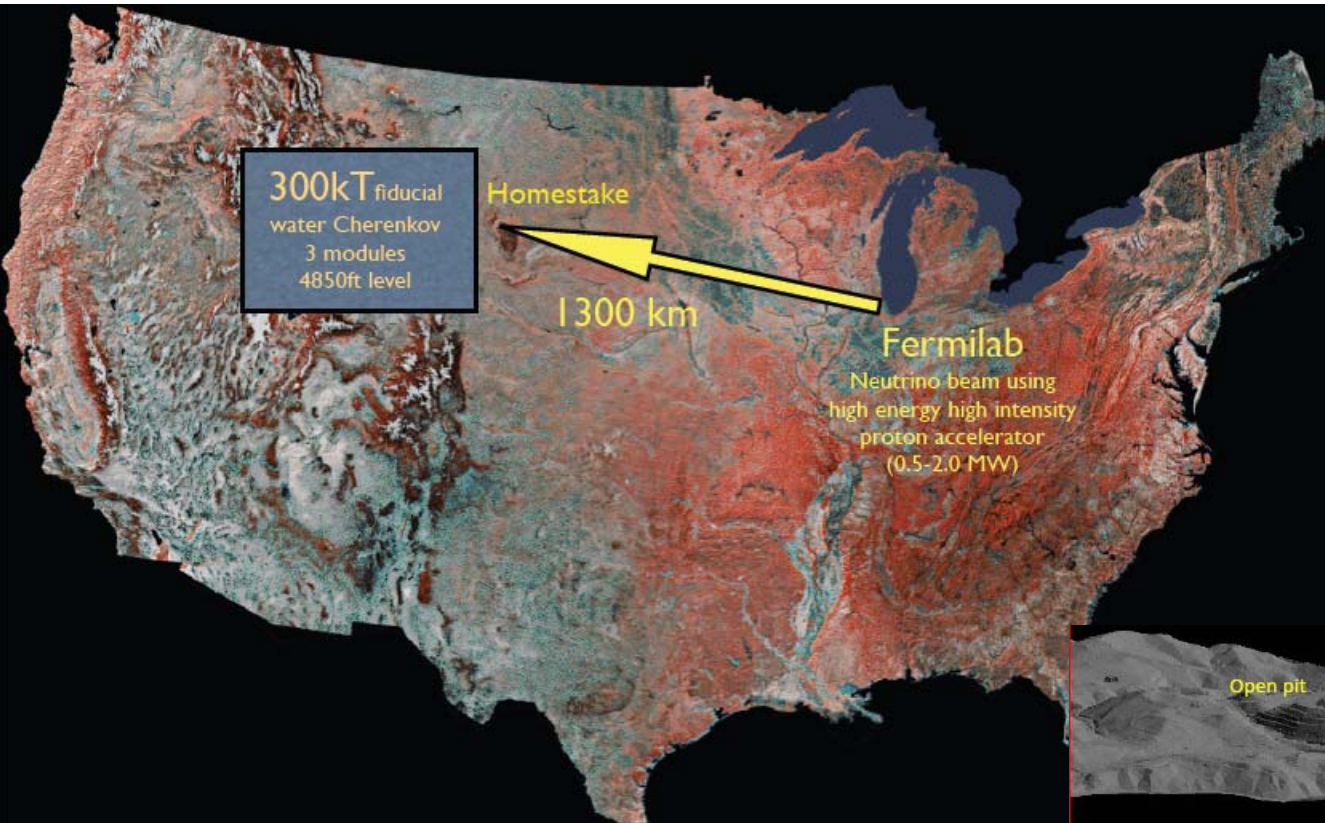
Multiple secondary tracks can be traced back to the same primary vertex

Use both topology and dE/dx to identify interactions

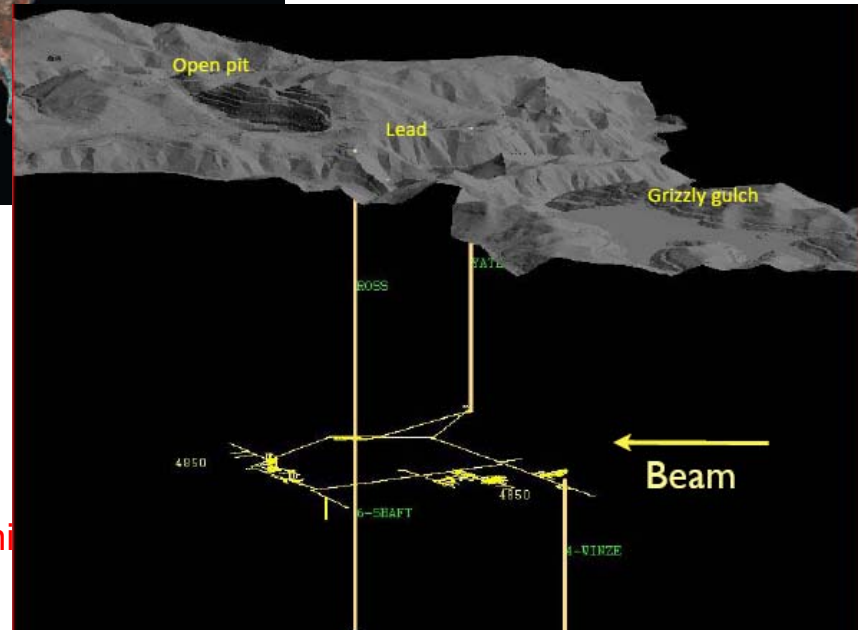
# Liquid Argon TPC R&D Path in the US



# The Future

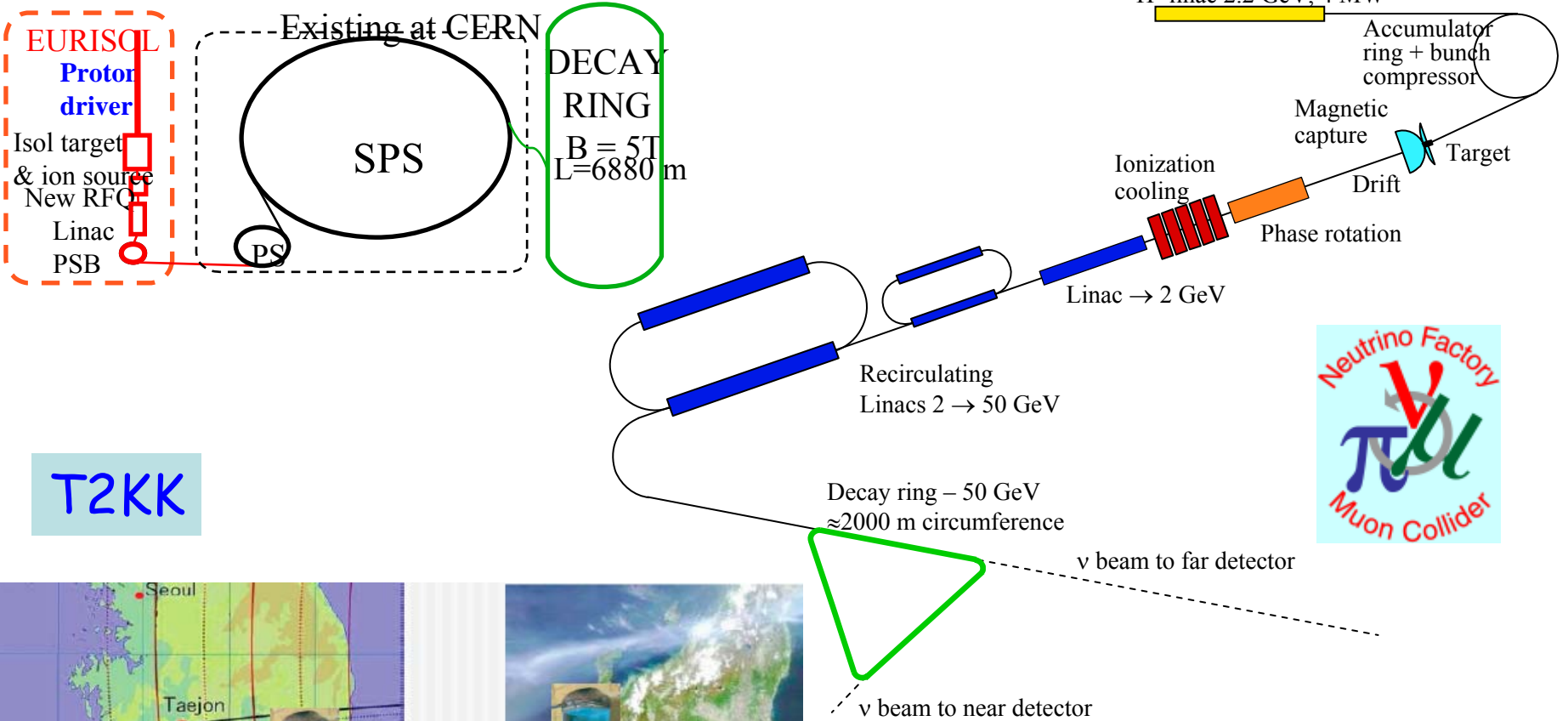


In the US, a wideband  $\nu_\mu$  beam from Fermilab to DUSEL (Homestake)

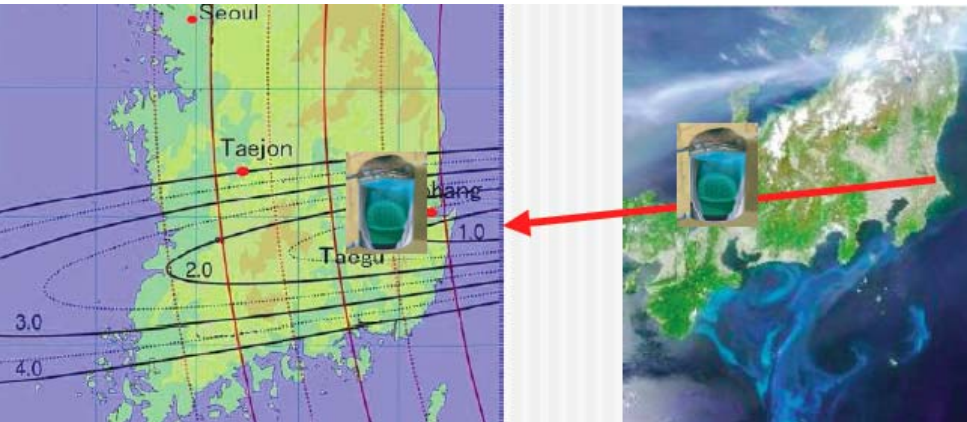


Beta-Beam

# Far (ever?) Future



T2KK



# Conclusions

- Long-baseline neutrino accelerator experiments will always have the advantage of control of the neutrino source
- There are a number of upcoming experiments designed to measure/improve the limit on  $\theta_{13}$  but, of course, one is constrained by a complete lack of knowledge of this angle. If it has a high value (near the Chooz limit) then “traditional” long-baseline experiments have a chance but if it is “small” then really only a Neutrino Factory has a chance of making useful measurements.