

(1)

Angular momentum  $\vec{L}$

(i)  $\vec{L}$  of a pointlike particle defined by

$$\vec{L} = \vec{r} \times \vec{p}$$

e.g. for the ball attached to massless rod discussed above we have:

$$|\vec{L}| = |\vec{r}| |\vec{p}| |\sin \phi|$$

$$= r m v \quad (\phi = \frac{\pi}{2} \text{ for circular motion})$$

$$= mr^2 \omega = I\omega$$

orientation: RH rule  $\leftarrow \vec{L} = \vec{I} \vec{\omega}$

Note that for uniform circular motion  $\vec{\omega} = \text{const} \Rightarrow \vec{L} = \text{constant}$ .

Now recall Newton-2:  $\frac{d}{dt} \vec{p} = \vec{F}$   
(see discussion of impulse theorem)

Try something similar with  $\vec{L}$ : product rule

$$\begin{aligned} \leftarrow \frac{d}{dt} \vec{L} &= \frac{d}{dt} (\vec{r} \times \vec{p}) \stackrel{!}{=} \frac{d\vec{r}}{dt} \times \vec{p} + \vec{r} \times \frac{d\vec{p}}{dt} \\ &= m \underbrace{\vec{v} \times \vec{v}}_{=0} + \vec{r} \times \vec{F} = \vec{\tau} \end{aligned}$$

Summary: 
$$\boxed{\frac{d}{dt} \vec{L} = \vec{\tau}}$$
 angular-momentum theorem

We can infer a conservation law:

if  $\vec{\tau} = 0 \Rightarrow \vec{L}$  conserved

$\vec{\tau} = 0$  if  $\begin{cases} \vec{F} = 0 & (\text{trivial}) \\ \vec{r} \parallel \vec{F} & (\text{e.g. uniform circular motion}) \end{cases}$

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(ii)  $\vec{L}$  for a particle system (extended object)

define  $\vec{L}_{\text{tot}} = \sum_i \vec{r}_i \times \vec{p}_i = \sum_i (m_i \vec{r}_i \times \vec{v}_i)$   
 $\vec{v}_{\text{tot}} = \sum_i \vec{r}_i \times \vec{F}_i$

show (similarly as above):  $\frac{d}{dt} \vec{L}_{\text{tot}} = \vec{T}_{\text{tot}}$  \*)

For rotations about a fixed axis, one finds (again):

$$\vec{L}_{\text{tot}} = I \vec{\omega} \quad \text{with } I = \sum_i m_i r_i^2$$

Examples for angular momentum conservation

(i) Spinning figure skater

$$\begin{aligned} L_{\text{initial}} &= I_{\text{initial}} \omega_{\text{initial}} \\ &= L_{\text{final}} = I_{\text{final}} \omega_{\text{final}} \end{aligned} \quad \left[ \begin{array}{l} \vec{r}=0 \\ \Rightarrow \end{array} \right] \quad \boxed{\omega_{\text{final}} = \frac{I_{\text{initial}}}{I_{\text{final}}} \omega_{\text{initial}}}$$

start with a small angular velocity  $\omega_{\text{initial}}$ , then pull arms and legs in to reduce  $I \Rightarrow$  end up with a higher angular velocity  $\omega_{\text{final}}$

Consider KE:

$$\begin{aligned} KE_g &= \frac{1}{2} I_{\text{final}} \omega_{\text{final}}^2 = \frac{1}{2} I_{\text{final}} \left( \frac{I_{\text{initial}}}{I_{\text{final}}} \omega_{\text{initial}} \right)^2 \\ &= \left( \frac{I_{\text{initial}}}{I_{\text{final}}} \right) \frac{1}{2} I_{\text{initial}} \omega_{\text{initial}}^2 = \boxed{\frac{I_{\text{initial}}}{I_{\text{final}}} KE_i} > 1 \end{aligned}$$

$$\therefore KE_g > KE_i$$

KE is not conserved, but increased (the skater does positive work when pulling in arms + legs)

\*) Angular momentum conservation, if  $\vec{T}_{\text{tot}} = 0$ :

$$\vec{L}_{\text{tot}} = \vec{L}_i = \vec{L}_g = \text{const}$$

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### (ii) Planetary motion

Recall: Earth orbits Sun on an elliptical orbit  $\Rightarrow$

$$\vec{F}_{\text{SunE}} = -\frac{GM_{\text{Earth}}}{r^3} \vec{r}$$

$\vec{r}$  Earth's position vector relative to resting sun

Q: does Sun exert a torque on Earth?

$$\vec{\tau}_0 = \vec{r} \times \vec{F}_{\text{SunE}} = -\frac{GM_{\text{Earth}}}{r^3} \underbrace{\vec{r} \times \vec{r}}_{=0} = 0$$

A: no  $\Rightarrow \vec{L}_{\text{Earth}} = \text{const}$

This has two important implications:

(i) Direction of  $\vec{L}_{\text{Earth}}$  never changes  $\Rightarrow$   
orbit is truly planar

(ii) Magnitude  $|\vec{L}_{\text{Earth}}|$  never changes;

one can show that Kepler's area law follows from this fact

(see book, Chap 9.4)

So, Sun doesn't exert a torque on Earth. It does do work though. For Earth's (slightly) elliptical orbit  $\vec{F}_{\text{grav}}$  is not always perpendicular to displacement (unlike for uniform circular motion).

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\* Assume Sun is infinitely heavy, i.e.  $M_s \rightarrow \infty$   
(such that it is at rest despite Newton-3.)