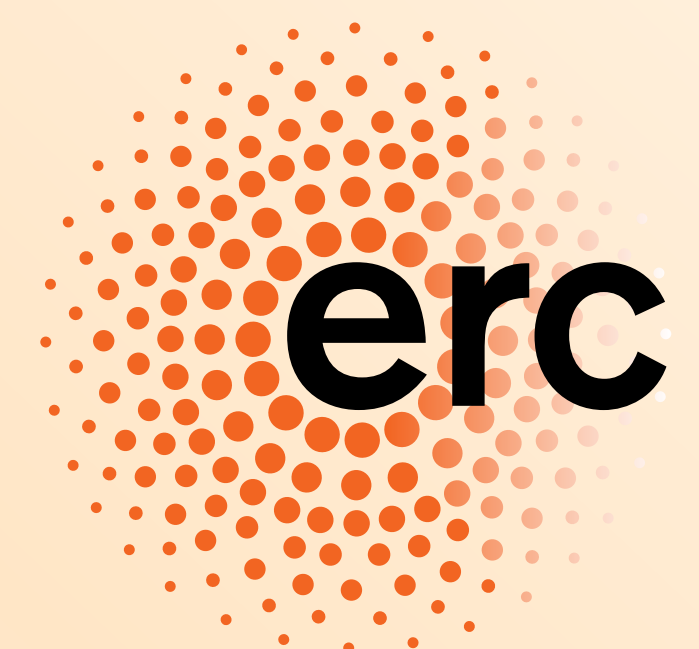


A new approach to cosmological simulations

Andrew Pontzen, UCL
gmgalaxies.org



GMGalaxies team & collaborators



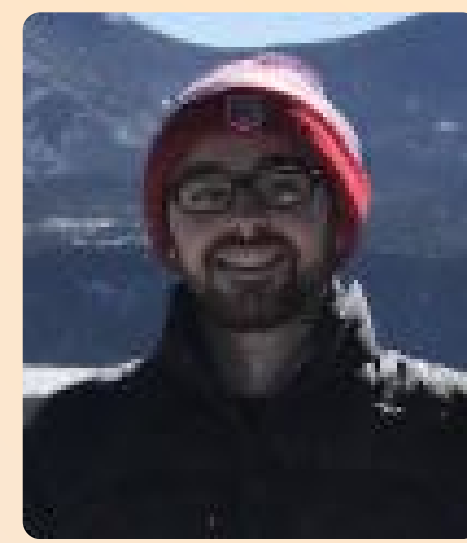
Oscar Agertz,
Lund



Alyson Brooks,
Rutgers



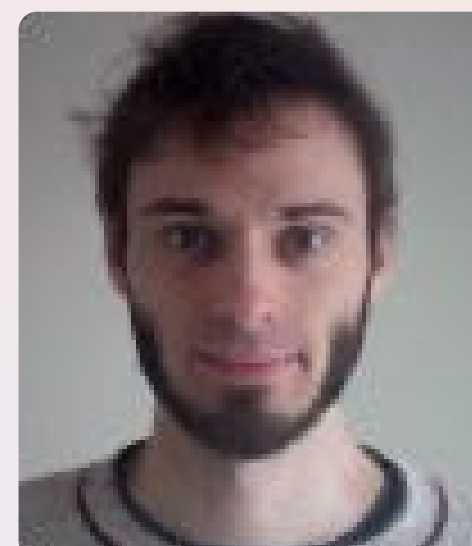
Corentin Cadiou,
UCL



Jonathan Davies,
LJMU → UCL



Luisa Lucie-Smith,
UCL → MPA



Matthew Orkney,
Surrey



Hiranya Peiris,
UCL



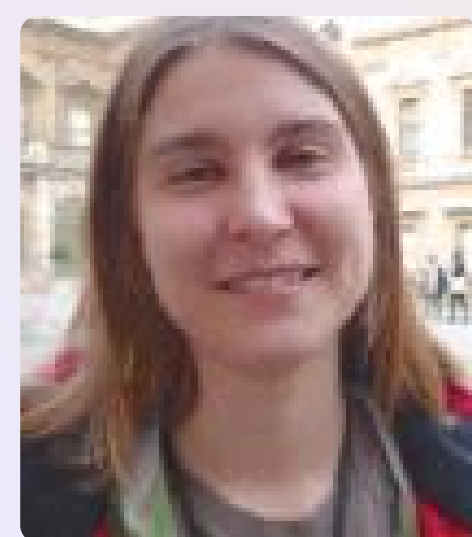
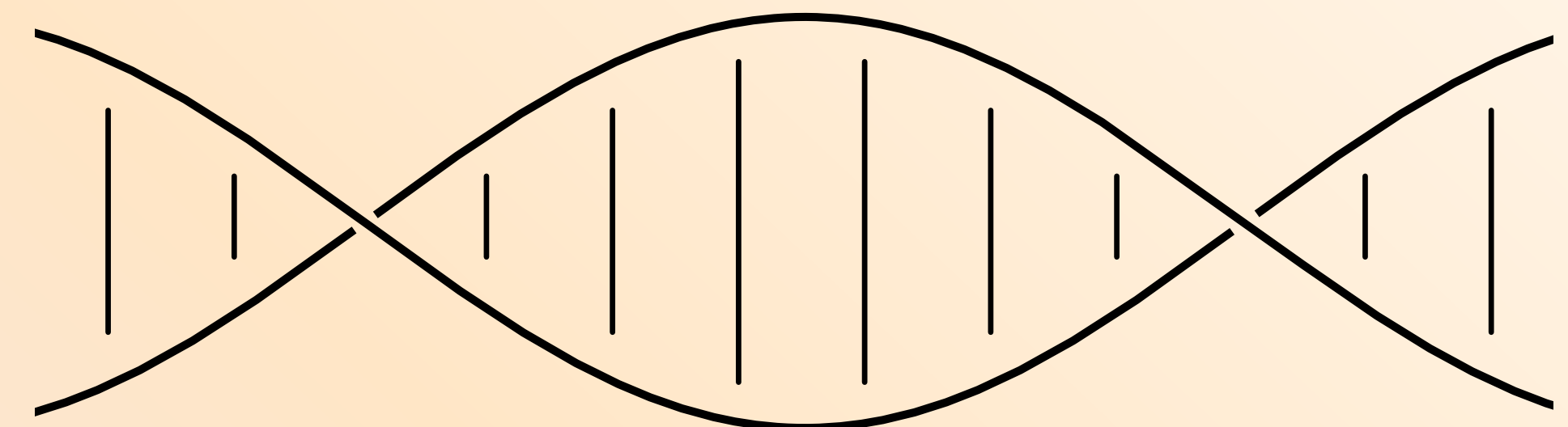
Andrew Pontzen,
UCL



Justin Read,
Surrey



Martin Rey,
UCL → Lund



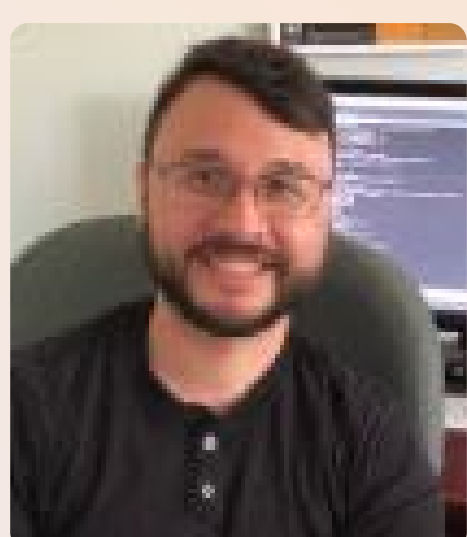
Nina Roth,
UCL → Industry



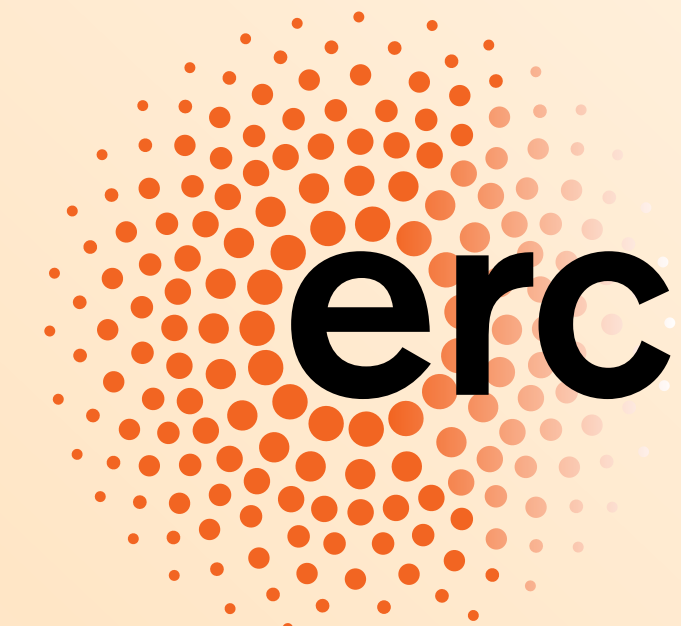
Nicole Sanchez,
U. Washington

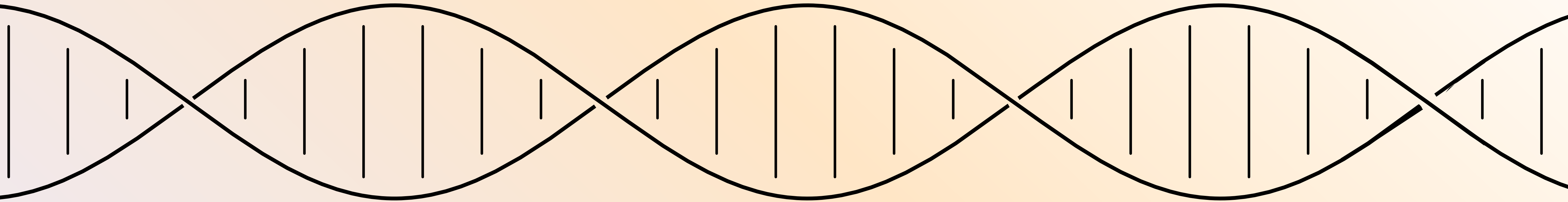


Stephen Stopyra,
UCL



Michael Tremmel,
Yale



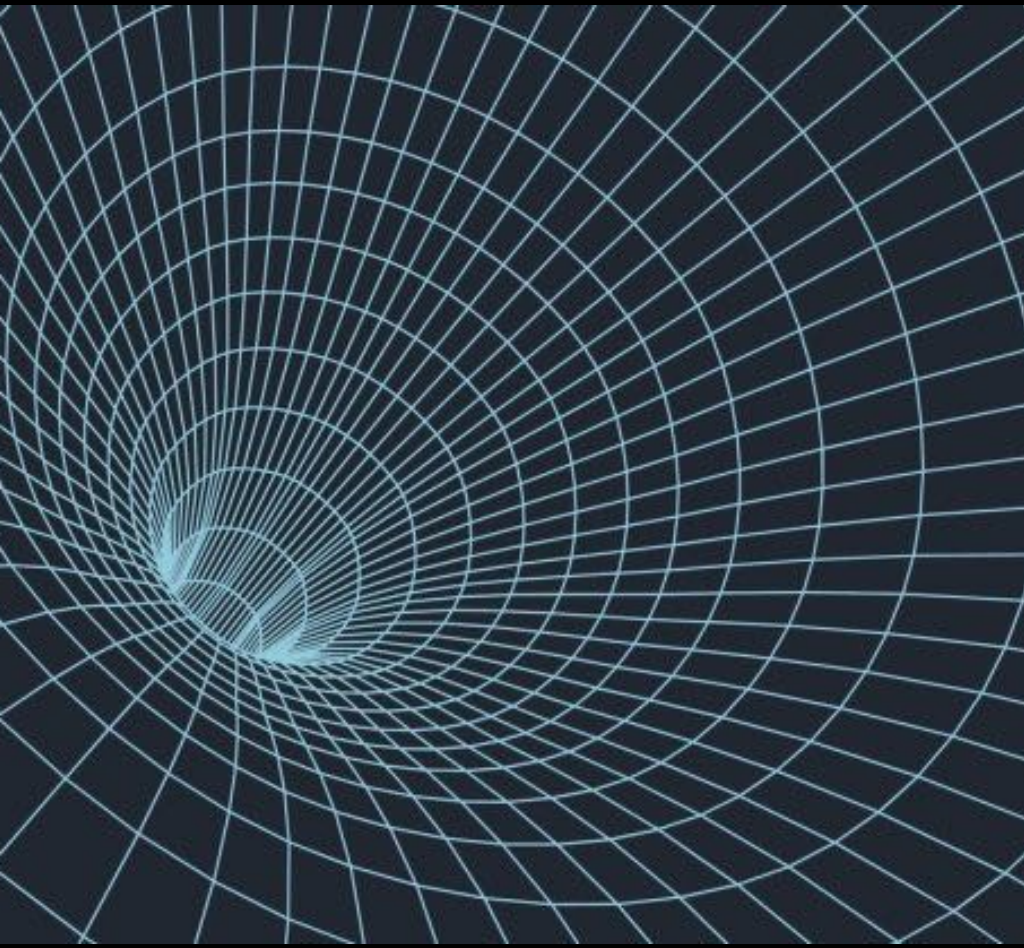


Part 1: Why dwarf galaxies are interesting

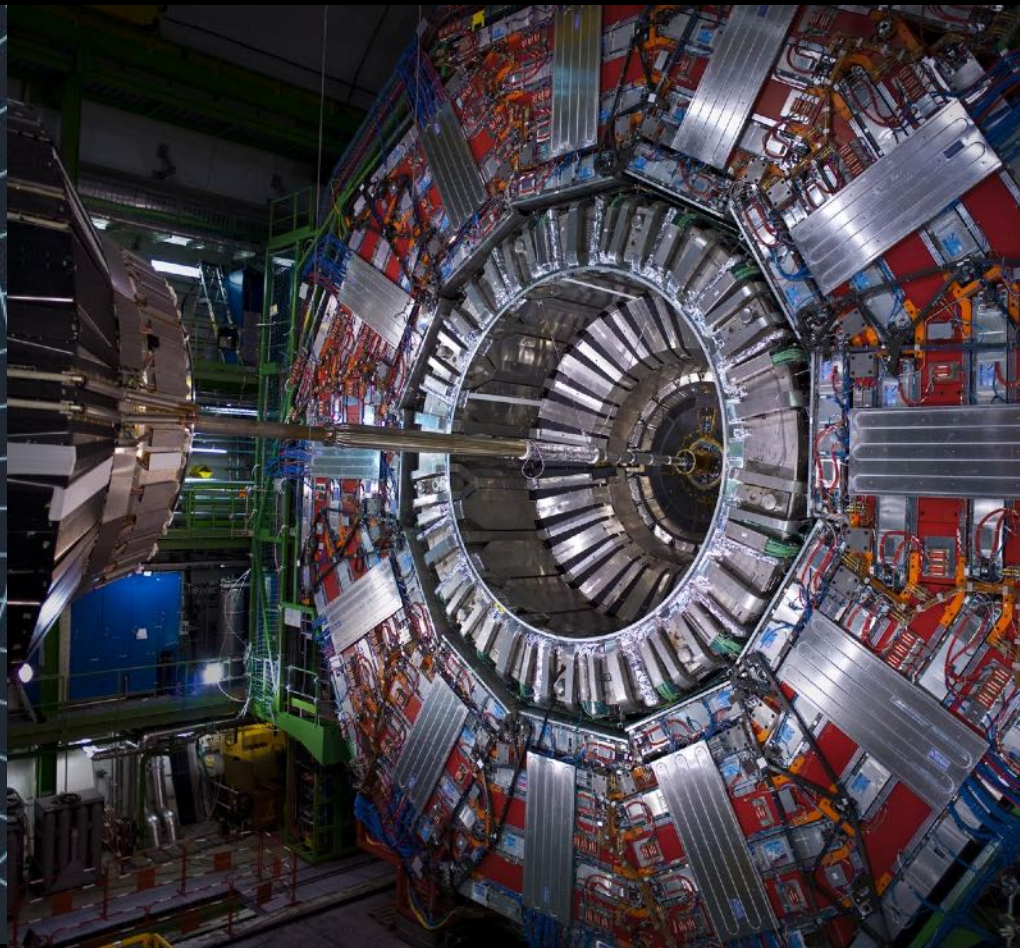
Part 2: “Genetically modified” simulations

Part 3: Results / predictions

Ingredients of Λ CDM cosmology



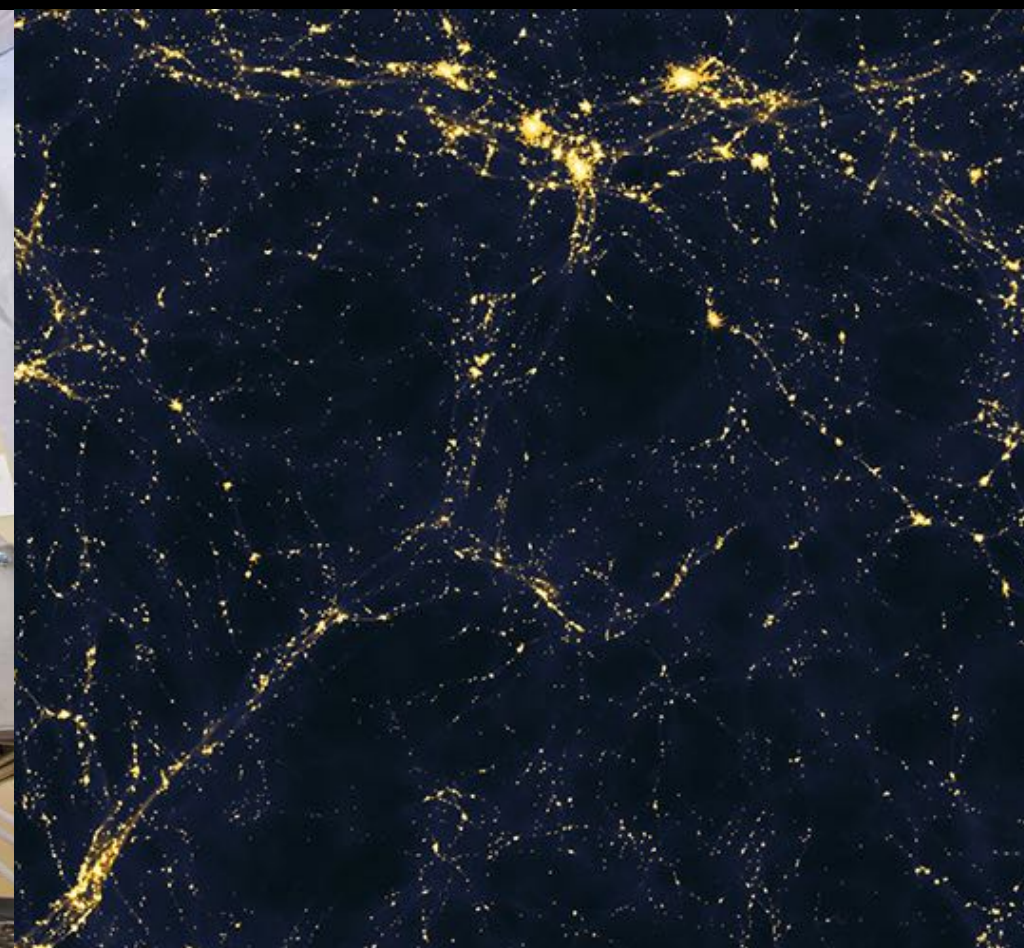
General
relativity



Standard model of
particle physics



Cold dark matter
(CDM)



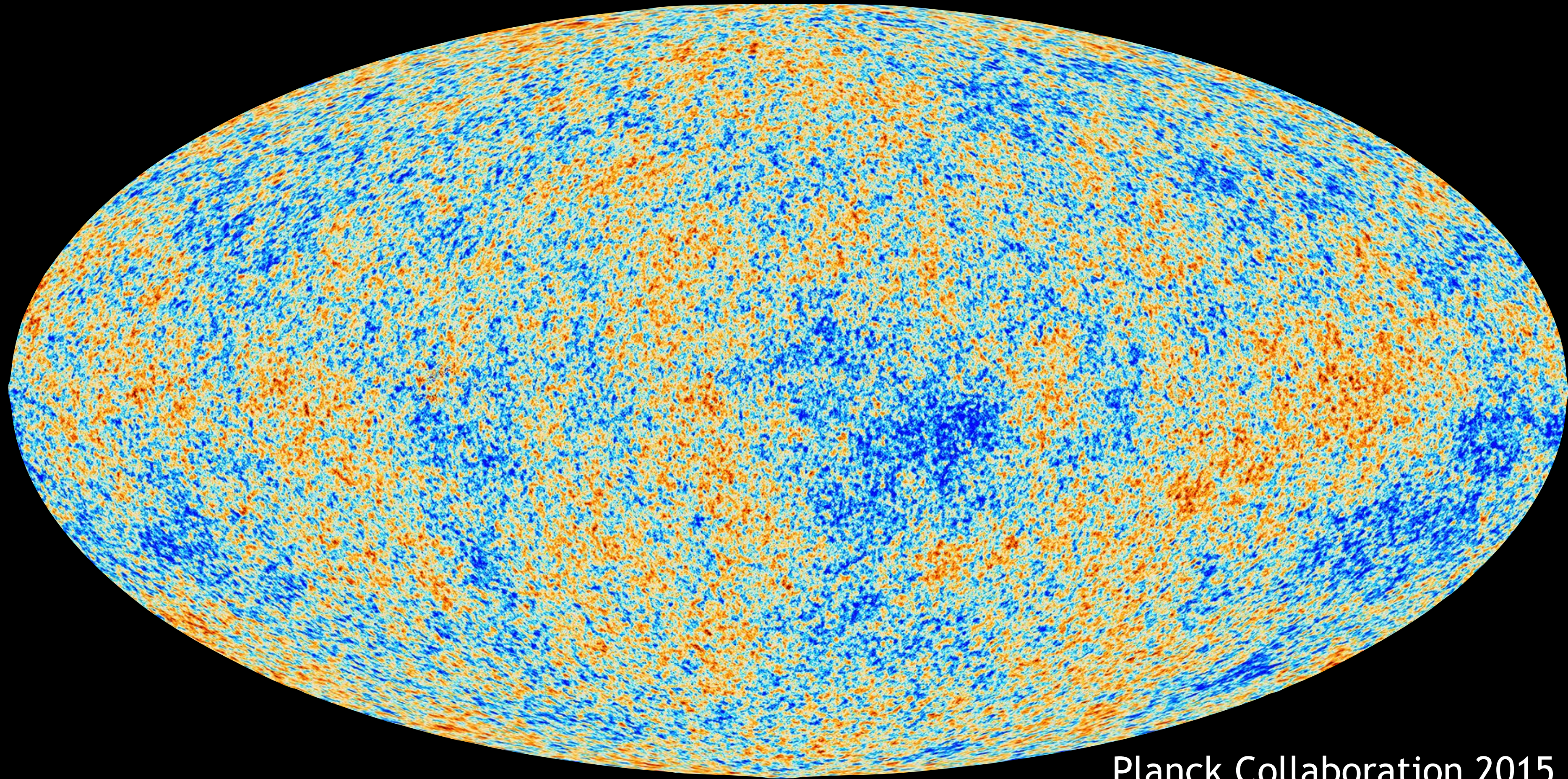
Dark energy (Λ)



Inflation

Highly successful phenomenology

Large scales/early times



Planck Collaboration 2015



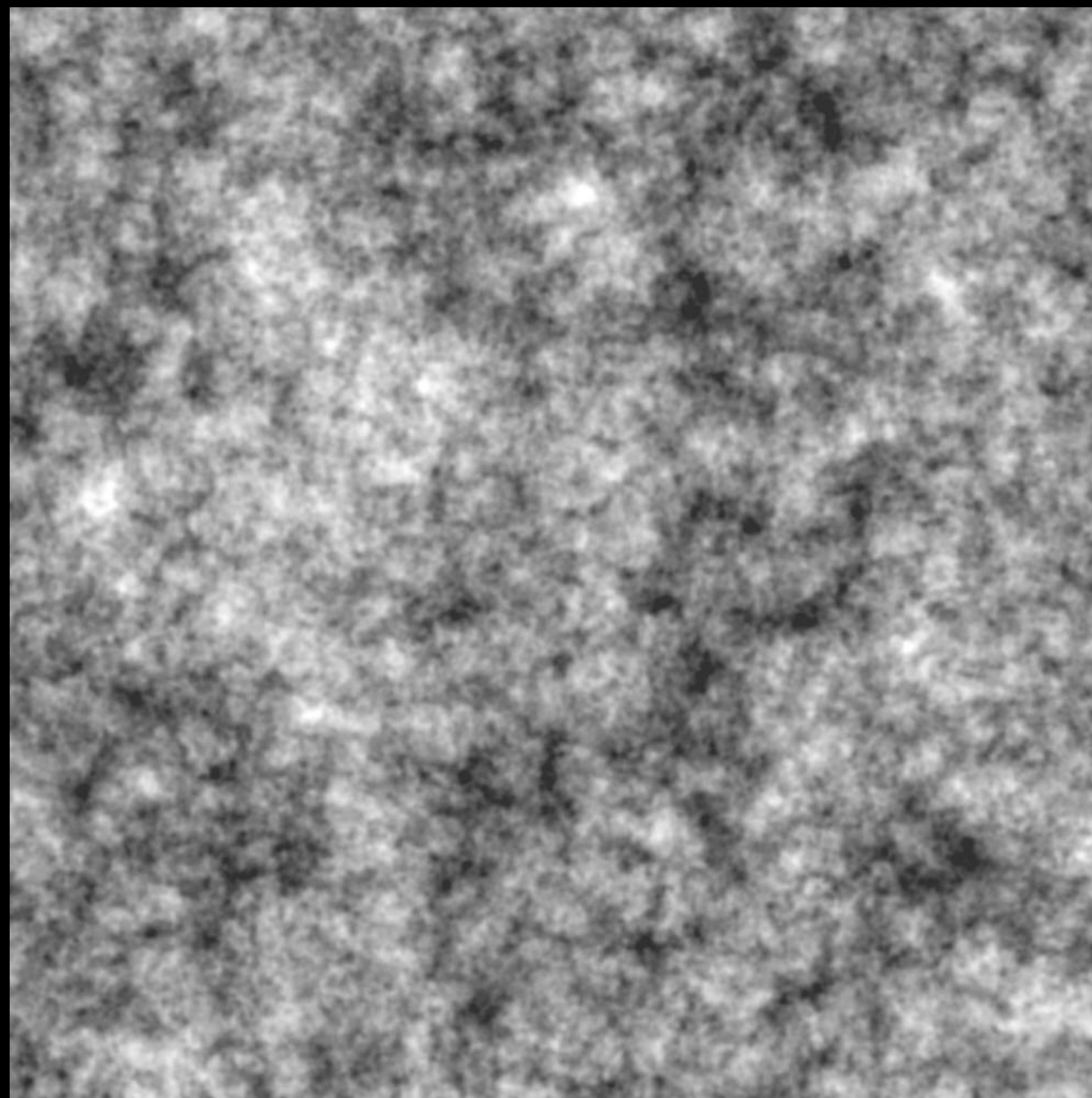
$2\pi d_A = 90$ Gpc comoving

Fluctuation scale: 10^{-5}

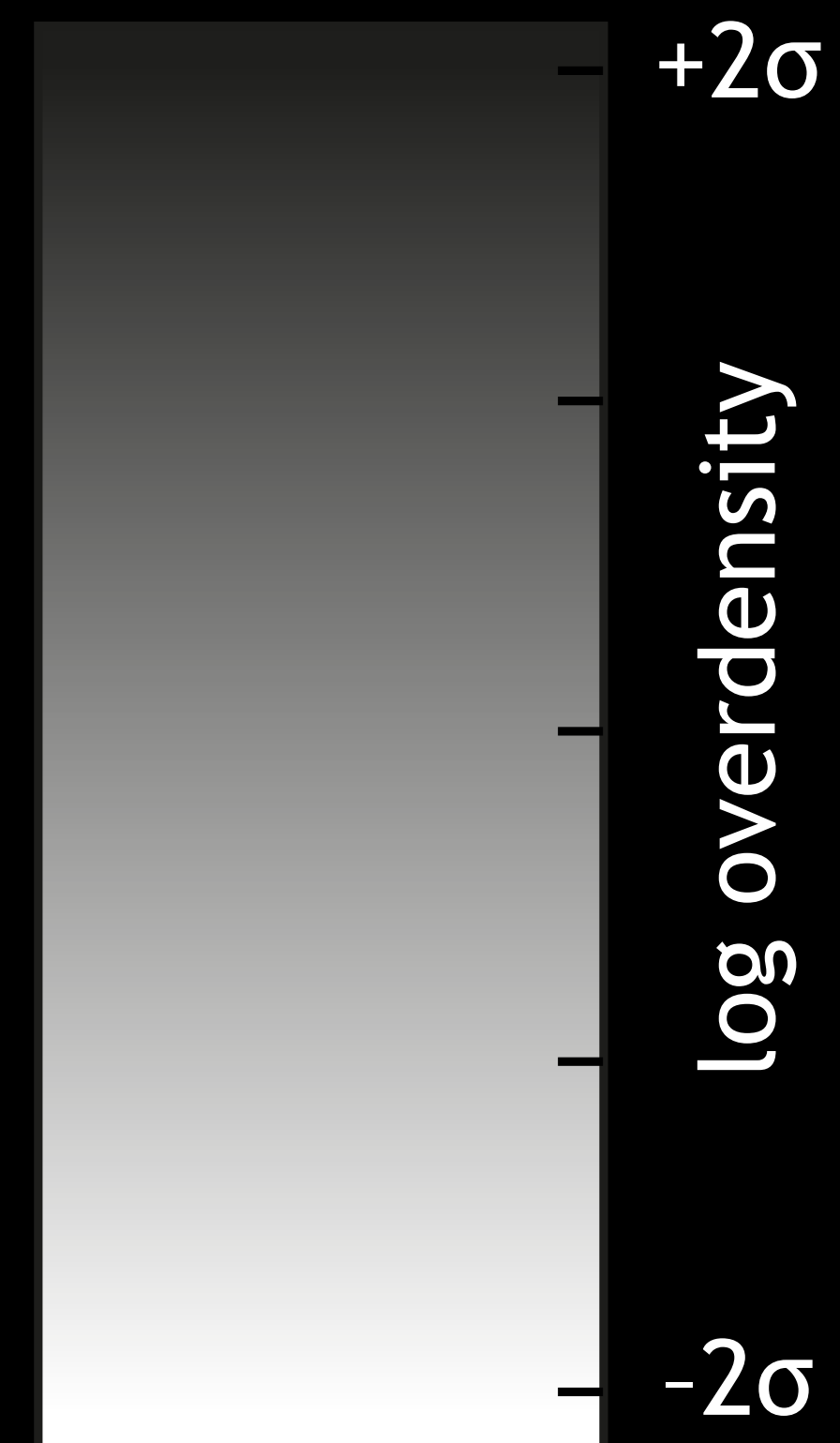
Small scales / late times

50 Mpc

Halos — 10^{+2}
Disks — 10^{+6}
Molecular clouds — 10^{+8}
Stars — 10^{+30}



← 200 comoving Mpc →



Pontzen+ 1511.04090

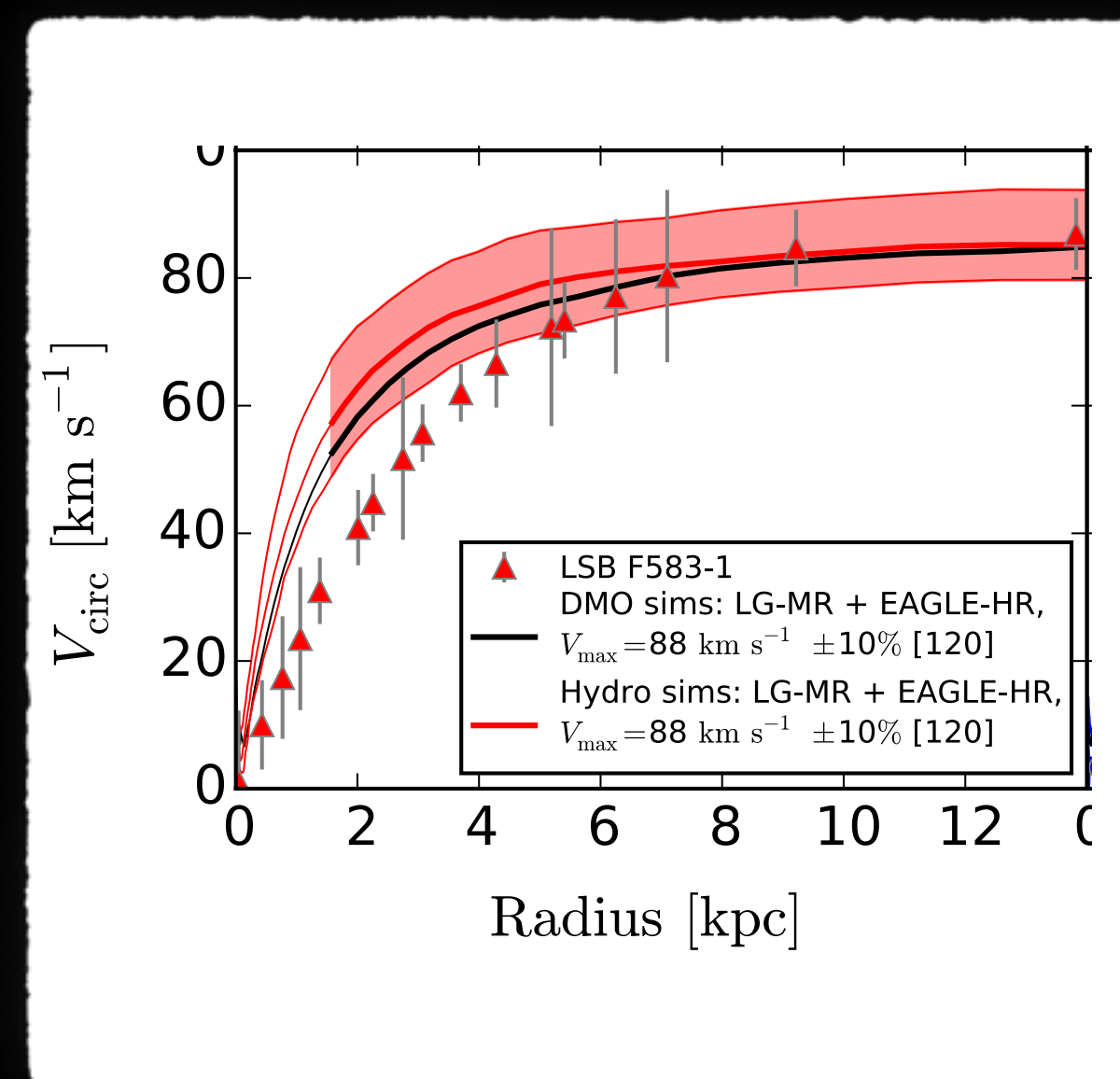
Why dwarf galaxies challenge Λ CDM



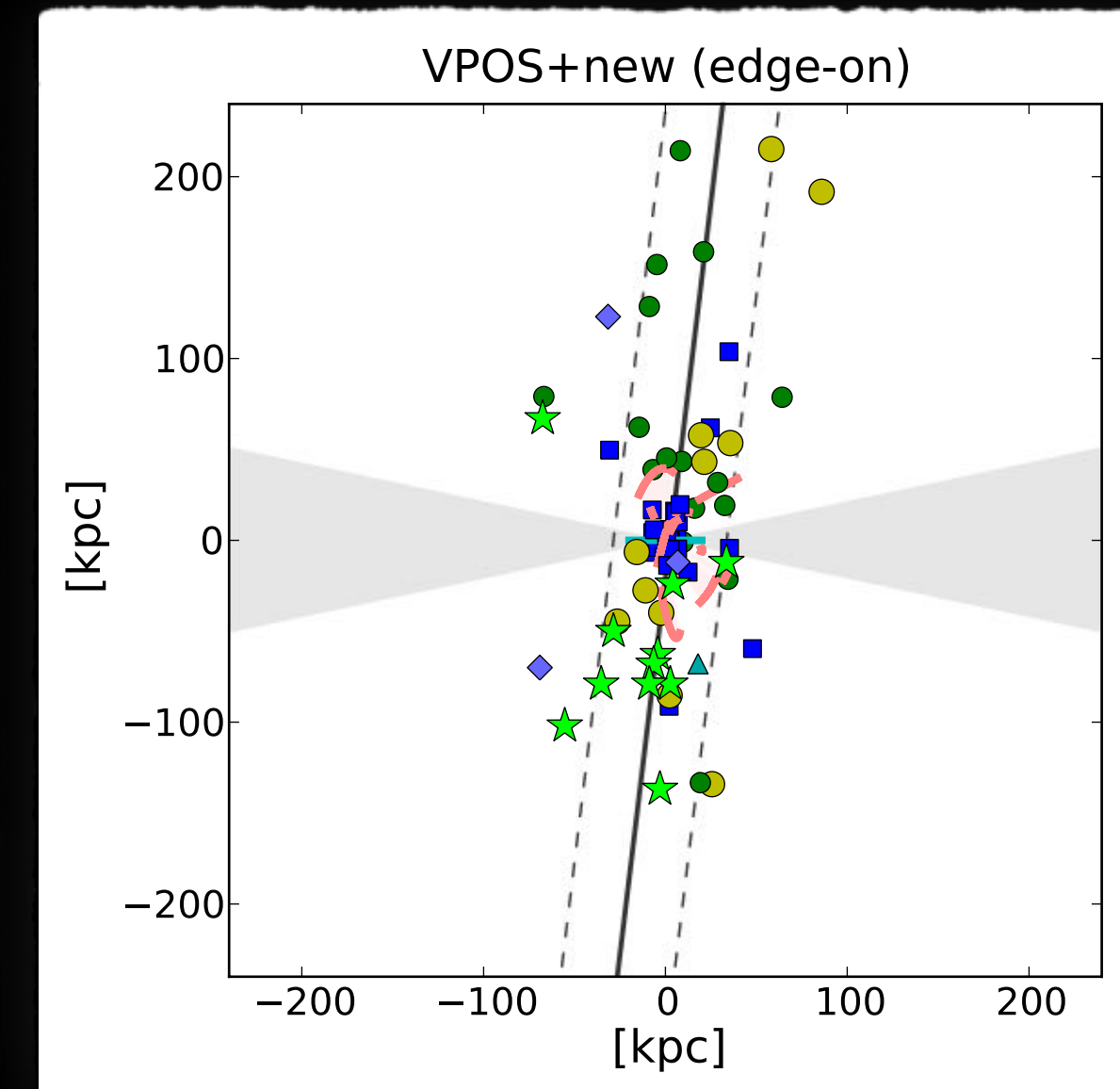
Dynamical masses



Abundance

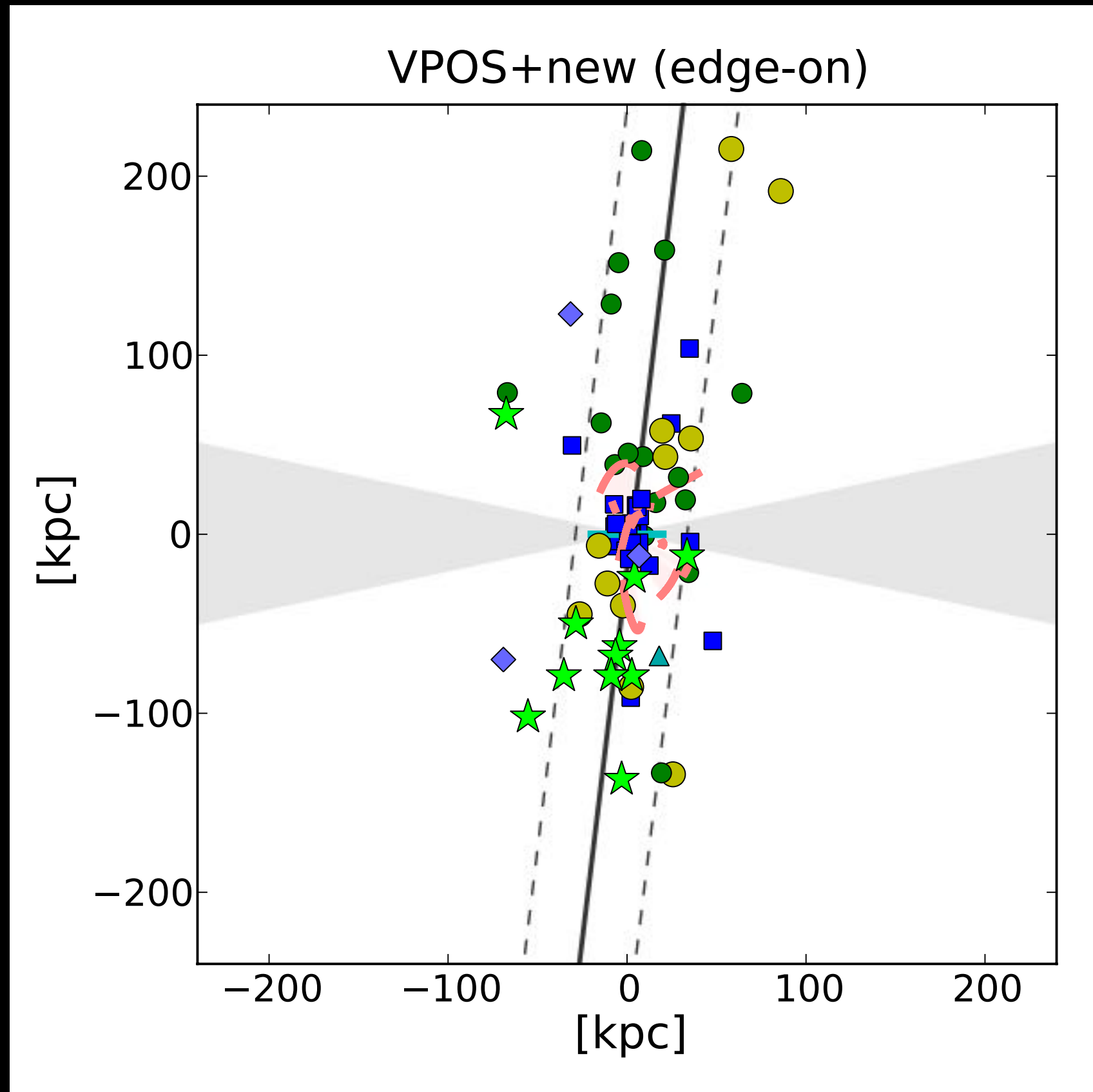


Central densities

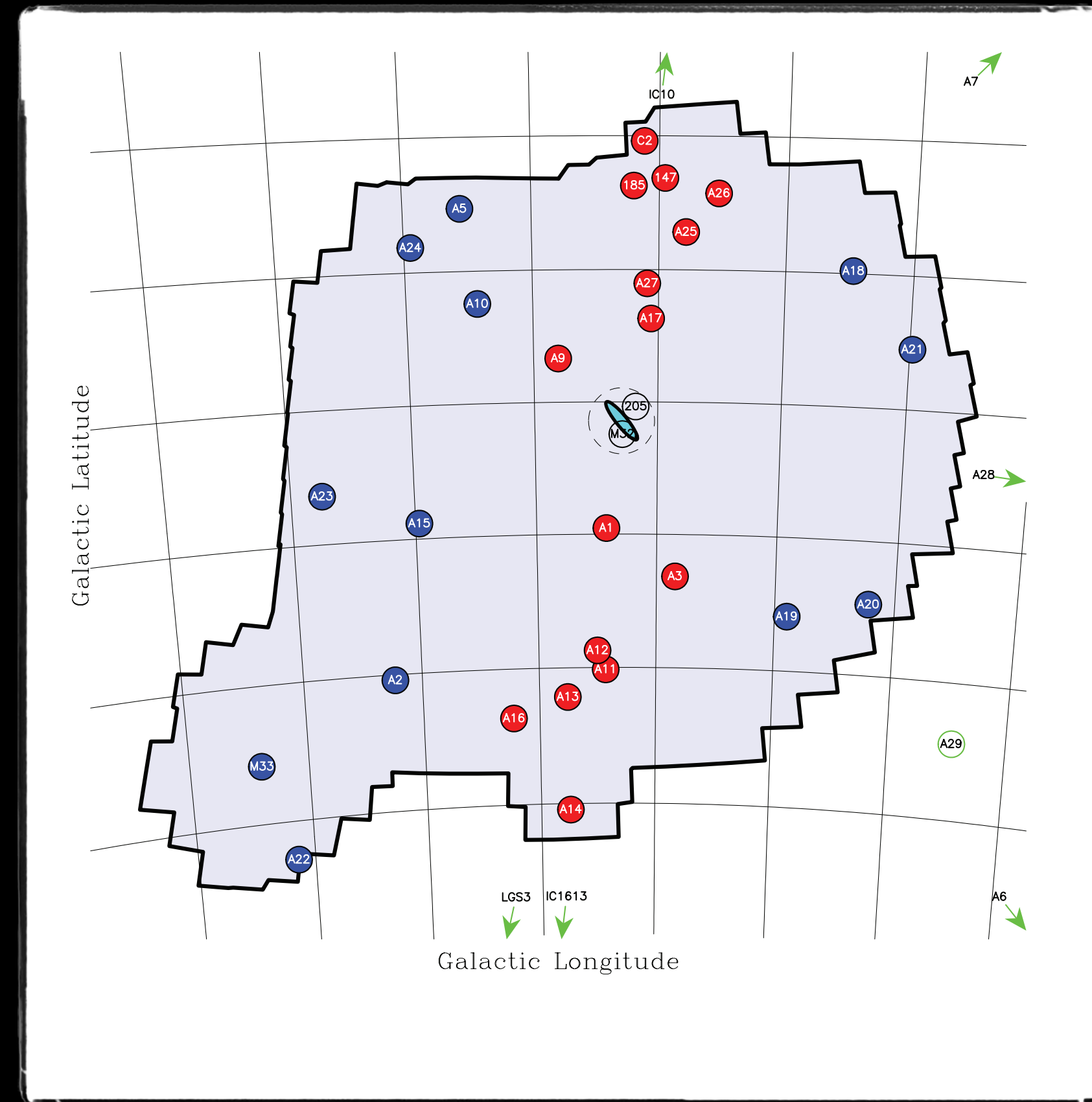


Alignments

Alignments



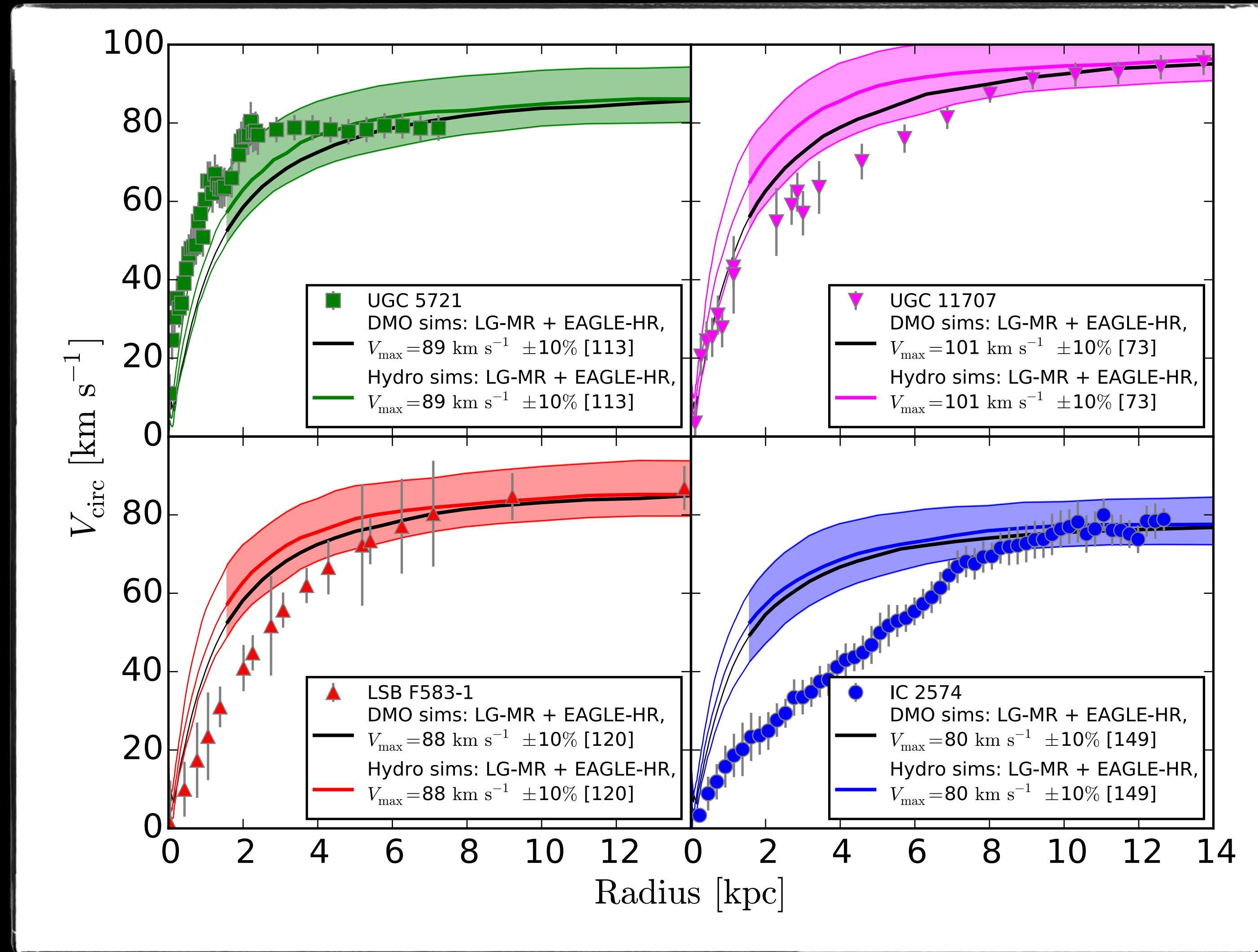
MW e.g. Pawlowski, McGaugh & Jerjen 15



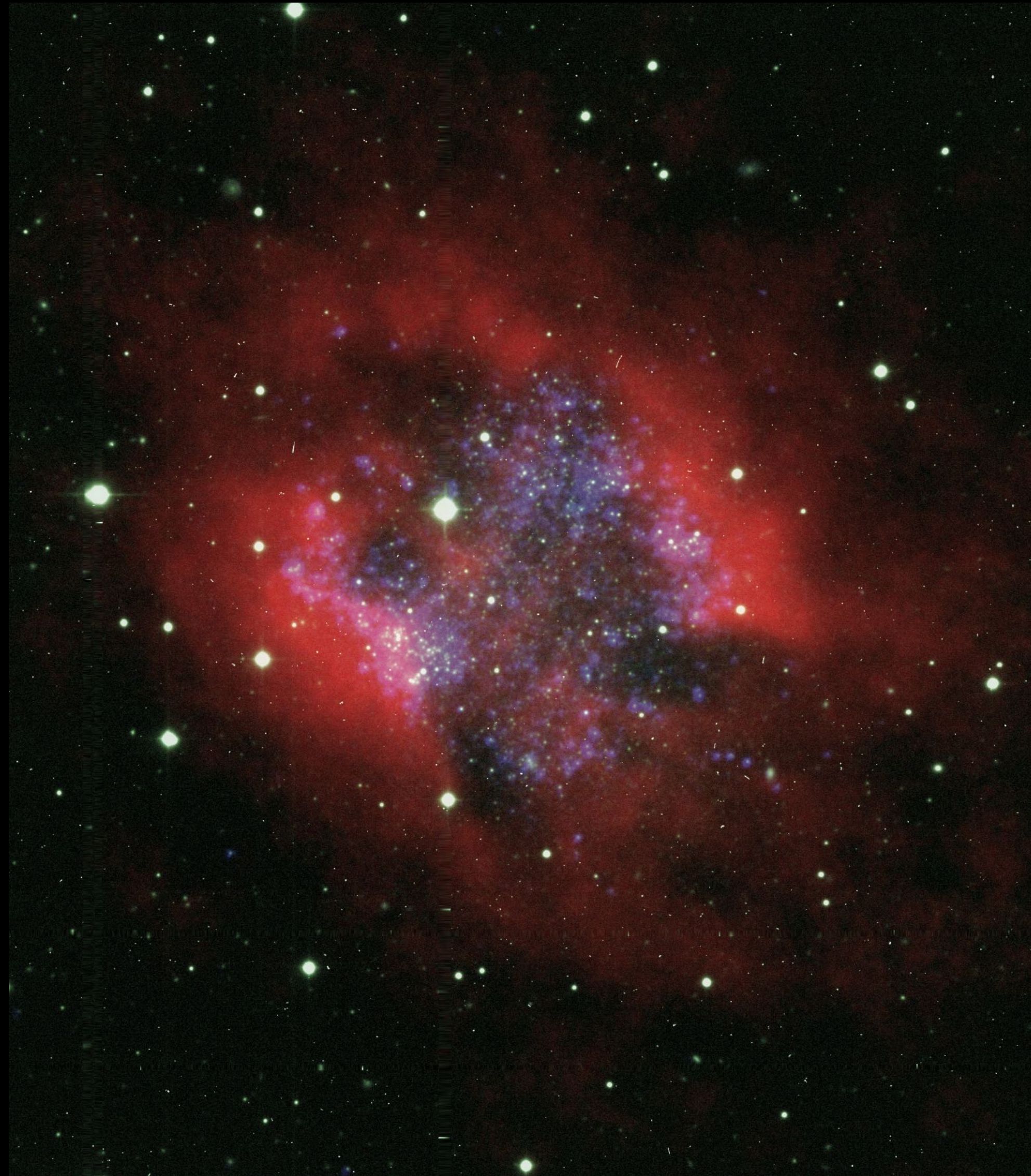
M31 e.g. Ibata+ 2013

Jury still out on whether this is a problem or not
e.g. Libeskind+ 2009; Buck, Dutton & Macciò 2016; Santos-Santos+ 2021

Central densities



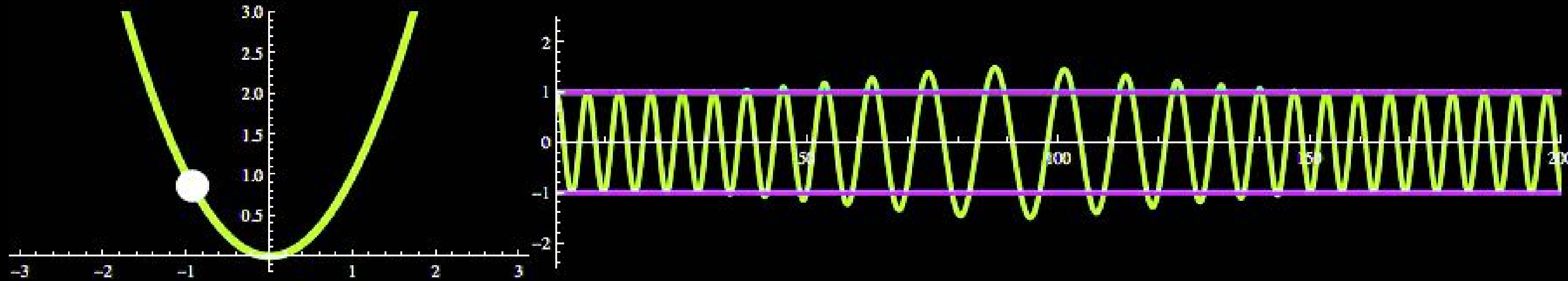
Central densities



DDO 75
“Little THINGS” survey
Hunter+ 2012

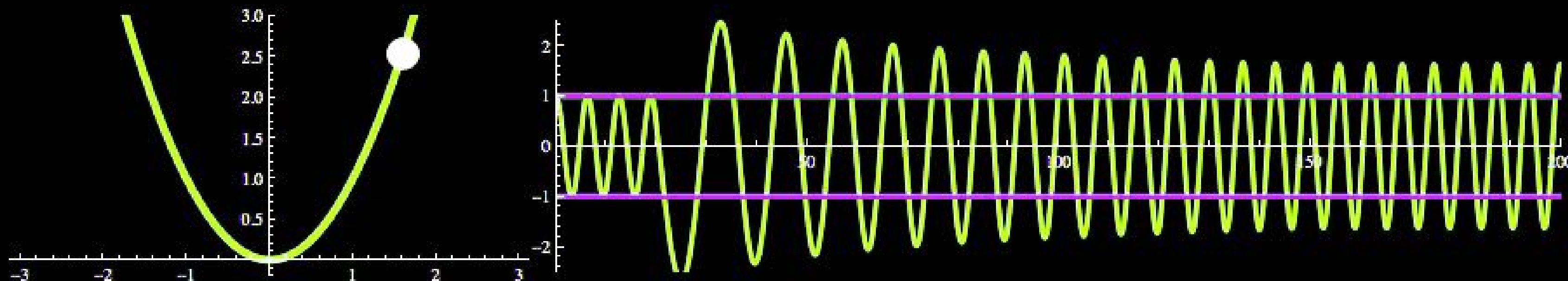
Composite image by
Lauren Hill

Central densities: slow vs fast feedback

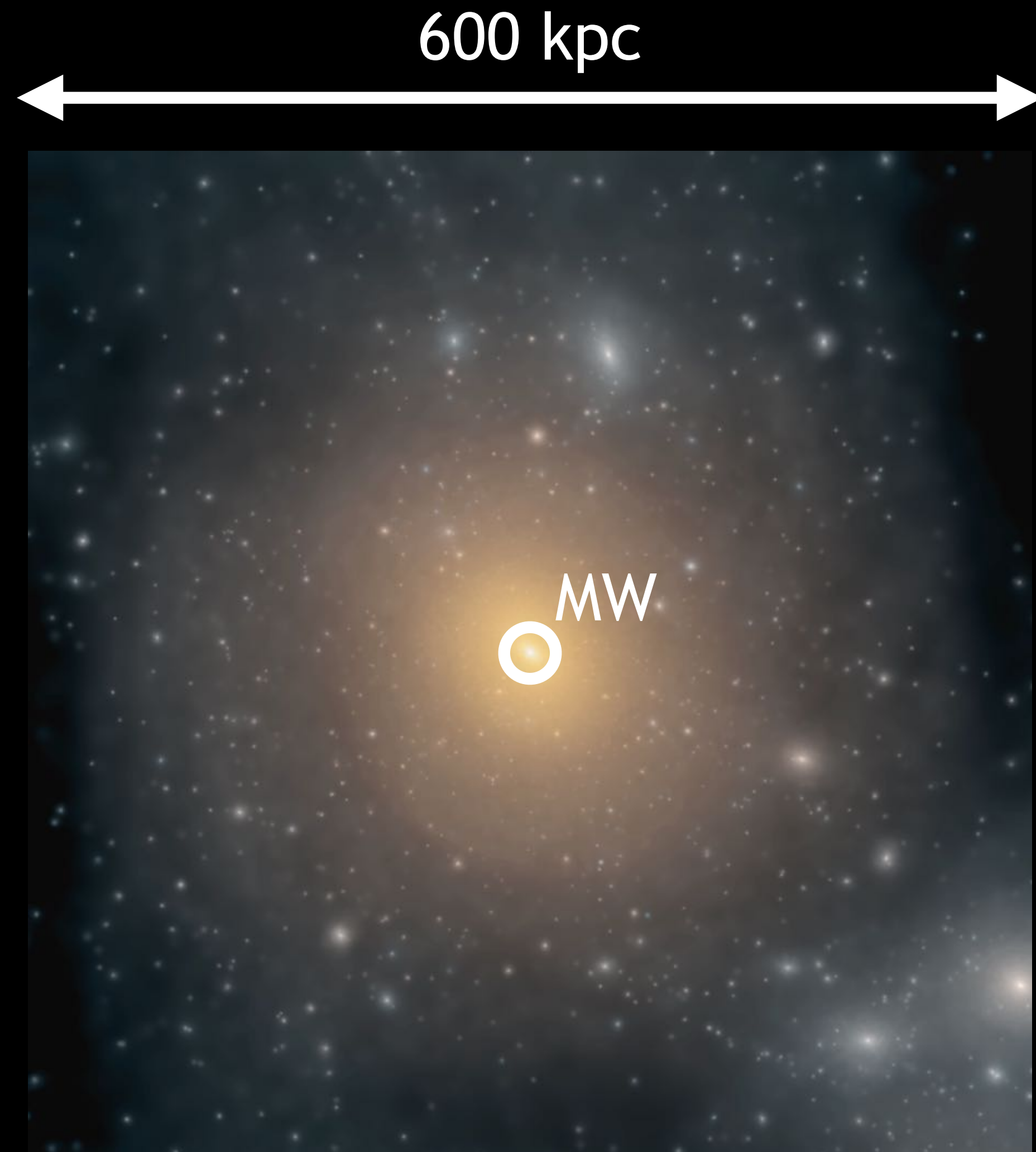


Slow outflows (adiabatic) $E_f = E_i$

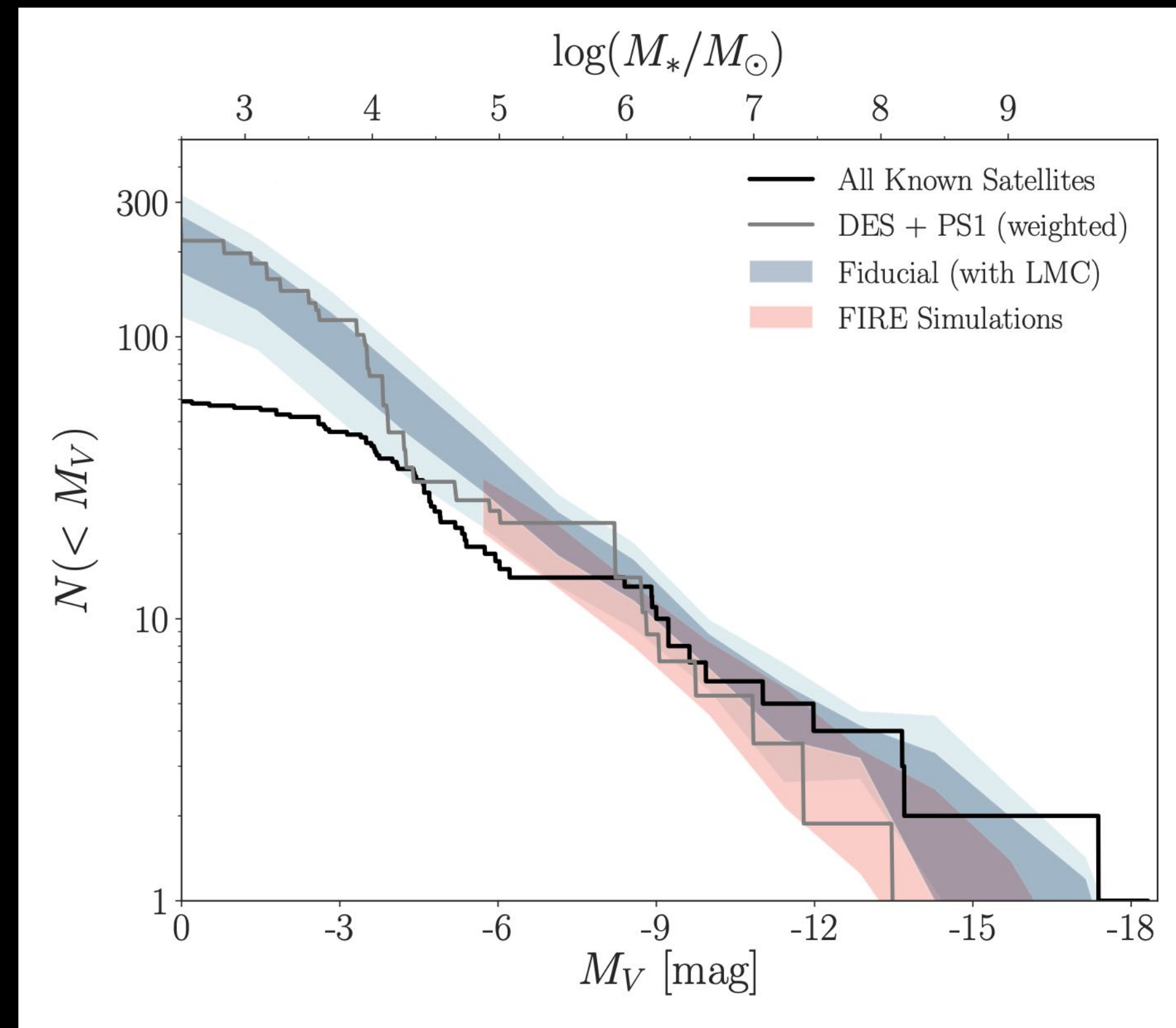
Fast outflows (impulsive) $\frac{\langle E_f \rangle}{E_i} = \frac{1}{2} \left(\frac{\omega_1}{\omega_0} + \frac{\omega_0}{\omega_1} \right)$



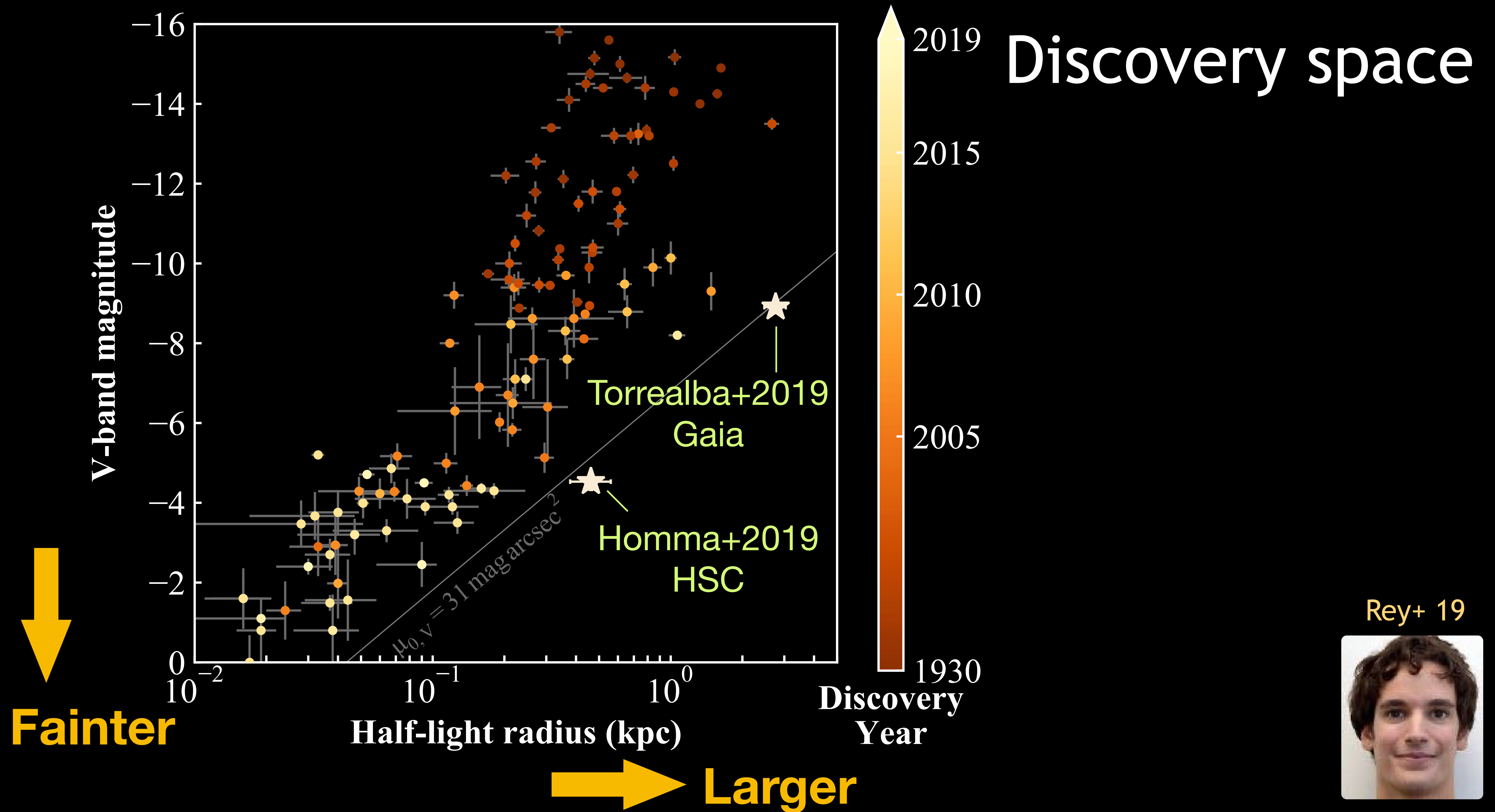
Missing satellites

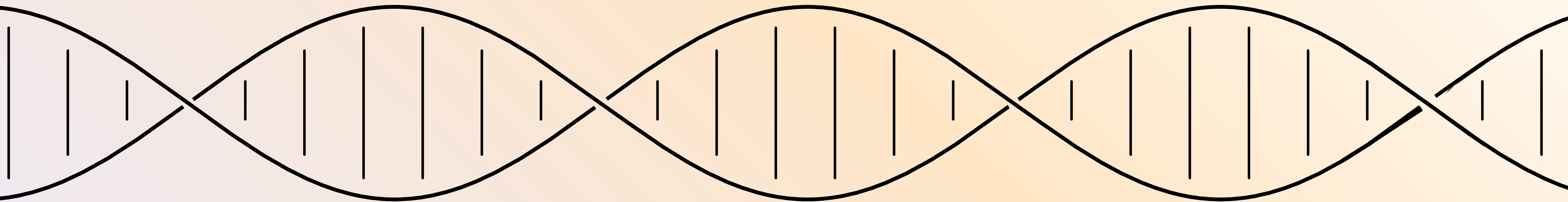


Pontzen+ 2013



Nadler+ 2020

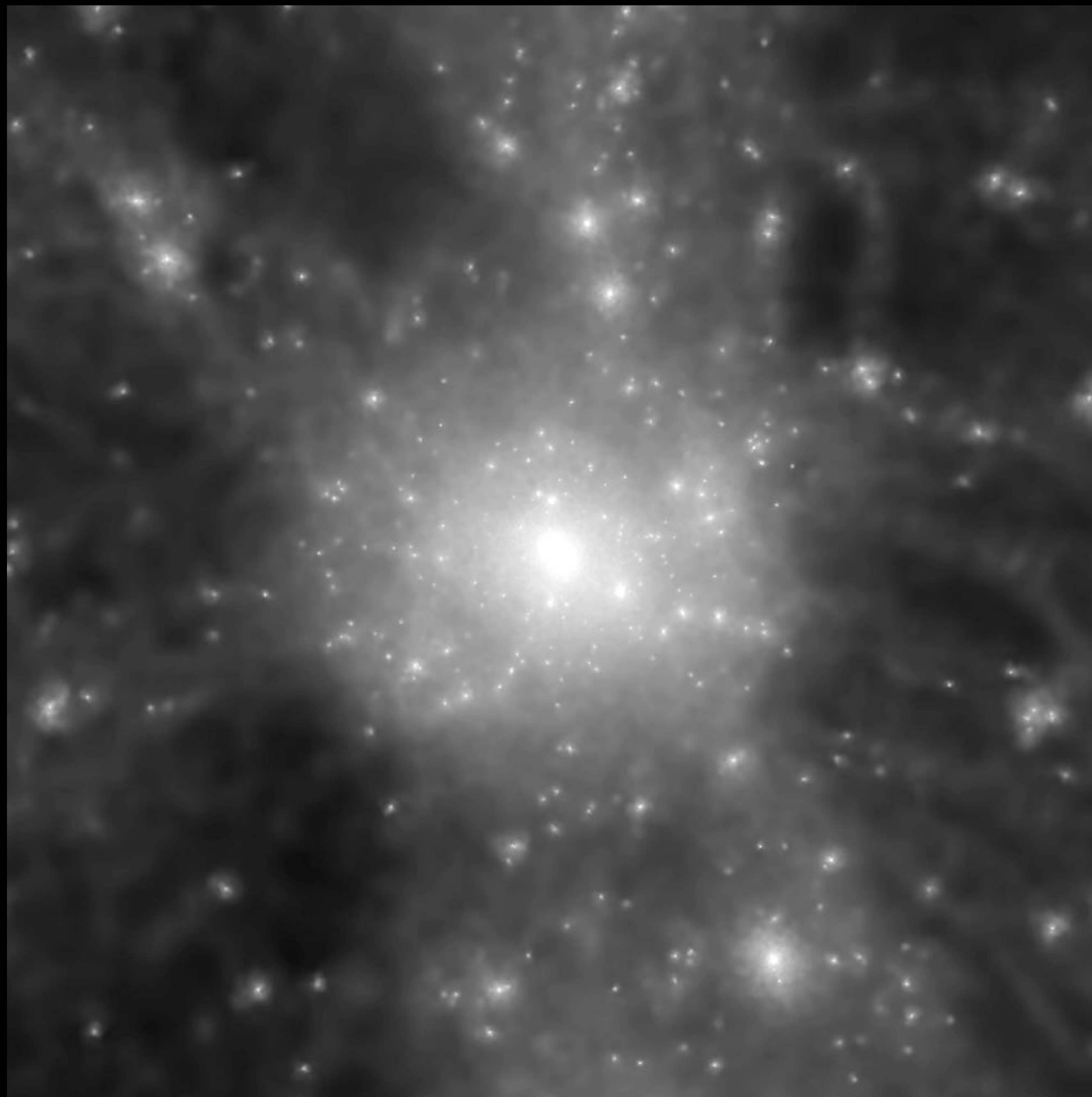




Part 2: “Genetically modified” simulations

Simulation Pros & Cons

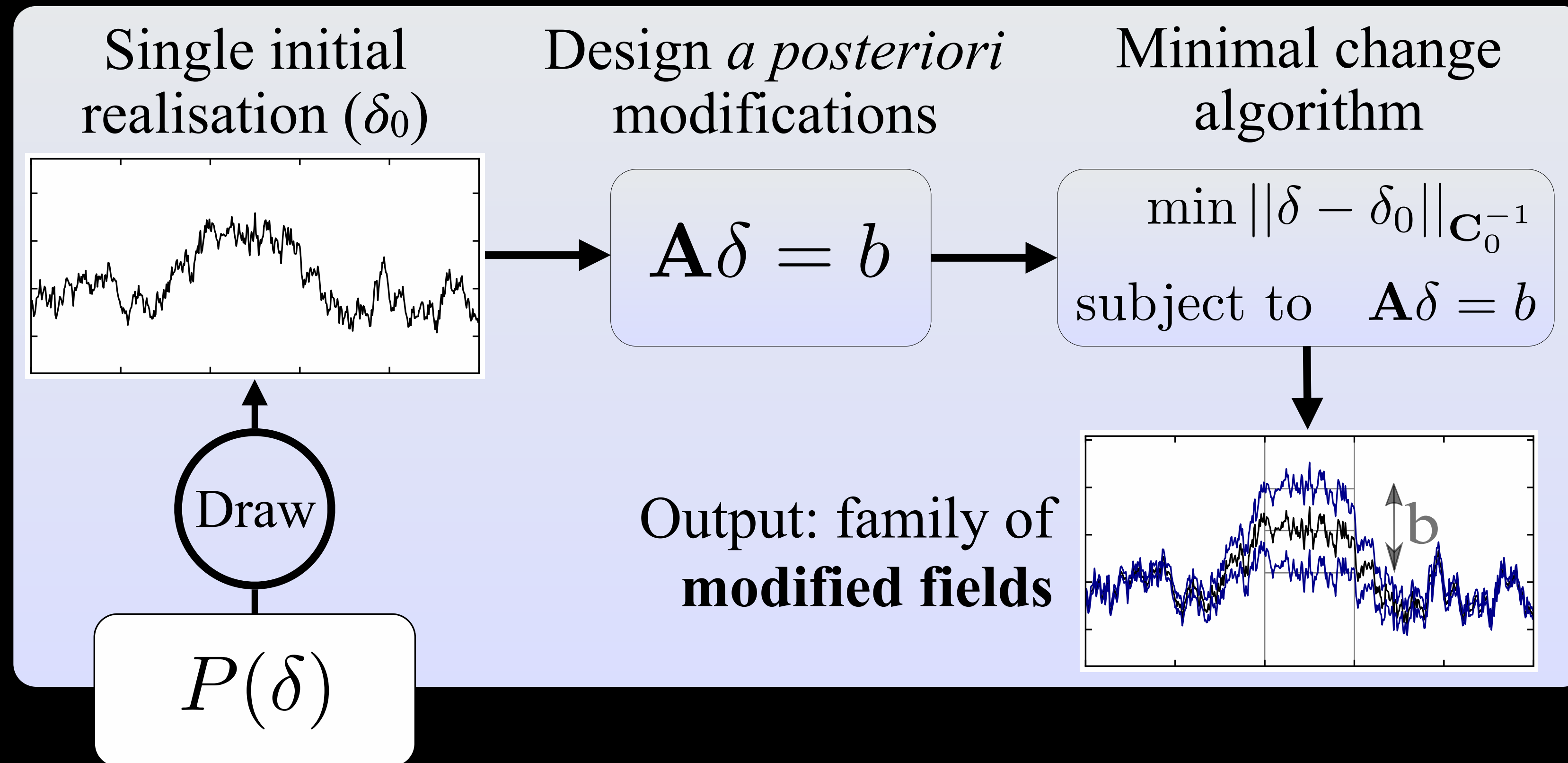
| | | Cosmology | | Resolution | | Control | |
|---|--|---|--|---|--|---|--|
|  Volume | |  | |  | |  | |
|  Zoom | |  | |  | |  | |
|  Idealised | |  | |  | |  | |
|  GM | |  | |  | |  | |



1 cMpc
(full box is 50 cMpc)

Formation
of a dwarf
galaxy halo

“Genetic modification” of early Universe

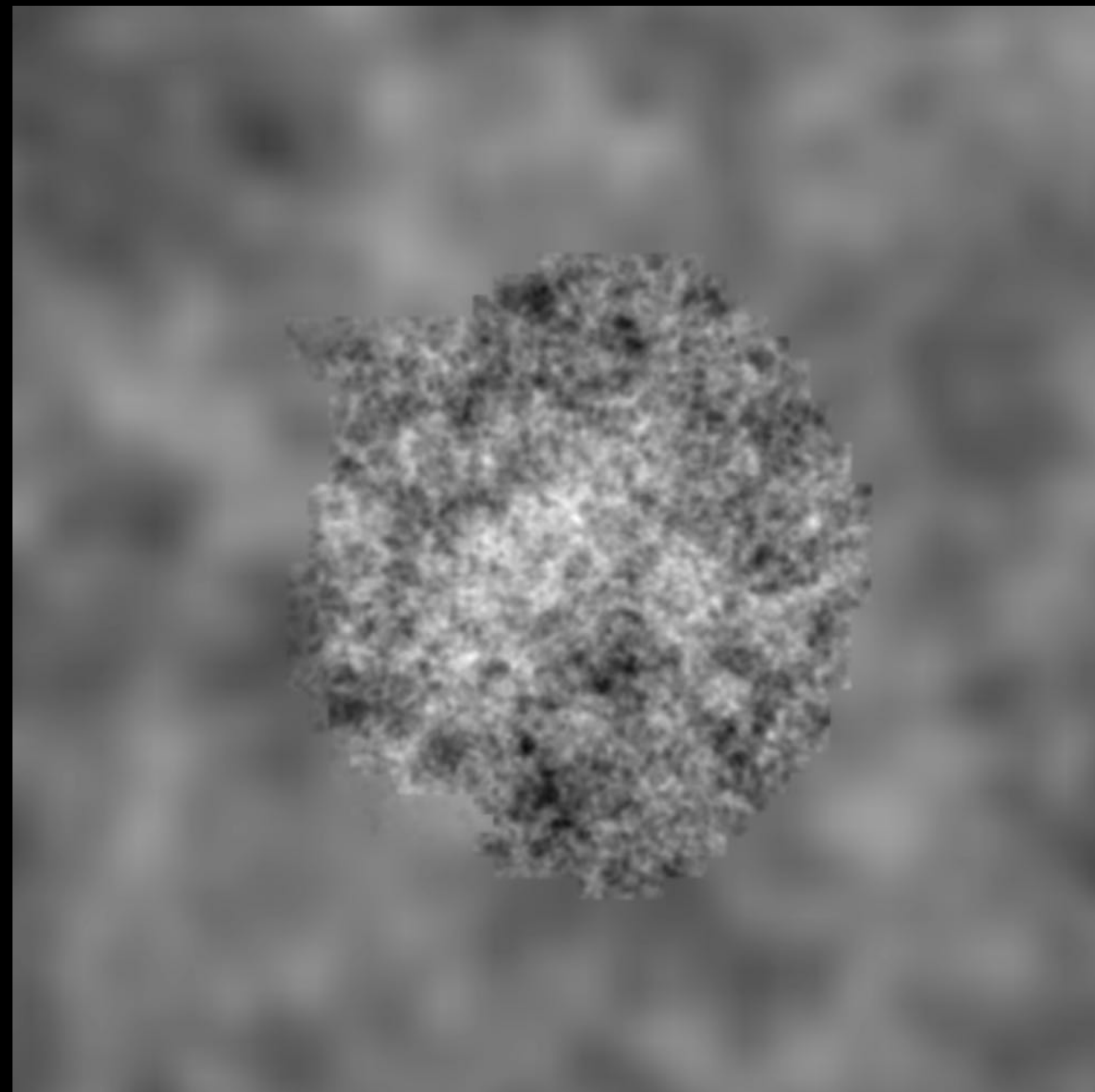


Rey+ 18



Modifications
are tiny
and smooth

Maintain consistency
with Λ CDM



Unmodified

$z=99$

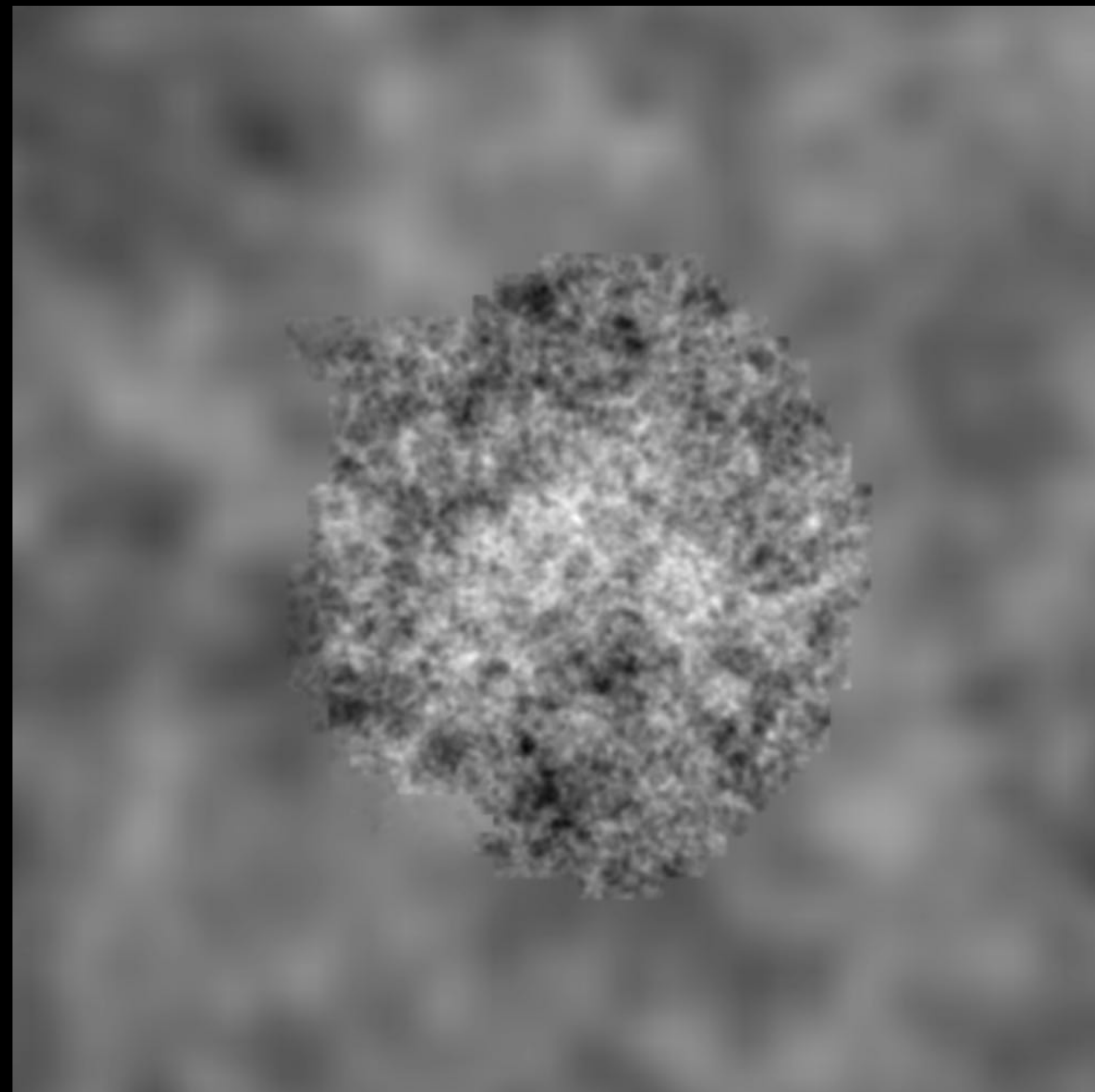
10 cMpc
(full box is 50 cMpc)

Roth+ 2016



Modifications
are tiny
and smooth

Maintain consistency
with Λ CDM



GM late formation

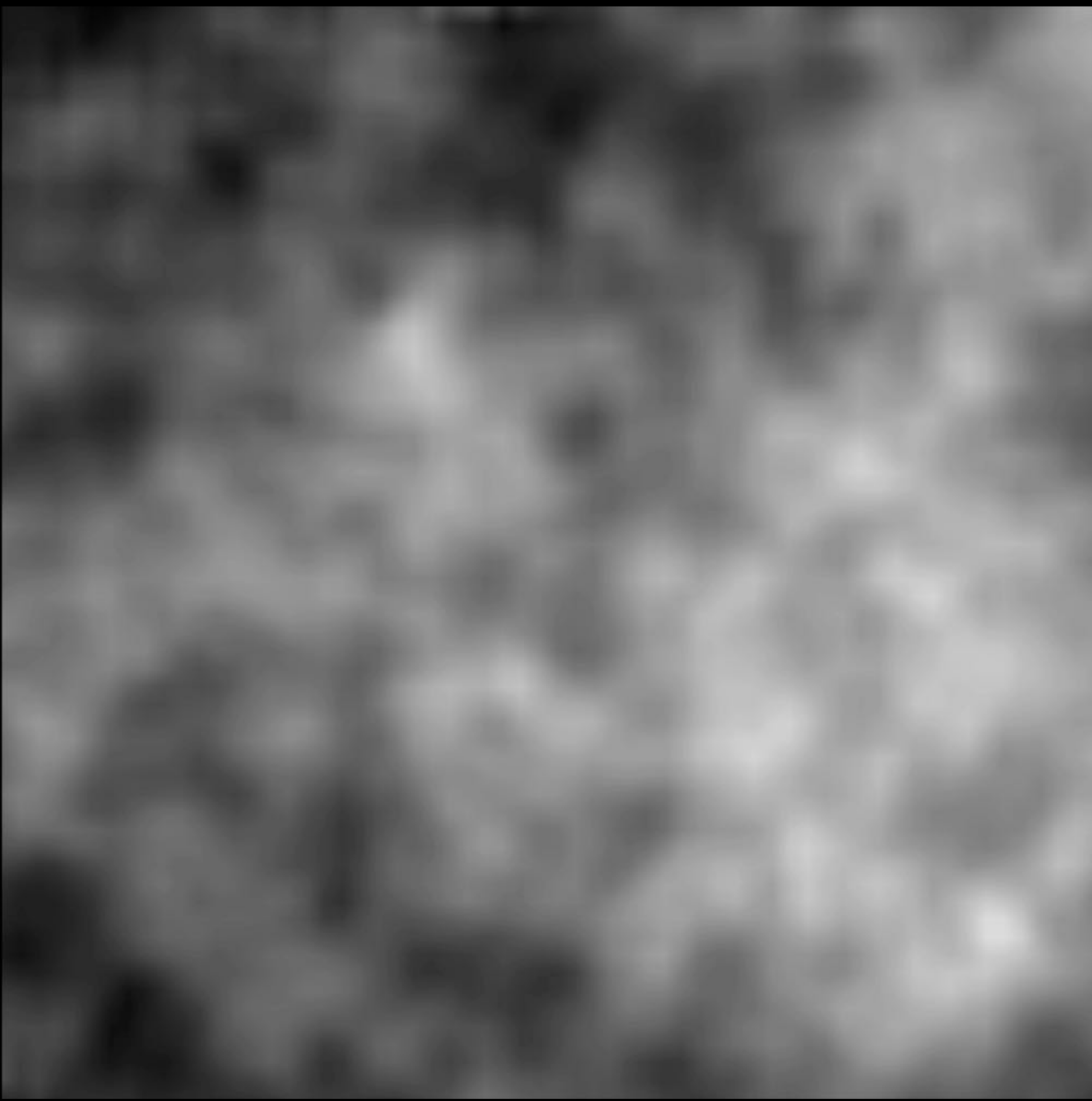
$z=99$

10 cMpc
(full box is 50 cMpc)

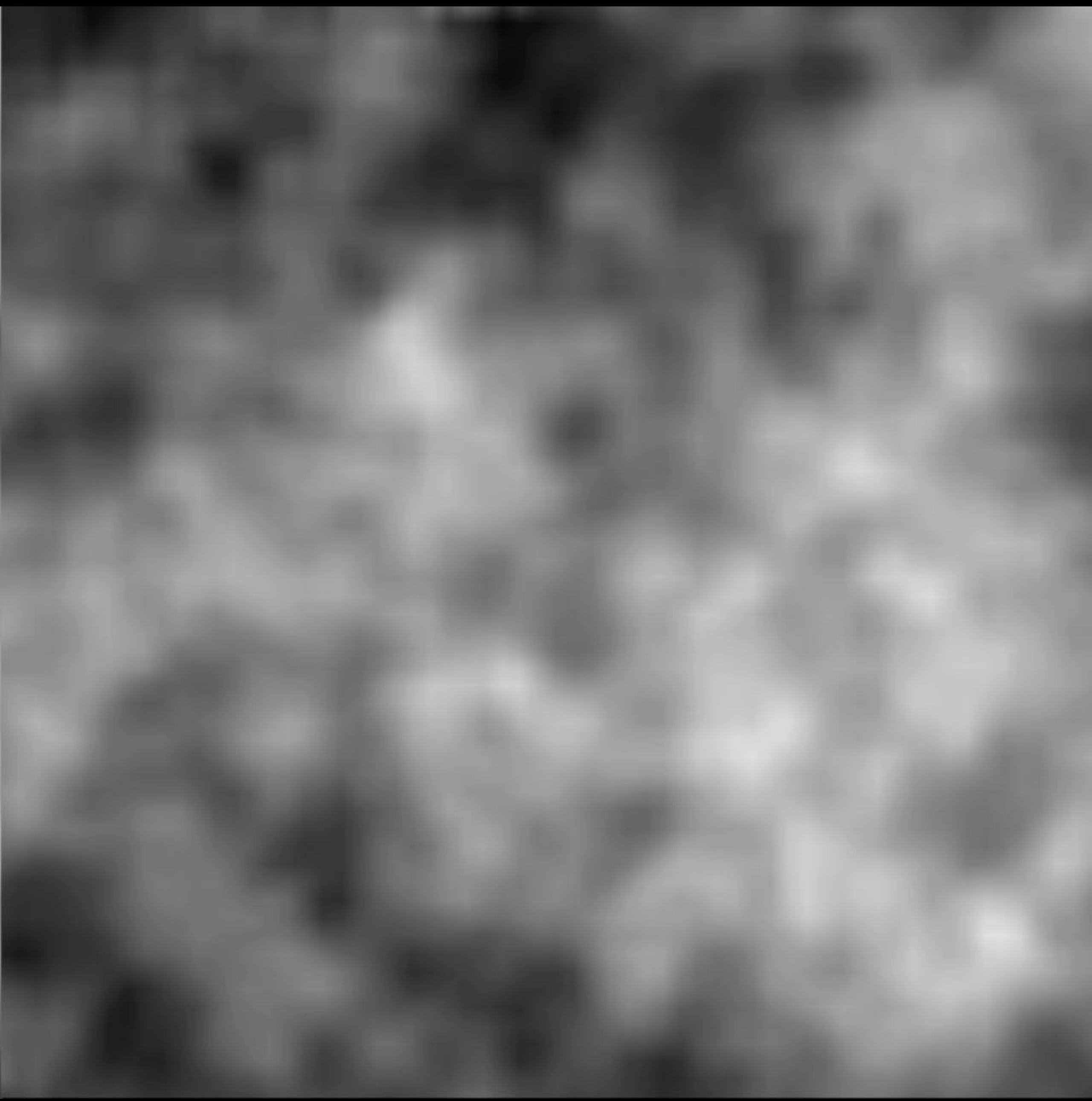
Roth+ 2016



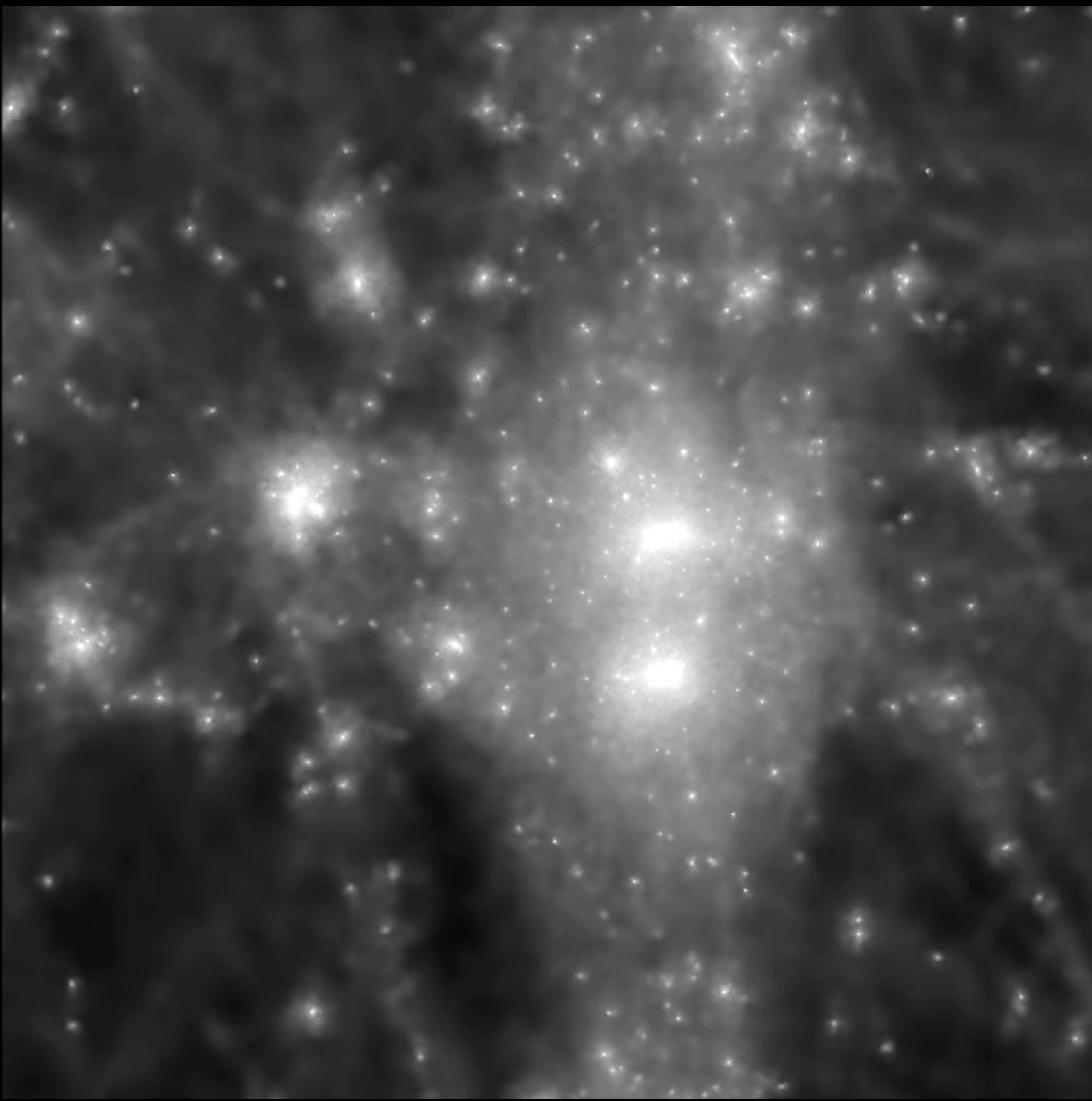
Unmodified



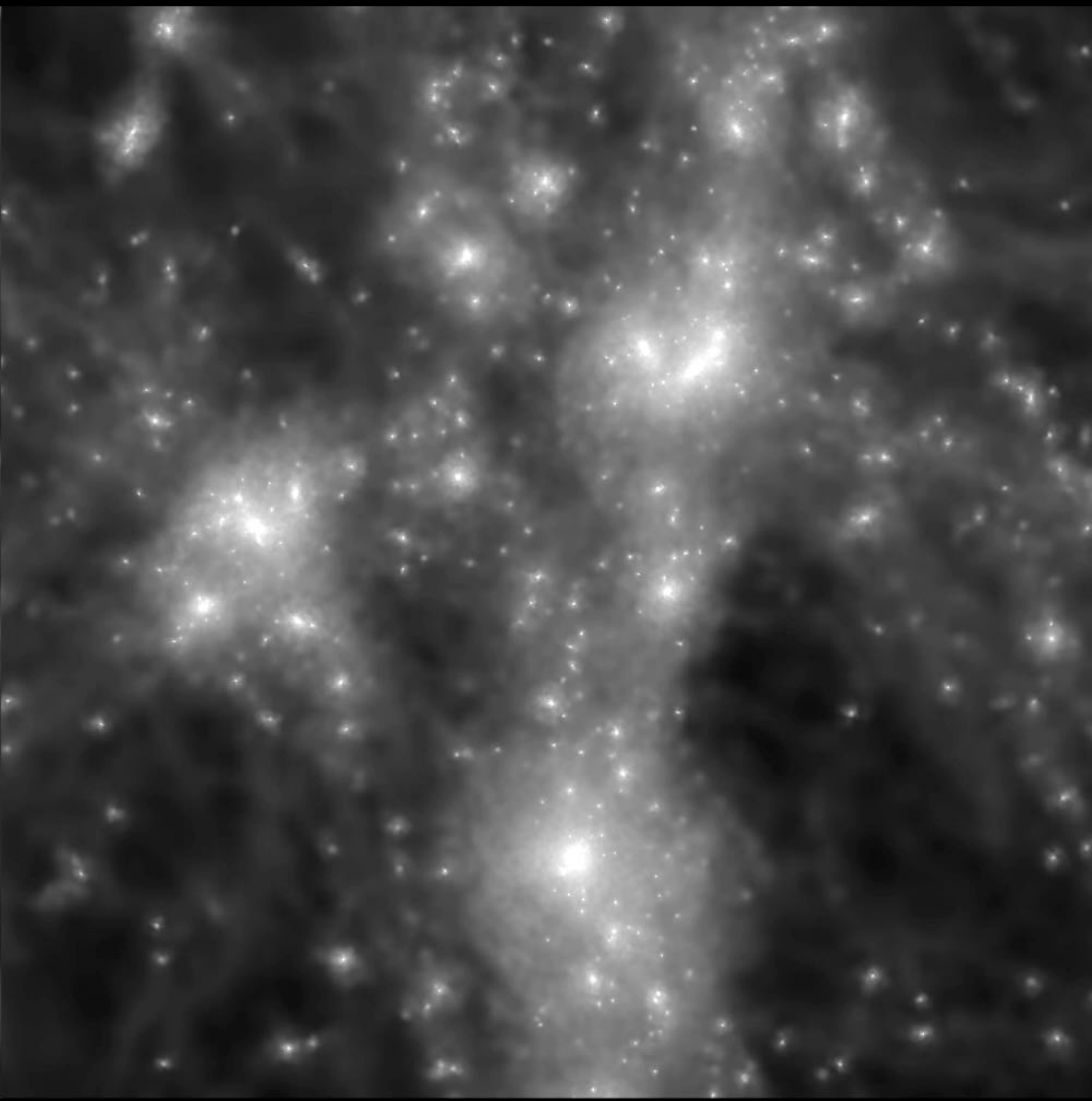
Late collapse



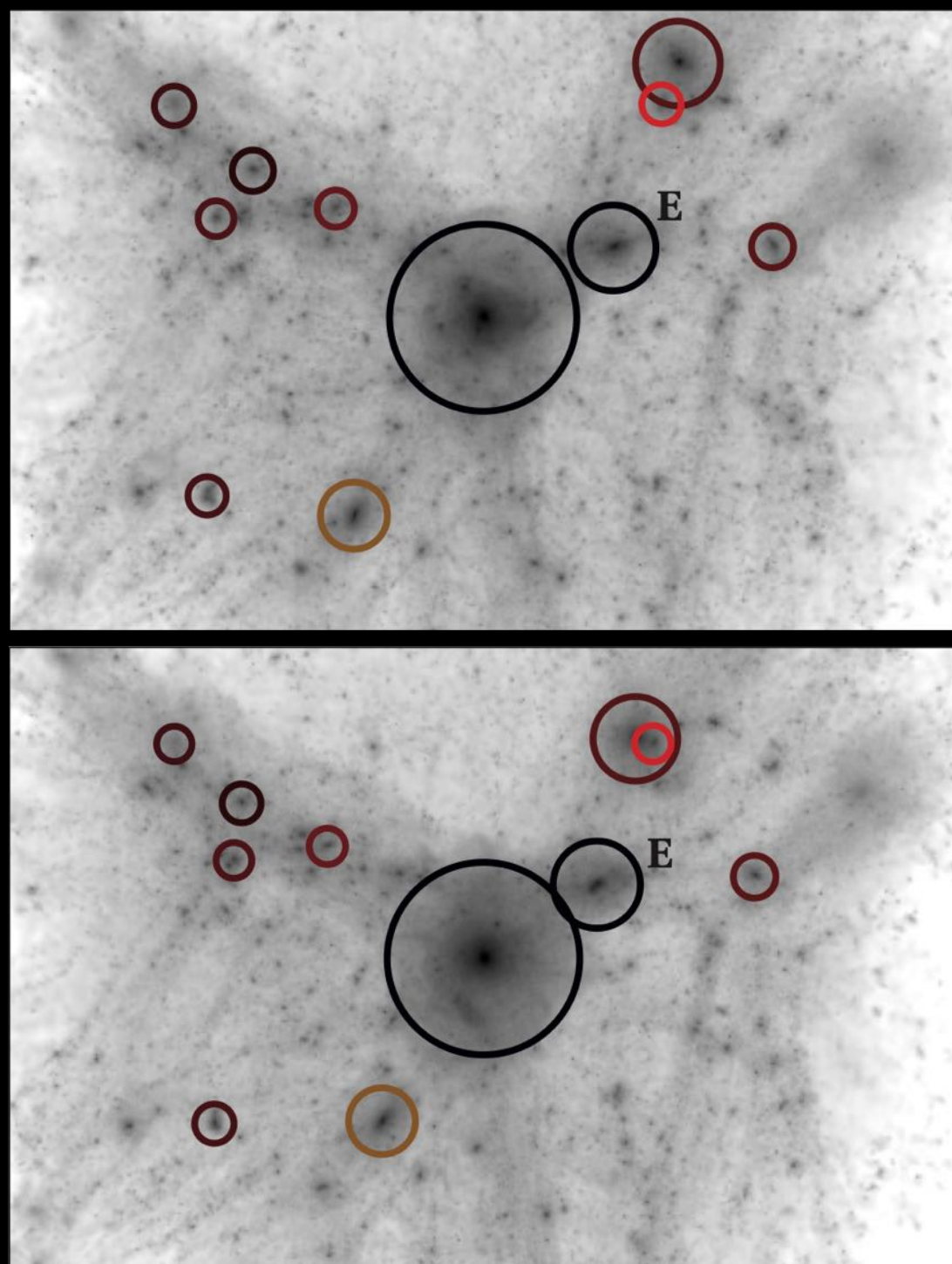
Unmodified



Late collapse

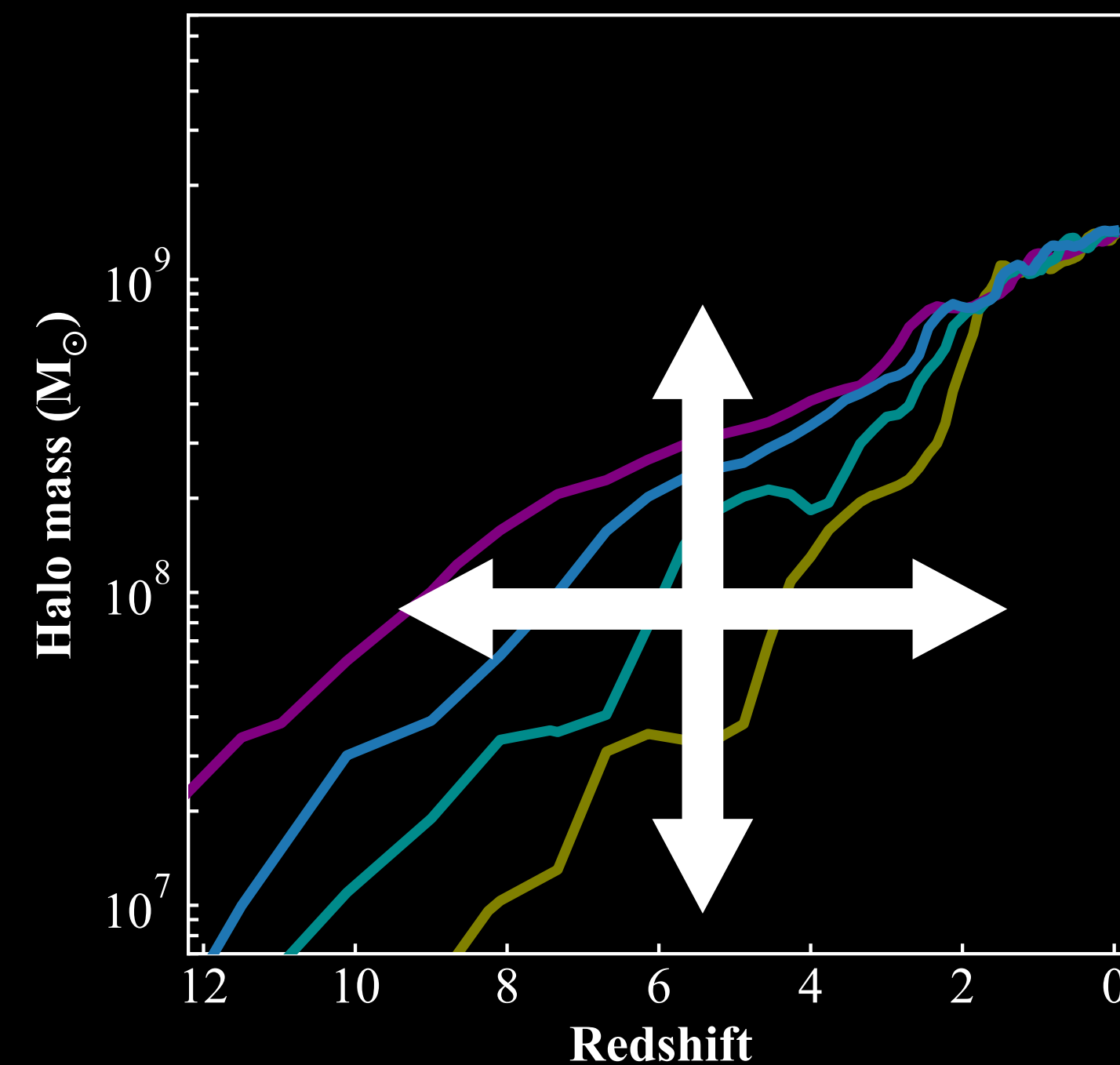


Why is this good?



Consistent
environment

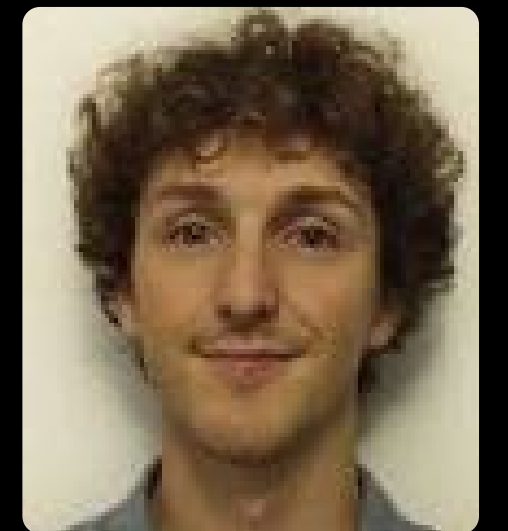
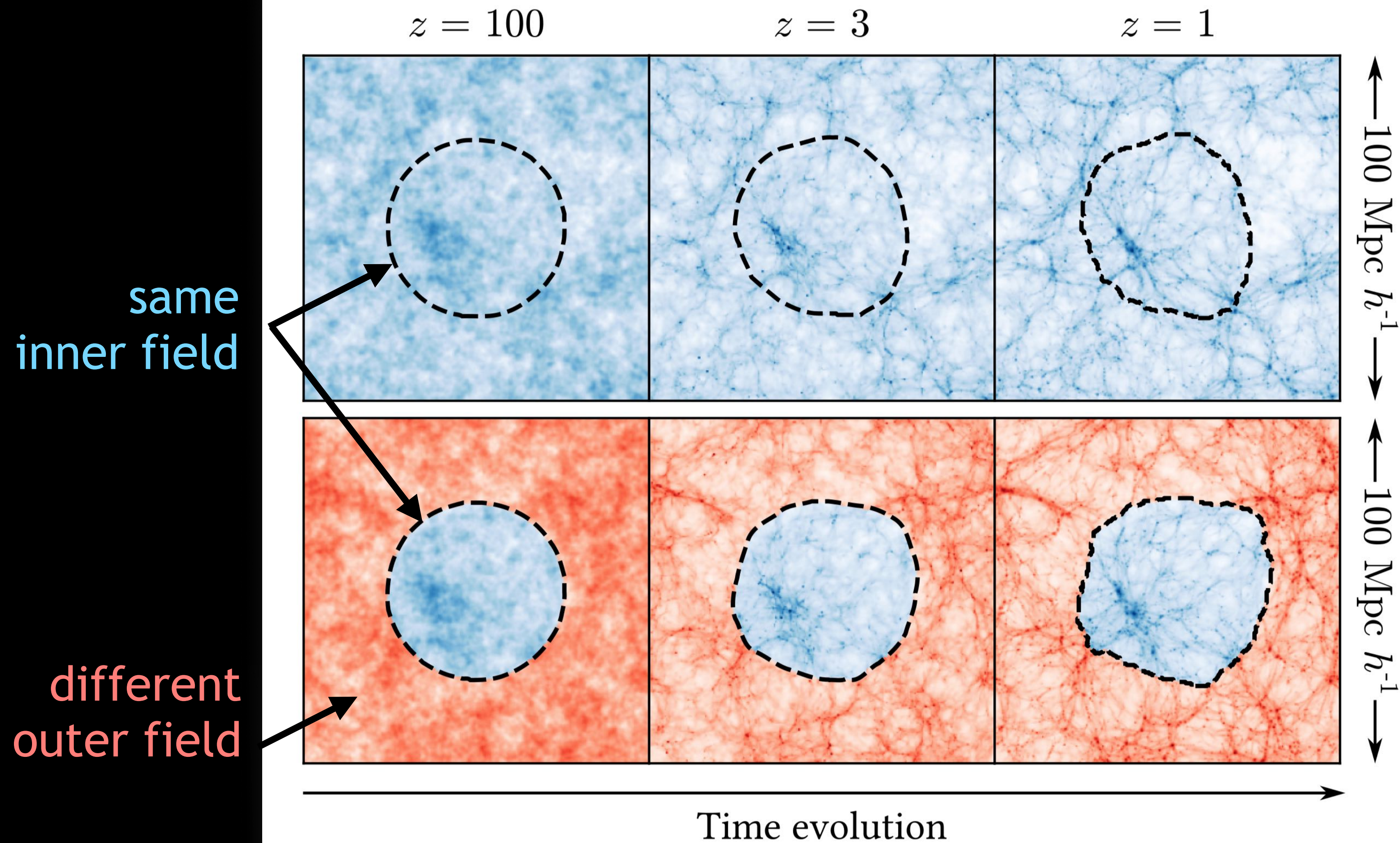
no confusion from
random sampling



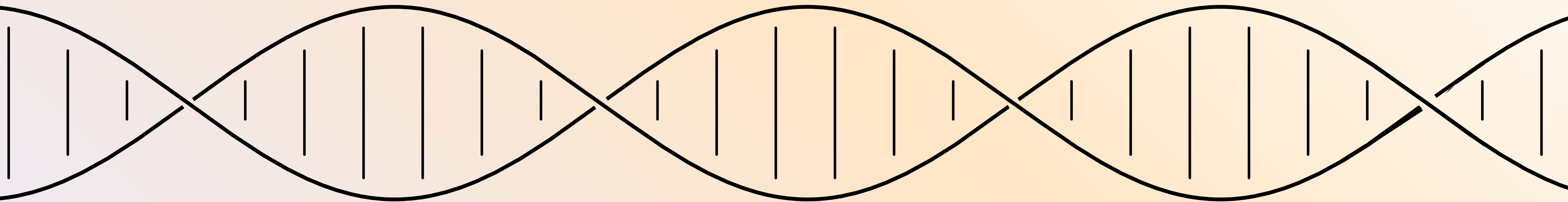
Infinite
tunability

pose and test
hypotheses

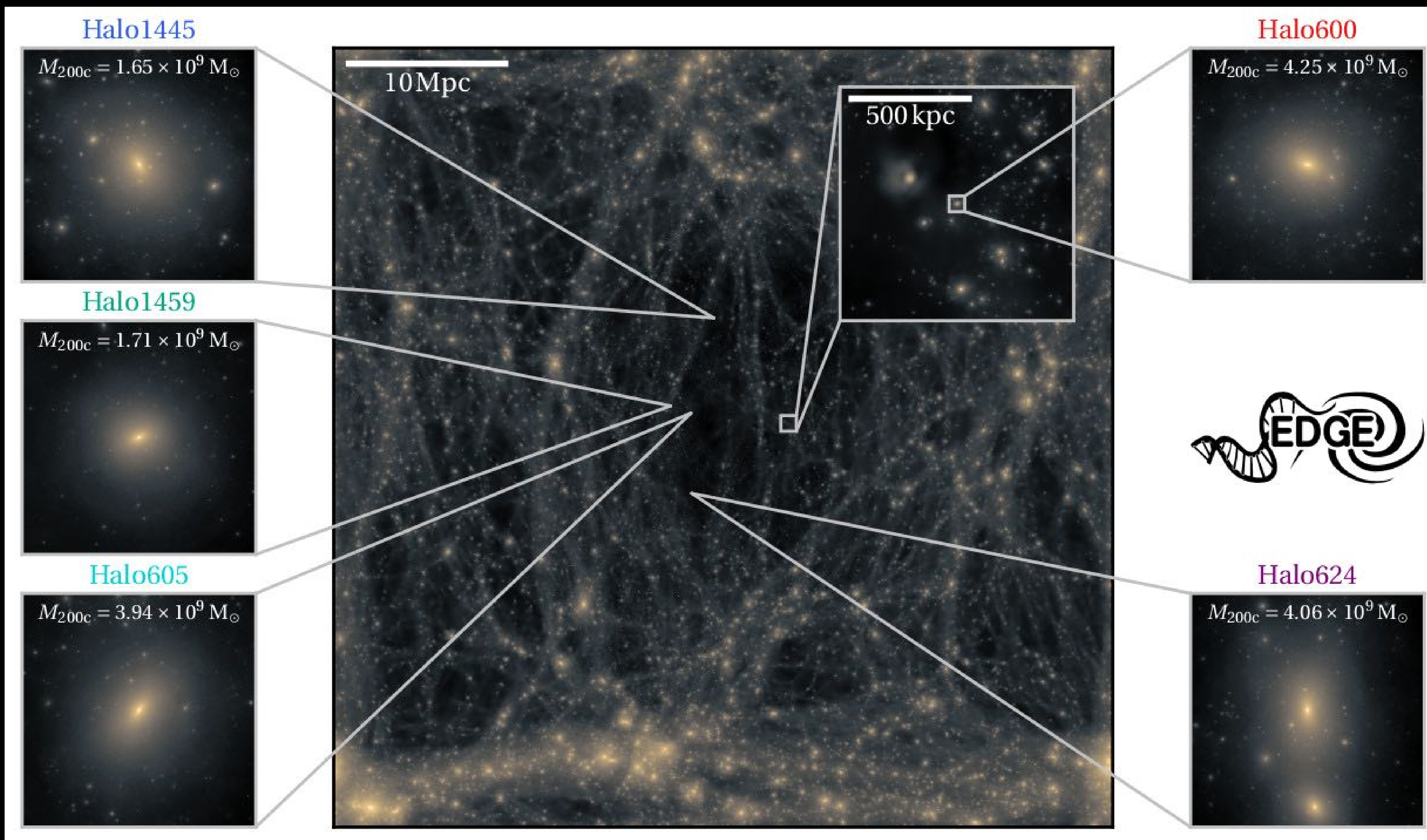
Also: “Gene Splicing”



Cadiou+ 21



Part 3: Results / predictions

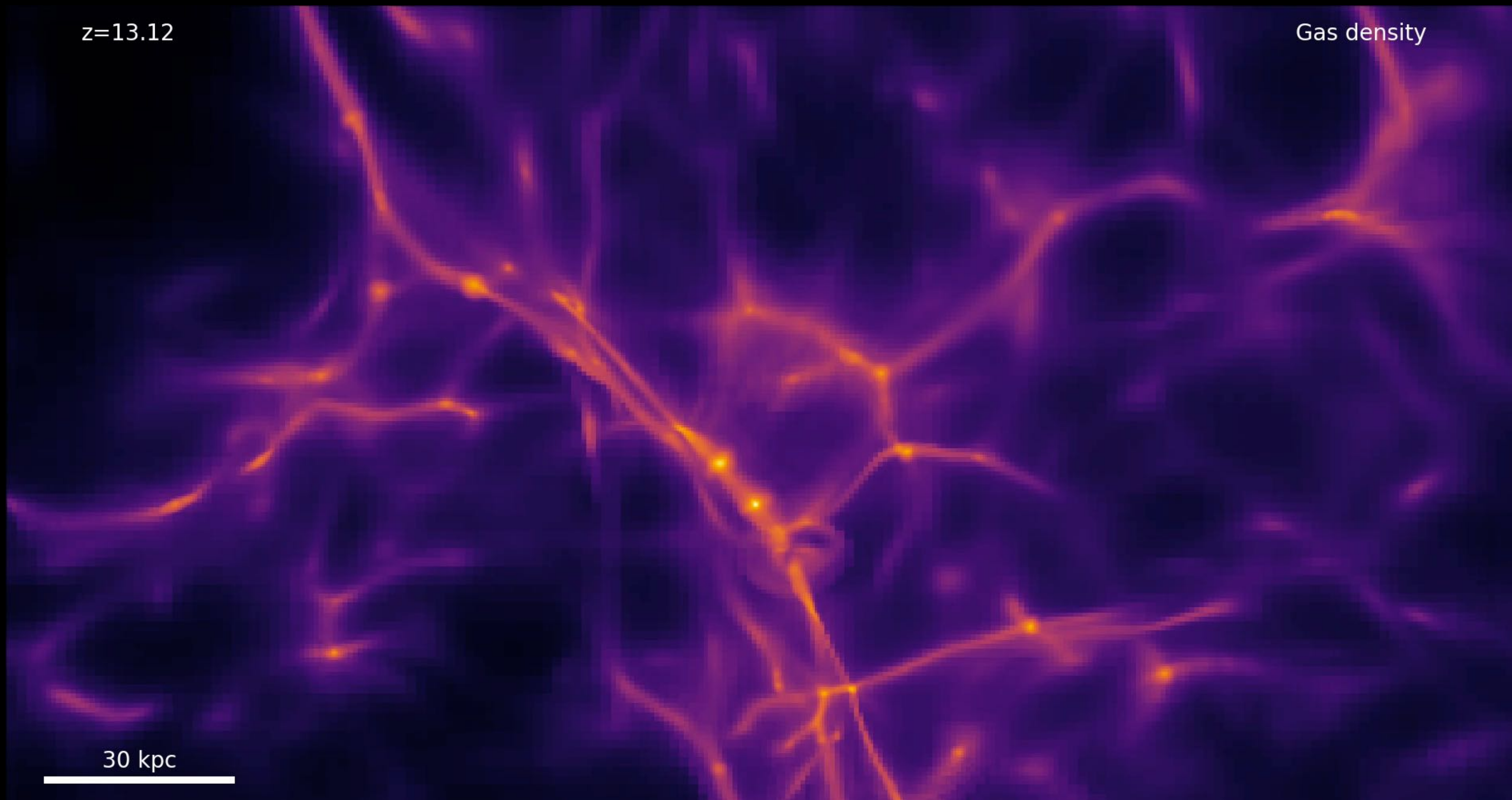


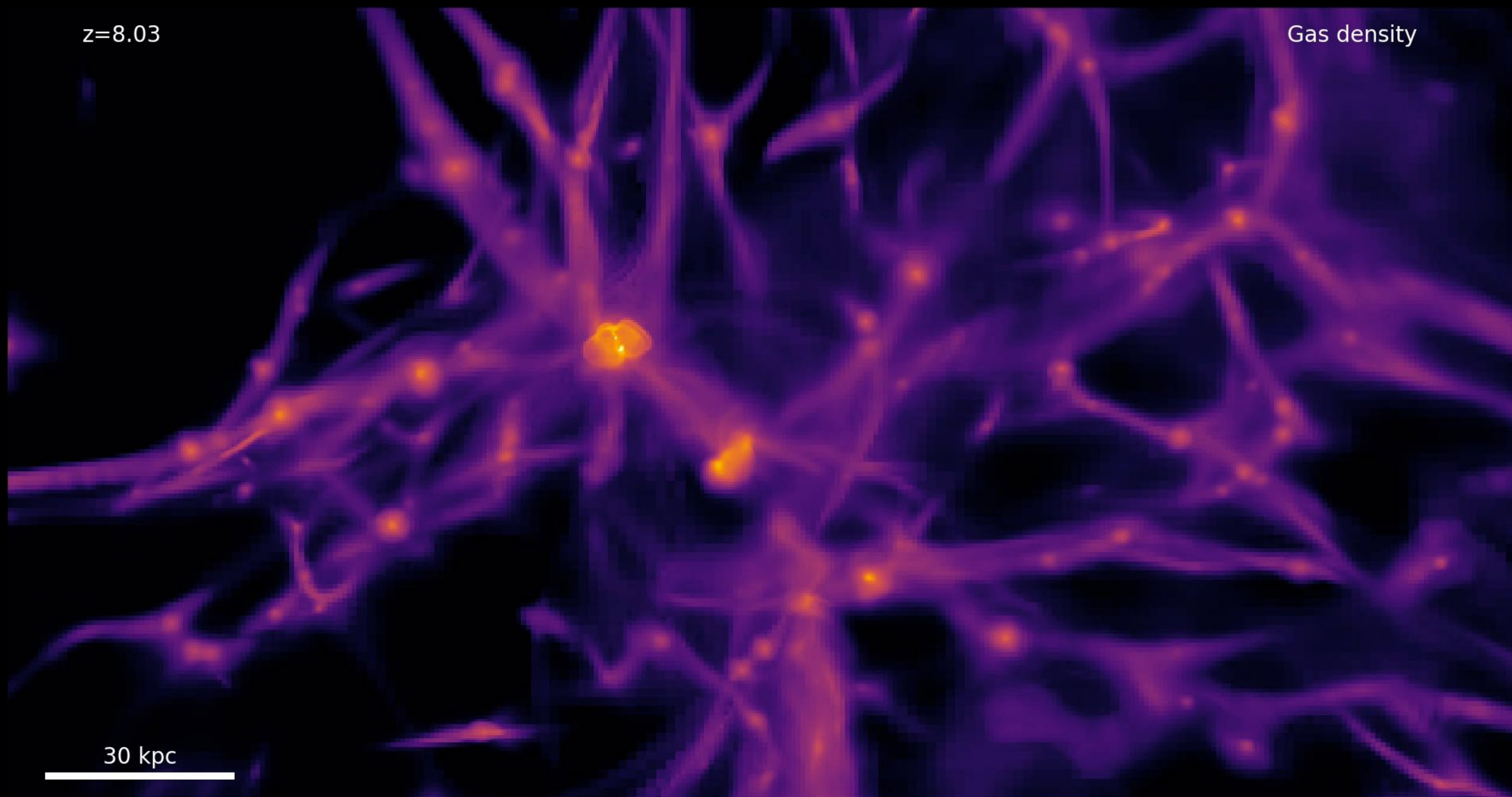


$z=13.12$

Gas density

30 kpc



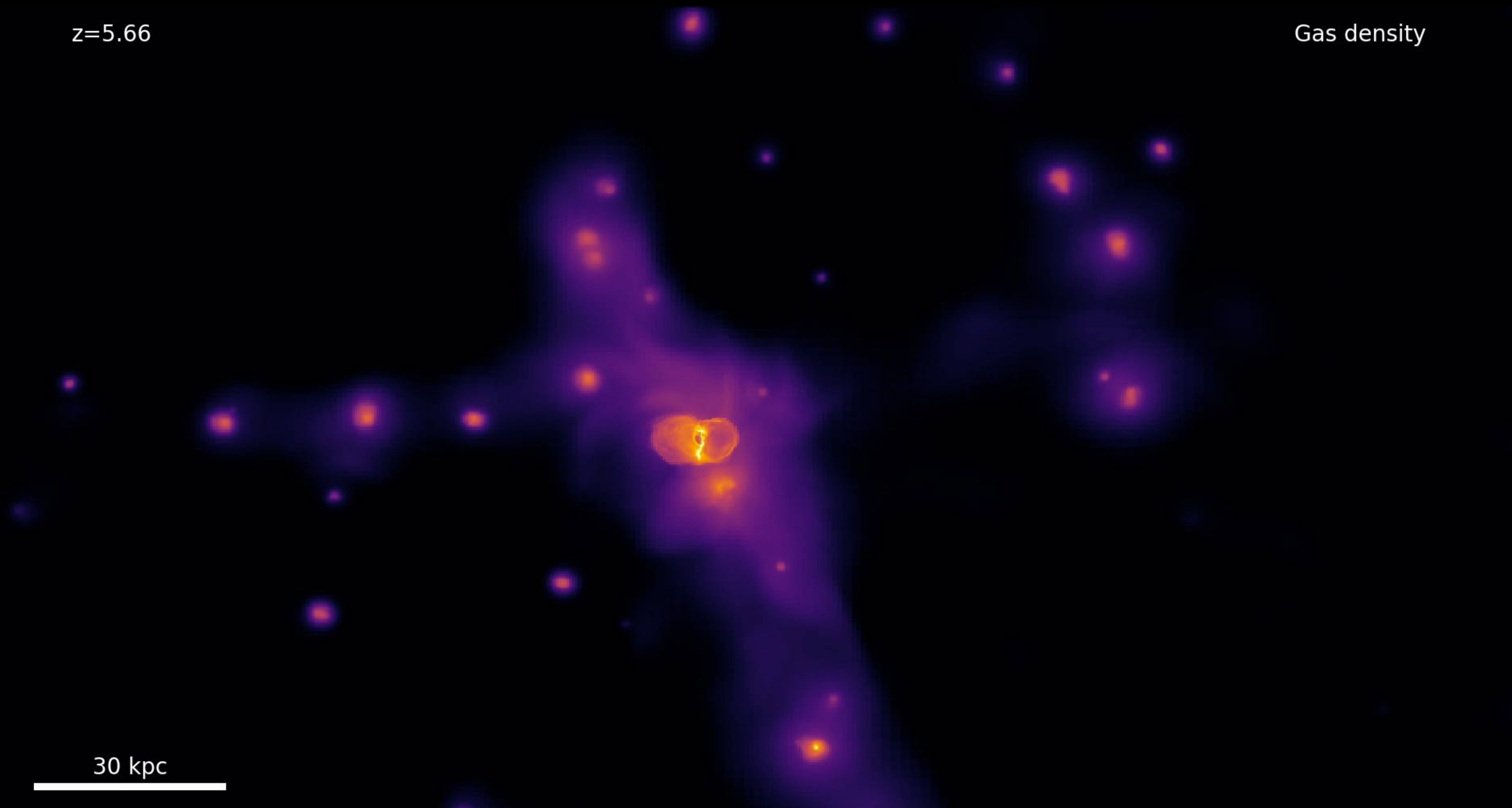




$z=5.66$

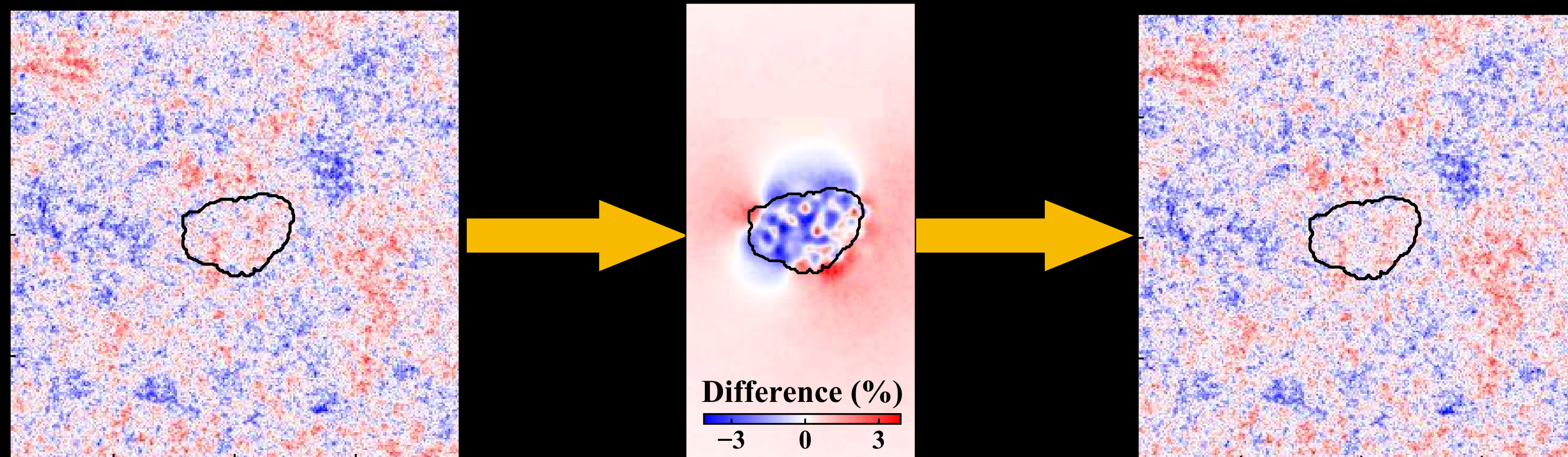
Gas density

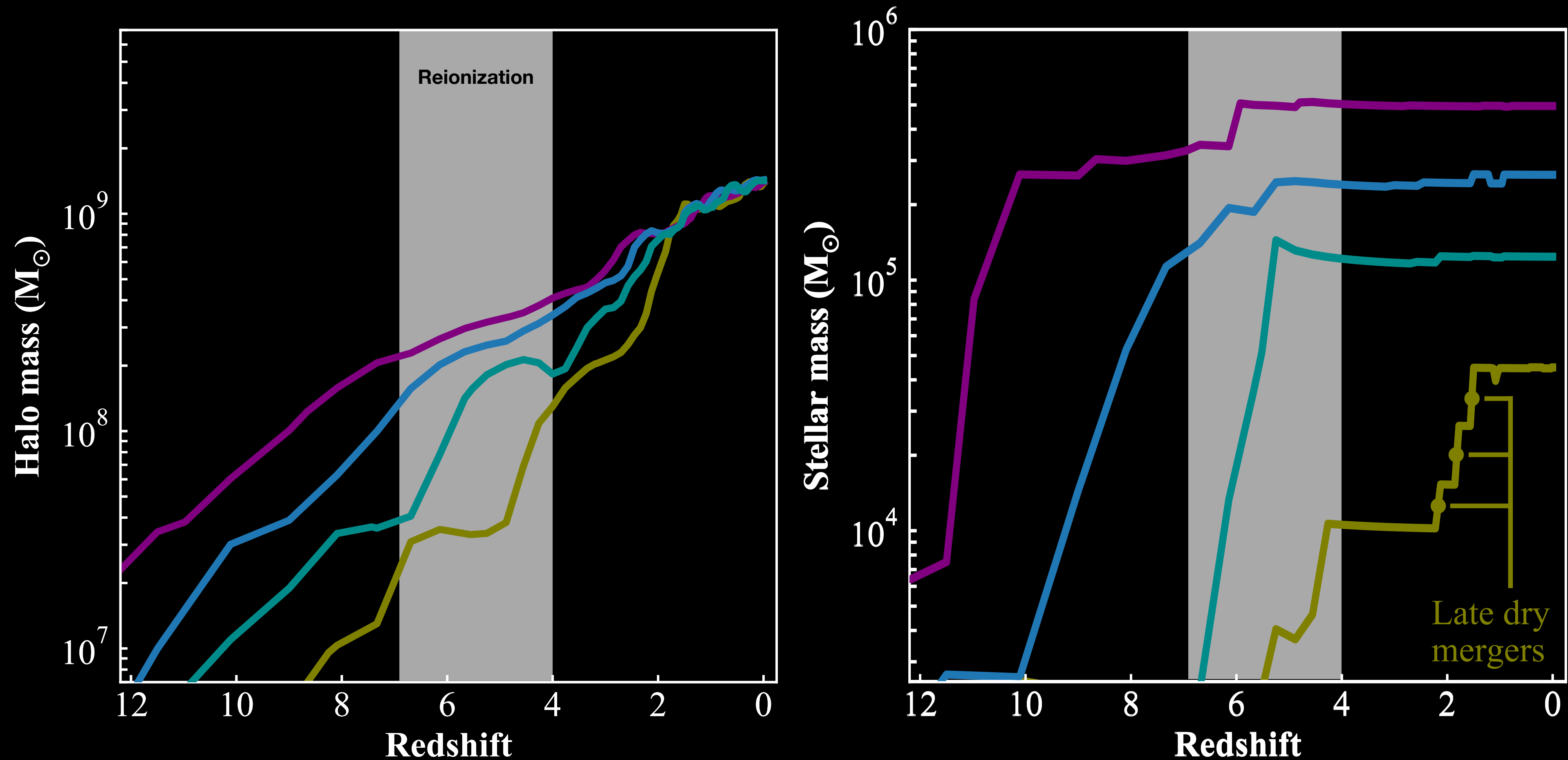
30 kpc





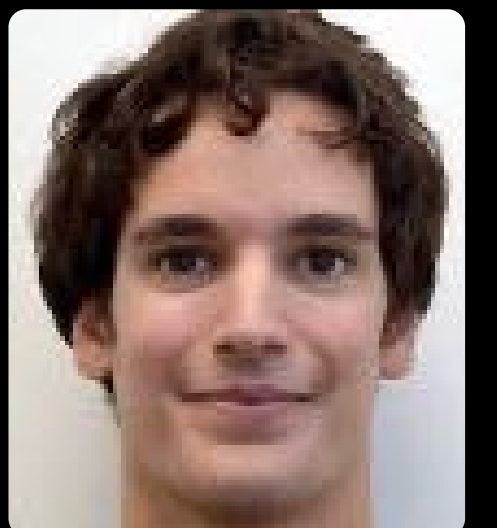
“Genetic modification” of early Universe



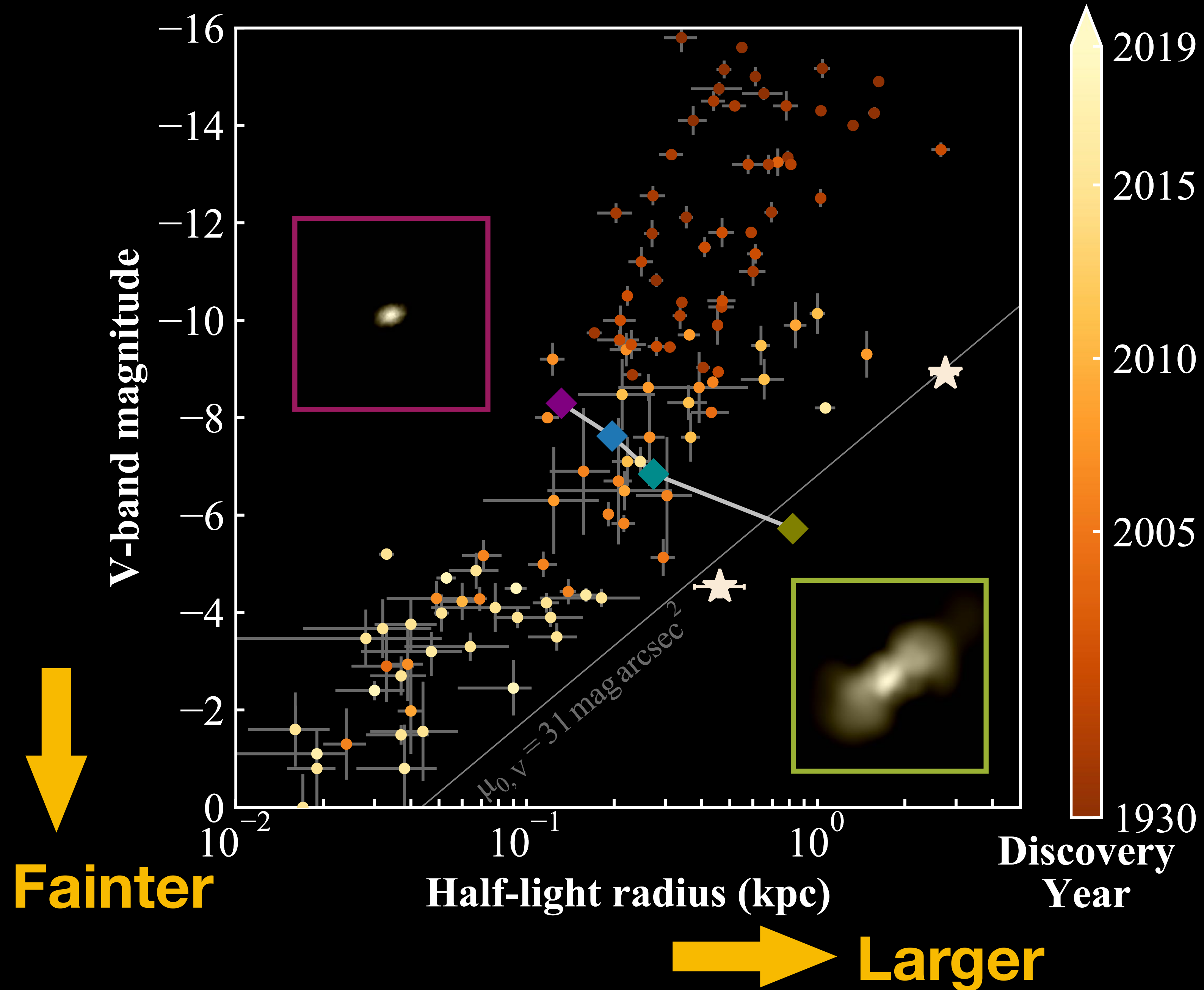


Earlier forming ultra-faint have higher stellar mass, at fixed halo mass today.

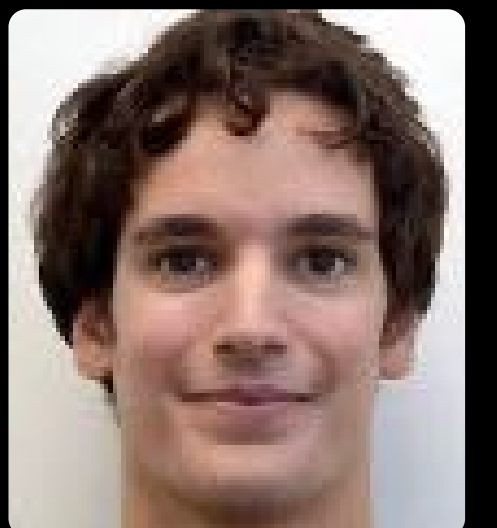
Rey+ 19



Finding dwarfs



Rey+ 19



Redshift 7.8
0.67 Gyr
Step 181

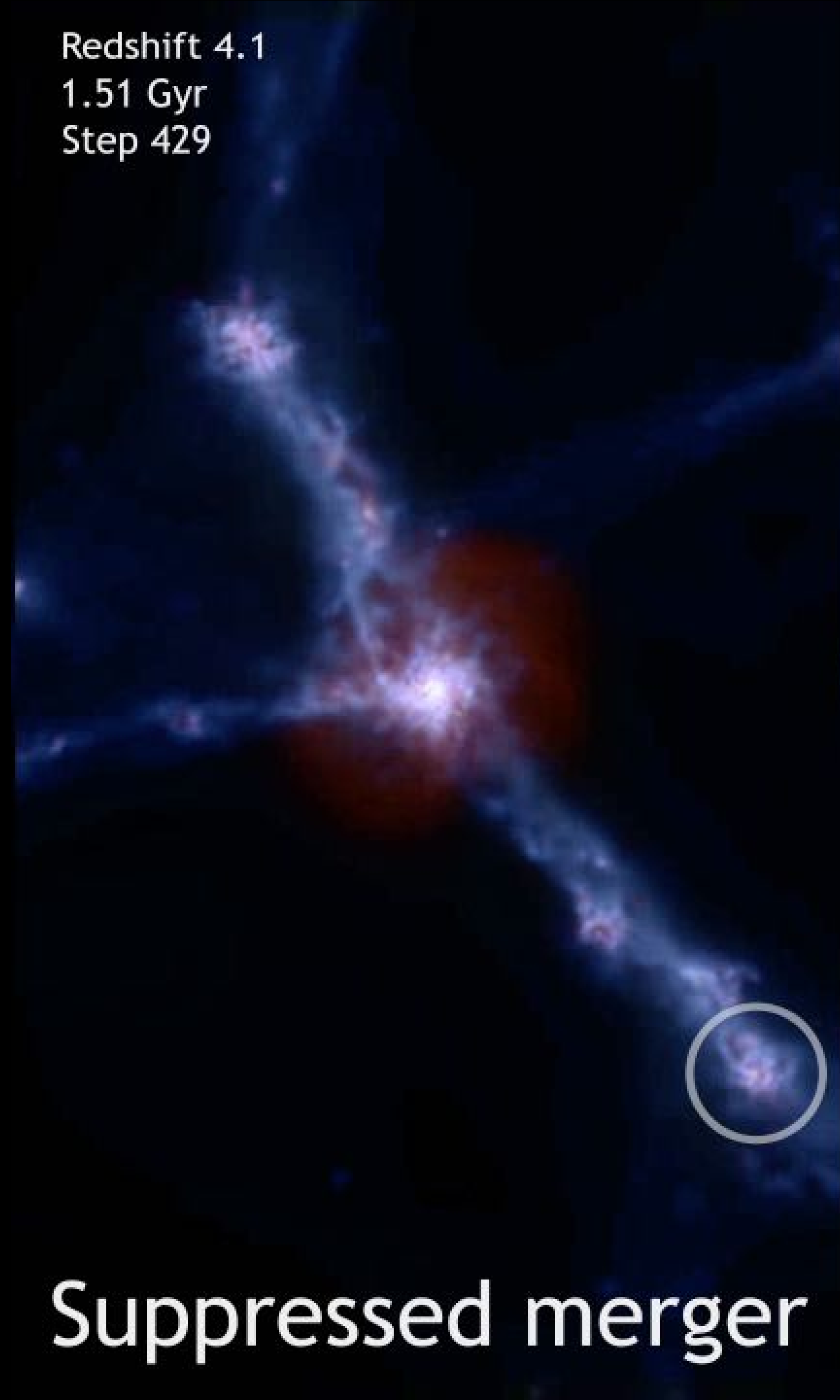
Suppressed merger

Reference

Enhanced merger

Pontzen+ 1607.02507; Tremmel+ 1607.02151

Redshift 4.1
1.51 Gyr
Step 429



Suppressed merger



Reference

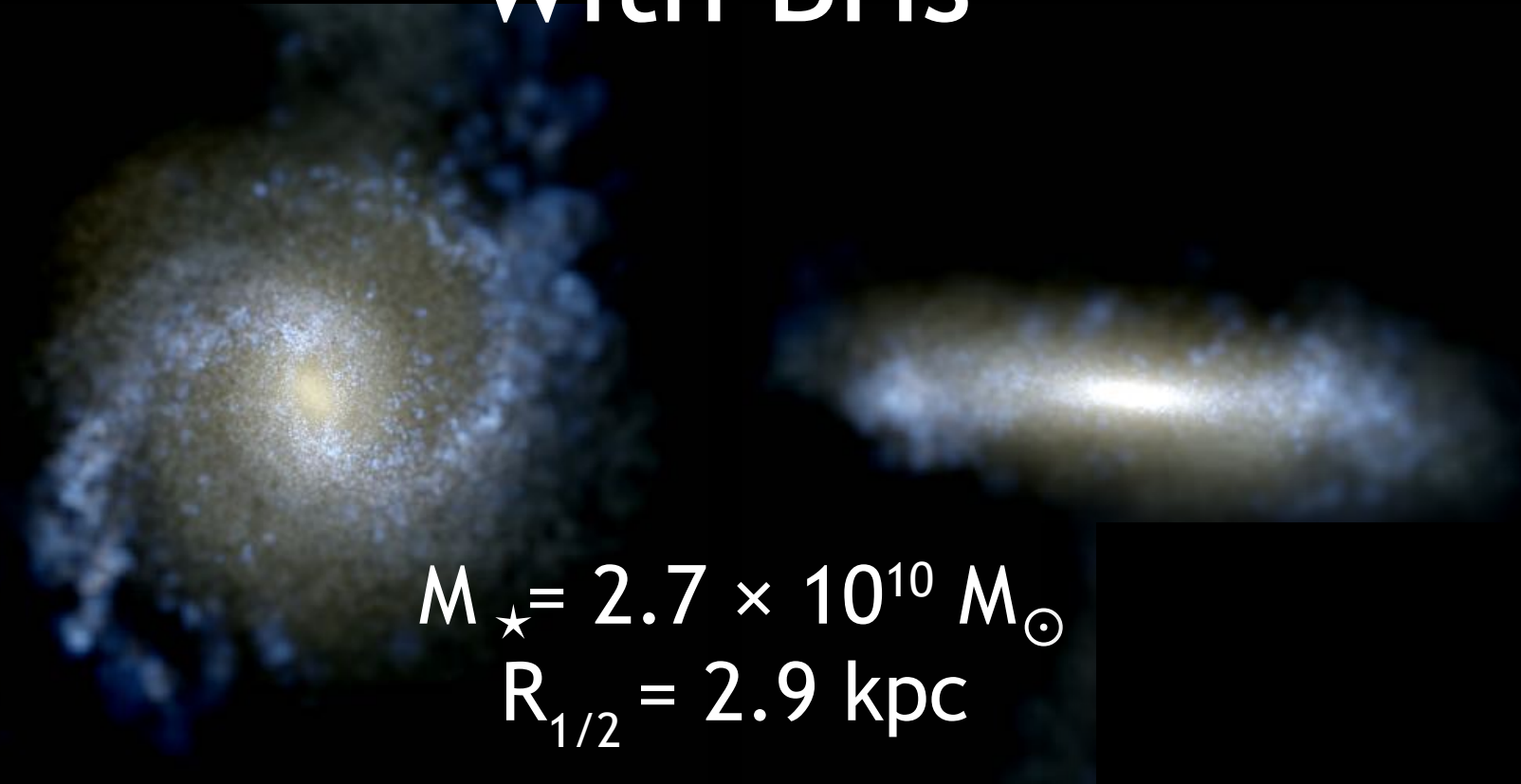


Enhanced merger

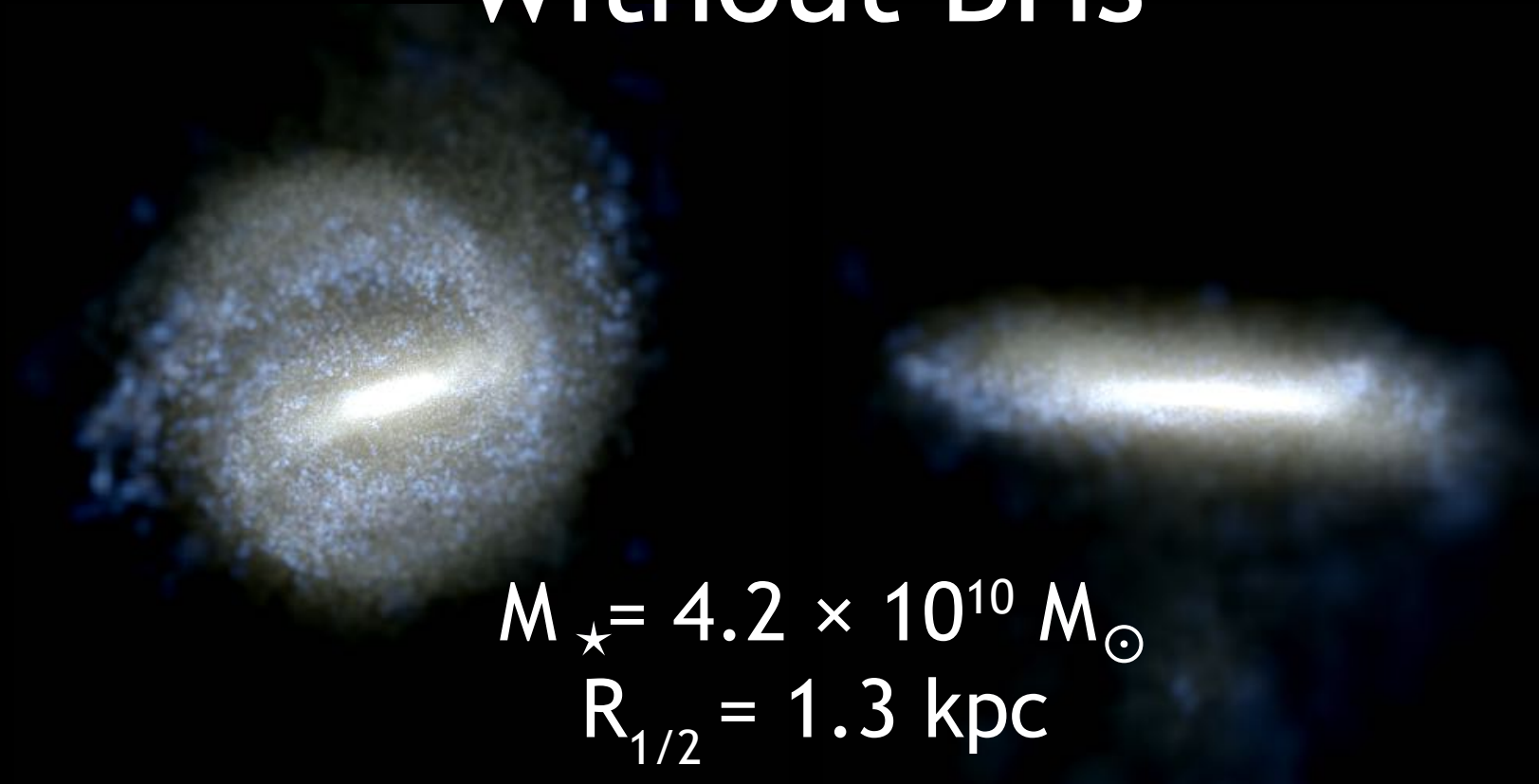
Pontzen+ 1607.02507; Tremmel+ 1607.02151

Suppressed

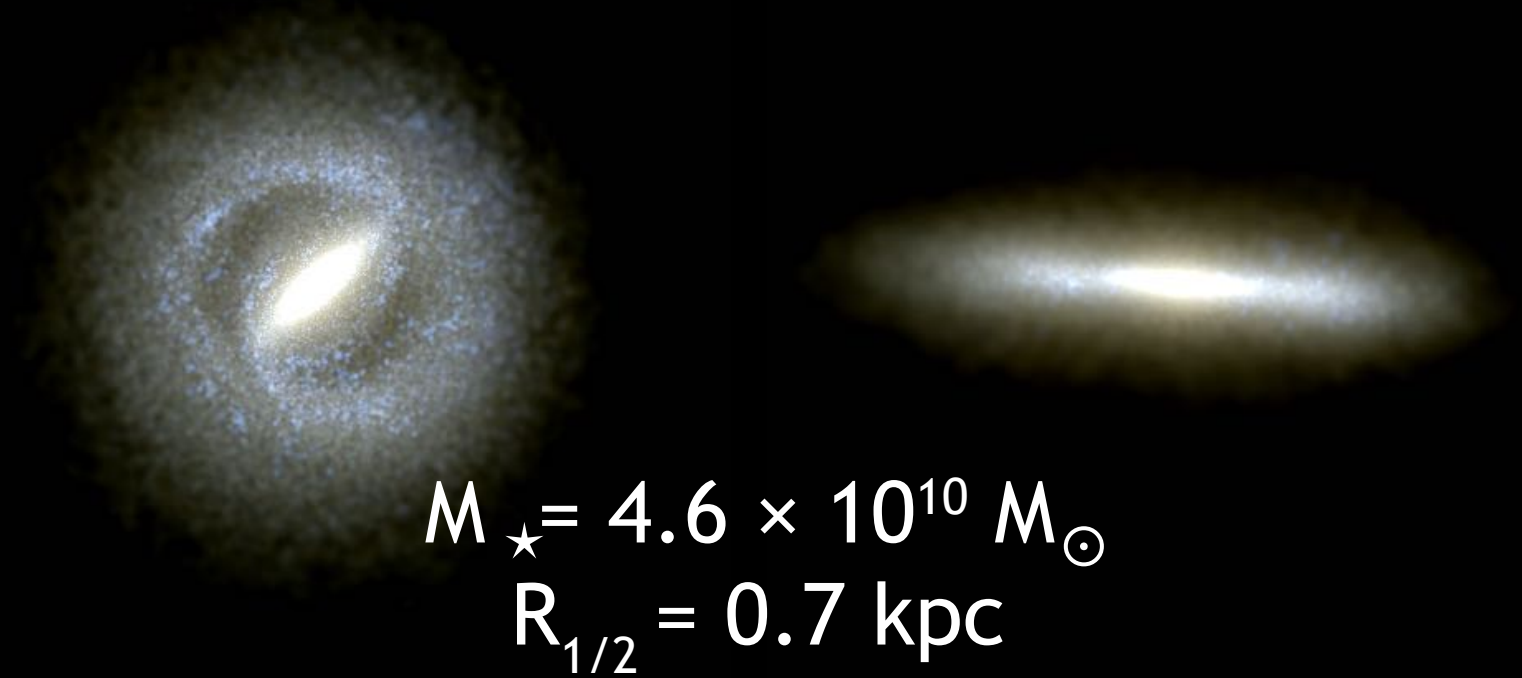
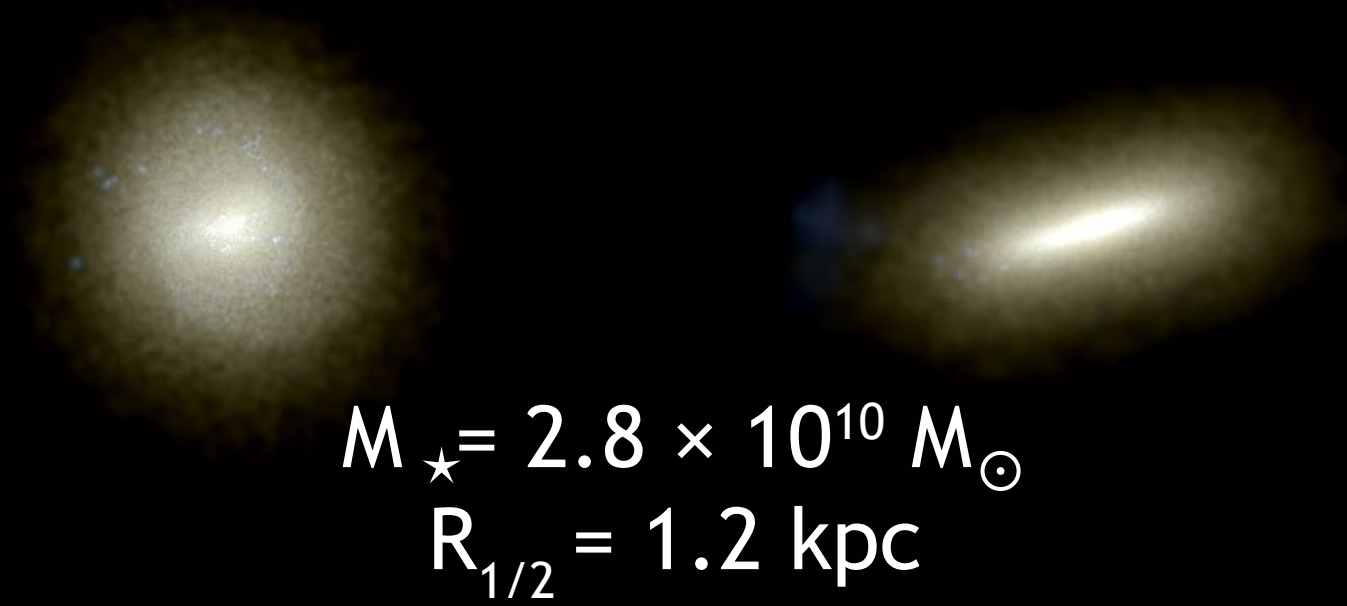
with BHs



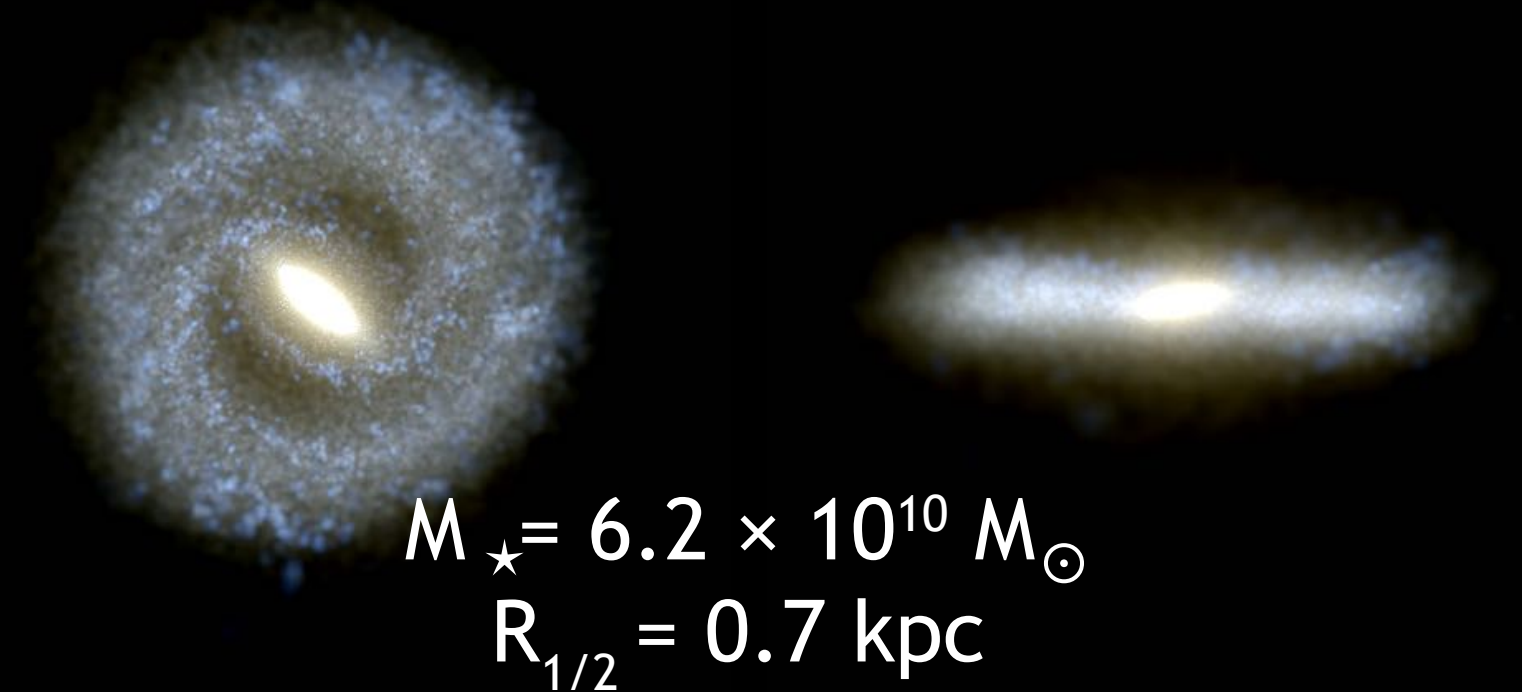
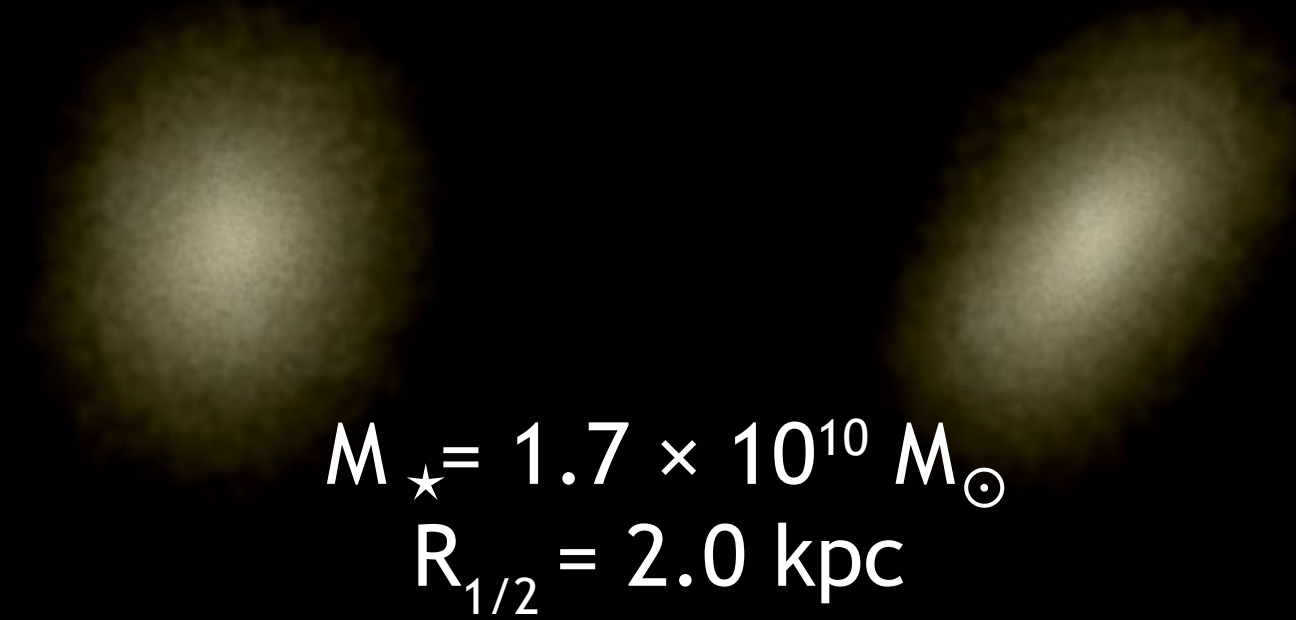
without BHs



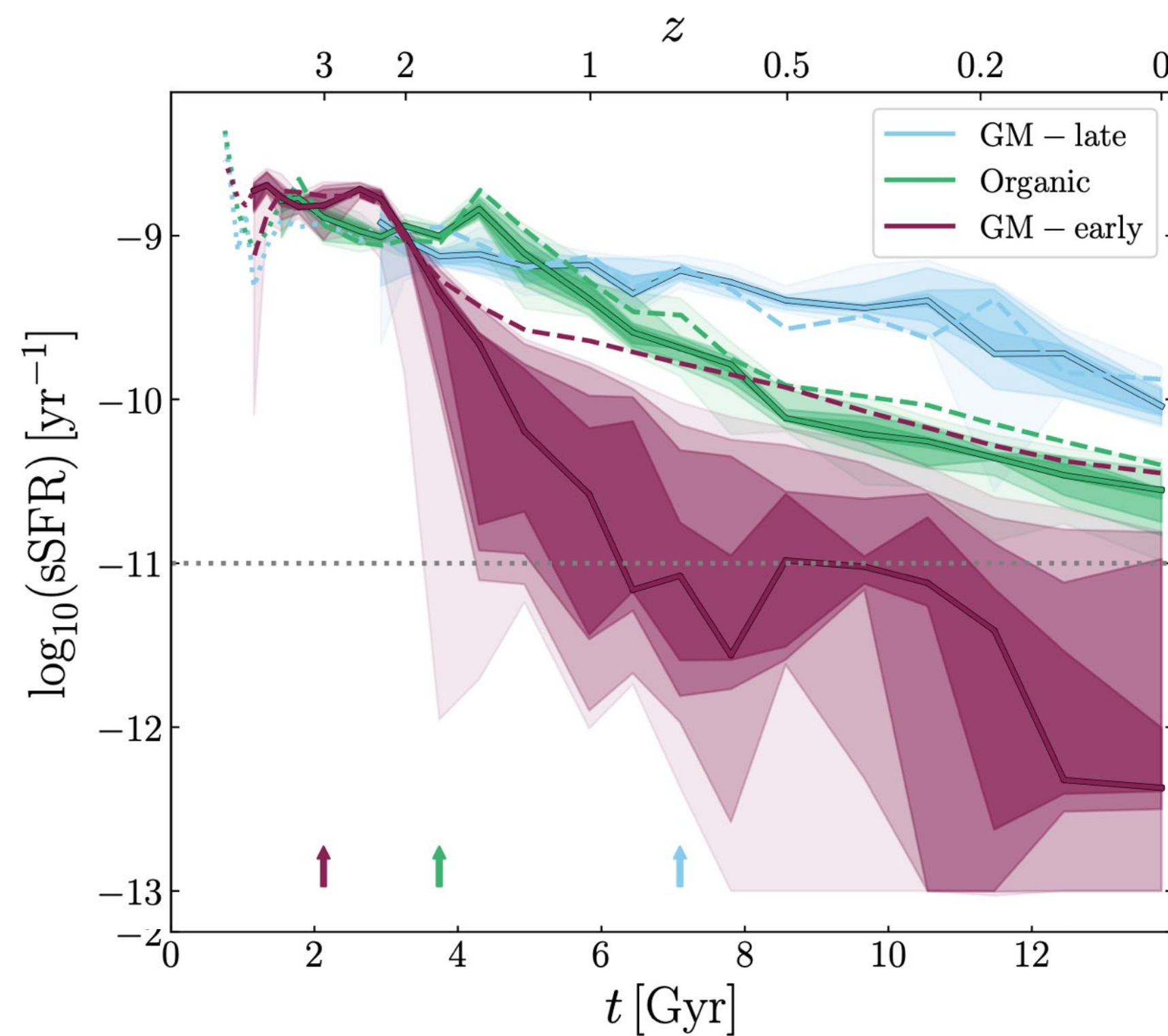
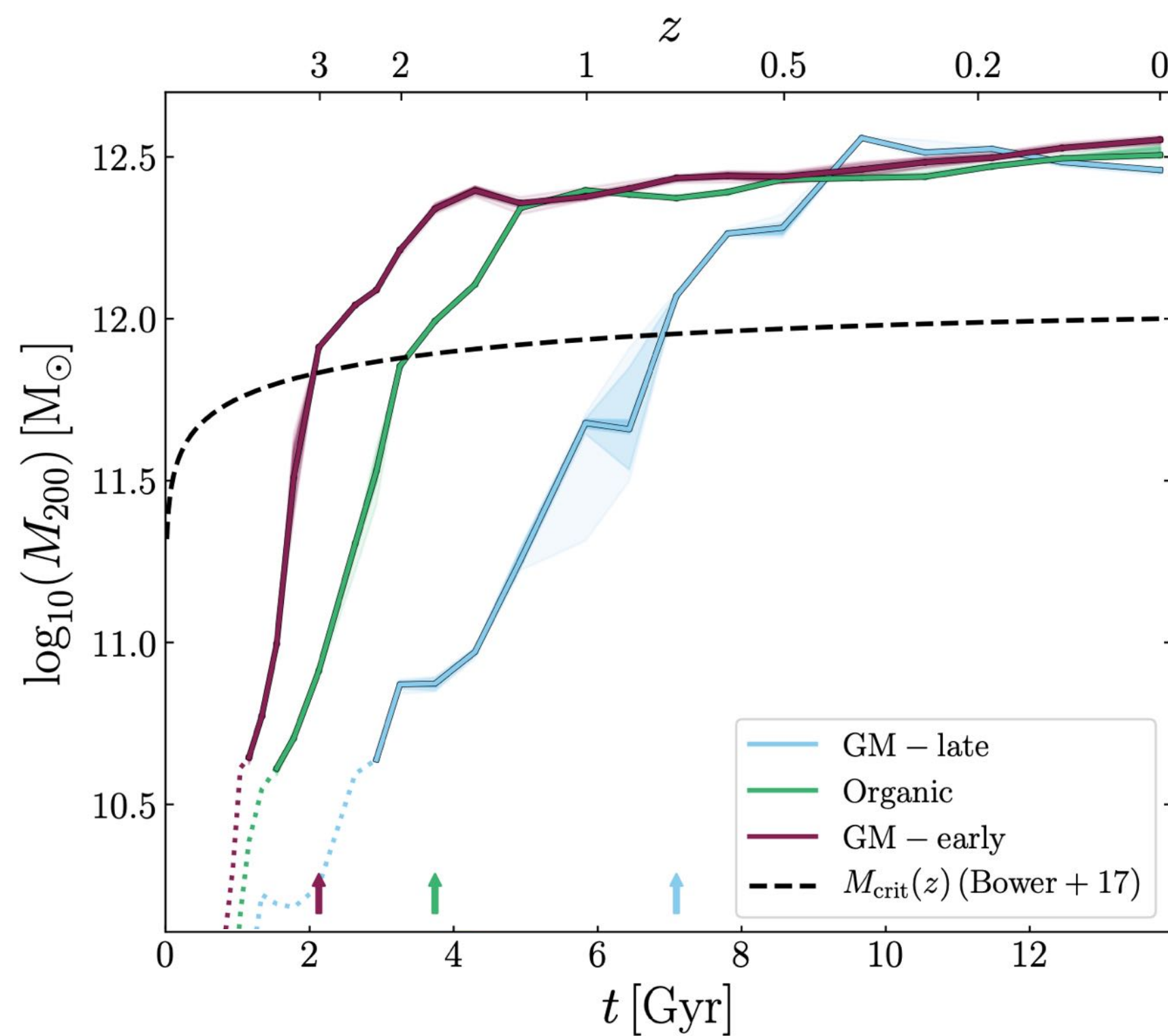
Reference



Enhanced



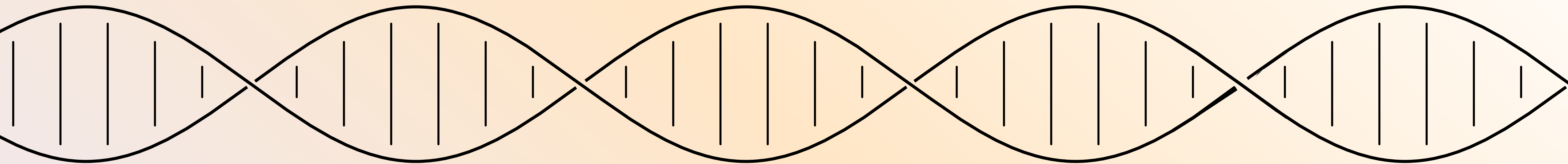
Pontzen+ 1607.02507



History + Feedback + Stochasticity
= Galaxy Formation



Davies+ 2020



- Dwarf galaxies: an exciting laboratory where we are trying to make predictions for forthcoming surveys (e.g. LSST)
- Crucial that we take advantage of this moment to make some predictions (e.g. low surface brightness dwarfs)
- “Genetic modification” – alternative versions of galaxies; controlled, cosmological numerical experiments

Andrew Pontzen, UCL
gmgalaxies.org

